# Geochemistry and magnetostratigraphy of Deccan flows at Anjar, Kutch

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Chemical analysis of nine Deccan flow basalts at Anjar, Kutch, western India, indicates that all, except the uppermost flow F-9, are alkaline. In their major and trace element composition, the alkali basalts resemble Ocean island basalts (OIB). Similarities of many diagnostic trace element ratios (e.g. Sm/Nd, Ba/Nb,Y/Nb and Zr/Nb) are similar to those found in the Réunion Island basalts. The uppermost basalt is tholeiitic and chemically resembles the least contaminated Deccan basalt (Ambenali type). The Anjar basalts have iridium concentration ranging between 2 and 178 pg/g. Some of these values are higher by about an order of magnitude compared to the Ir concentration in other basalts of the Deccan. A synthesis of chemical, palaeomagnetic and geochronologic data enables us to construct a chemical and magnetic stratigraphy for these flows.

The three flows below the iridium enriched intertrappean bed (IT III) show normal magnetic polarity whereas all except one of the upper basalts show reversed magnetic polarity. The sequence seems to have started in polarity zones 31N and probably continued up to 28R or 27R. The results presented here support the view that Deccan volcanism in Kutch occurred on a time span of a few million years.

## 1. Introduction

Deccan volcanism in India represents one of the largest areas  $(>10^6 \text{ km}^2)$  of continental flood basalts. The initial area covered by Deccan volcanics is estimated to be about  $2 \times 10^6 \text{ km}^2$ , about half of which has been eroded during the last 60 million years. The initial volume of the basalt, considering its unexposed part and erosional losses, is estimated at over  $\sim 2 \times 10^6 \text{ km}^3$  (Sen 1995) although some estimates are extremely high at about  $8 \times 10^6$  km<sup>3</sup> (Coffin and Eldholm 1993). The flows of western Ghats (Ambenali, Mahabaleshwar, Igatpuri, Poladpur etc.) have been mapped in detail and characterized using chemical, mineralogical, palaeomagnetic and geochronological studies (Cox and Hawkesworth 1985; Duncan and Pyle 1988; Subbarao et al 1988; Vandamme et al 1991, Venkatesan et al 1993, Baksi 1994, Venkatesan and Pande 1996). Though some work has also been done on basalts from the northeastern Trap province (Nagpur, Jabalpur, Salbardi etc; Yedekar

et al 1996; Deshmukh and Nair 1996; Peng et al 1998), the basalts from Kutch and Saurashtra regions, except for a few studies (Krishnamurthy and Cox 1977, 1980; Krishnamurthy et al 1988; Melluso et al 1995; Peng and Mahoney 1995), have not been investigated in detail. We have carried out chemical and palaeomagnetic studies of the Trap sequence at Anjar in Kutch (figure 1). Geochronological dating of this sequence has been done by Venkatesan et al (1996); Venketasan and Pande (private communication) and Hofmann (1997). The Anjar volcano-sedimentary sequence has attained importance in the context of identification of iridium-rich layers in one of the intertrappean beds (IT-III) and its possible association with the Cretaceous-Tertiary (K/T) boundary event since this is the only K/T site in Deccan identified so far (Bhandari et al 1995, 1996). All these results are synthesized and discussed here.

The Deccan volcanics are predominantly composed of tholeiitic basalts and alkaline, acidic and carbonatitic lavas form only minor components. It

 $\label{eq:Keywords.} \mbox{Anjar flows; K/T boundary; alkali basalts; magnetostratigraphy; geochemistry; ocean island basalts; Réunion plume.}$ 

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Figure 1. Map showing occurrence of different basalt flows in Anjar area (after Ghevariya 1988). Sample locations (AJ series) are marked. The Ar-Ar ages (in million years) measured by the PRL group and the French group (table 1) are given. The inferred polarity of various flows (N or R) is also marked. Intertrappeans (IT) are shown by hatched regions. The composite section (after Bhandari *et al* 1995) is shown as inset.

has been proposed that Deccan traps are related to the Réunion Plume (Morgan 1972, 1981; Duncan 1981; Campbell and Griffiths 1990; Basu *et al* 1993). The plume head formed the flood basalts and the trail of the plume has generated the track that ends in the active volcanoes on the Réunion Island. The plume origin is supported by the high  ${}^{3}\text{He}/{}^{4}\text{He}$  ratio, which is characteristic of the deep mantle (Basu *et al* 1993) besides other geochemical signatures (Peng *et al* 1994; Peng and Mahoney 1995).

The duration of Deccan volcanism has been a matter of intense debate. Based on magnetoand chrono-stratigraphy, Duncan and Pyle (1988), Courtillot et al (1988), Vandamme et al (1991), Allègre *et al* (1999) and Hofmann *et al* (2000)have proposed a short ( $\leq 0.5$  Ma) duration for the peak volcanic activity of the Deccan. This conclusion relies on the fact that most of the Deccan flows show a reverse magnetic polarity and only in a few sections, a maximum of two reversals (N-R-N sequence) has been seen. There is some controversy about true radiometric ages and duration of Deccan flows. The frequency distribution of <sup>40</sup>Ar-<sup>39</sup>Ar ages compiled by Courtillot et al (1986, 1988, 1996) and Duncan and Pyle (1988) show a peak at 65Ma. Venkatesan and Pande (1996), considering only the reliable ages of the basal part of the Western Ghats composite section comprising Jawhar, Igatpuri, Neral and Thakurwadi formations, generated in various laboratories based on <sup>40</sup>Ar-<sup>39</sup>Ar technique, have suggested that this basal section predates the Cretaceous/Tertiary transition by about 1.3 Ma. They further argued that this basal part of the section belongs to the 30R reversed polarity magnetic chron rather than 29R as suggested by Courtillot et al (1986, 1988). As this basal part of the Western Ghats comprises most of the flow volume, it is likely that the main pulse of Deccan came before the K/T transition. The considerably younger ages of the upper part of the section imply that the Deccan volcanism extended over 3 to 4 million years.

Raju *et al* (1991, 1995) have used the fossil evidence of intertrappean sediment beds in the Krishna-Godavari basin and concluded that the Deccan flows were emplaced during an extended time period of 5–7 Ma. Their conclusions generally support the longer duration proposed by Venkatesan *et al* (1993) and are not consistent with rapid volcanism ( $\leq 0.5$ Ma). The dates of Deccan flows from the eastern margin of Deccan (Takli, Salbardi etc) also show a span of 3–4 Ma, the ages lying between 63 and 67 Ma ( Shukla *et al* 1988; Kaneoka *et al* 1996). However, little <sup>40</sup>Ar-<sup>39</sup>Ar work has been done on Kutch basalts. A few dates by Pande *et*   $al\;(1988)$  show ages ranging between 64.4 and 67.7 Ma.

The iridium (osmium and other platinum group elements) contribution of the Deccan is another question that is quite important in connection with the K/T event and mass extinction. The global inventory of iridium at KTB is estimated to be around 250,000 tons taking its average integrated flux to be 50 ng/cm<sup>2</sup>. Orth *et al* (1990) find that iridium concentration of Deccan basalts is low, around 3 pg/g. Their measurements are confined to tholeiitic basalts. Not only is the iridium content of basalts low, but even the intertrappean sediments (marl, clay and shale etc) at a number of sites show a low (usually <30 pg/g) iridium concentration (Rocchia et al 1988) although values as high as 120 pg/g have been found in Takli intertrappeans at Nagpur (Bhandari *et al* 1993). The intertrappeans of Anjar, however, show high concentration of iridium up to 1333 pg/g (Bhandari et al 1995, 1996). They found three thin primary sub-centimeter thick ferruginous layers separated by 25–32 cm, rich in iridium, in the third intertrappean bed at Anjar. Not only are the Ir and Os high in these layers, but the intertrappean bed as a whole generally has a higher concentration of Ir, around 100 pg/g. Various arguments based on geochronology, anomalous PGE (platinum group elements) concentrations and fossil evidence suggest that the third intertrappean bed at Anjar encompasses the Cretaceous-Tertiary boundary (Bhandari et al 1995, 1996; Shukla et al 1997; Shukla and Bhandari 1997).

The Anjar volcano-sedimentary sequence has been mapped by Ghevariya and Srikarni (1989). Based on their work, it appears that there are at least seven, and possibly as many as eleven basalt flows, five of which have well developed intertrappean beds, several meters in thickness. The sequence was deposited in a lacustrine environment under semi-arid conditions (Khadkikar et al 1999). Volcanogenic clay minerals in this section have been reported by Shrivastava et al (2000) which can be expected since the whole sedimentary sequence was deposited in a volcanic environment. Some of the basalt flows are thin and limited in extent whereas others are compound flows. In the absence of a vertical section containing the entire sequence, it is difficult to identify or distinguish various basalts based on field observations. Bhandari et al (1995) have constructed a composite section based on the field observations of Ghevariya (1988) considering the seven well recognized flows. To characterize different flows, we have now carried out their detailed chemical analysis and discuss here some questions related to their stratigraphic relationship and origin.

Table 1. Flow assignments of various samples and their radiometric ages.

Flow $\#$	Sample code	Age $(Ma)^2$	Remark	Assigned age
F9	KP-7 KP-7' AJ6-3 AJ 6-8 AJ6-9	$61 \pm 1.6^{5} (WR)  60 + 2.0  60^{4}  (PL & WR)$	Integrated Age Estimated Age	$61 \pm 1.6$
F8	F-6 H.S. KP-6 KP- 6'	$\begin{array}{l} 65.0 \pm 1.2^5 \ (\mathrm{WR}) \\ 64.5 \pm 1.6^4 \ (\mathrm{WR}) \end{array}$	Integrated Age Estimated Age	$64.8 \pm 1.0$
F7	AJ9-11 AJ9-12 AJ9-2			
F6	AJ 8-12 AJ 8-7 AJ 8-4 KP-5 KP-5'	$65.0 \pm 0.4^4$ (WR) $65.9 \pm 1.6^5$ (WR)	Estimated Age Integrated Age	$65.1 \pm 0.4$
F5	AJ10-1 F-4B			
F4	F-4 KP' F-4 KP4 AJ 4-1 AJ 4-5 F-4A' F-4	$65.7 \pm 0.7^{3}$ (WR) $64.9 \pm 1.0^{4}$ (PL)	Plateau like Age Plateau Age	$65.4 \pm 0.6$
F3	F-3 B KP-3 KP3' AJ3-10 AJ3-11 AJ2-14 AJ2-3 F3#32 F3#32	$\begin{array}{c} 65.2 \pm 0.6^3 \; (\mathrm{WR}) \\ 64.9 \pm 0.8^3 \; (\mathrm{WR}) \\ \end{array}$ $\begin{array}{c} 66.9 \pm 0.2^4 \; (\mathrm{WR}) \\ 67.4 \pm 0.3^4 \; (\mathrm{WR}) \end{array}$	Plateau Age Estimated Age Total Age	$65.1 \pm 0.5$
F2	KP-2 AJ 1-10 AJ1-15	$\begin{array}{l} 66.6 \pm 0.6^5 (\mathrm{WR}) \\ 66.5 \pm 0.8^4 \ (\mathrm{PL}) \\ 65.3 \pm 0.6^4 \ (\mathrm{WR}) \\ 66.9 \pm 0.1^4 \ (\mathrm{WR}) \end{array}$	Plateau Age Plateau Age Estimated Age Total Age	$66.6 \pm 0.5$
F1	KP' KP-1 AJ11-1 AJ11-8 AJ11-2	$\begin{array}{c} 68.7 \pm 0.8^5 (\mathrm{WR}) \\ 66.8 \pm 0.4^4 \ (\mathrm{WR}) \\ 67.1 \pm 0.1^4 \ (\mathrm{WR}) \\ 67.3 \pm 0.6^4 \ (\mathrm{PL}) \end{array}$	Plateau Age Estimated Age Total Age Plateau Age	$67.8 \pm 0.5$

 $^1$  The Flow assignment is based on this work and is different in some cases from that given by Venkatesan *et al* (1996) and Hofmann (1997).

 $^{2}$  PL (plagioclase) and WR(whole rock).

<sup>3</sup> Venkatesan et al (1996).

<sup>4</sup> Hofmann (1997).

 $^{5}$  Venkatesan and Pande (personal comm.).

# 2. Sampling details

Basalt samples were collected from exposed sections at several locations around Anjar town. The various flows are exposed in a sub-parallel manner in east-west direction, younging towards south as shown in figure 1. Existence of several flows is also clear from the satellite imagery (Shukla *et al* 1997). On the eastern margin of the region, seven flows can be identified in the field whereas on the western margin, there is additionally a cluster of dykes. The dinosaur pit, where dinosaur bones were first found (Ghevariya 1988), is located south of a ring dyke near the eastern side. For the present

Flow	Source	Ir	Fe	Co	$\operatorname{Cr}$	Na	Ca	Ba	$\mathbf{Sc}$	Th	Hf	Zn	Ta	La	Ce	Nd	$\operatorname{Sm}$	Eu	Gd	Tb	Yb	Lu
F-1	AJ 11-13	139	11.0	49.9	43.5	2.12	5.22	410	22.6	6.43	6.48	147	4.56	49.3	97.0	39.4	9.01	2.76	6.20	1.26	2.91	0.50
	AJ 11-8		9.07	49.1	36.1	2.11	5.44	301	21.0	5.24	2.34		3.14	47.2	67.1	37.6	8.73	2.19	4.90	1.00	2.95	0.46
	KP-1D		11.2	50.9		2.17			24.2	6.36	6.94	138	4.52	51.0	101	49.7	9.31	2.82	8.79	1.08	2.60	0.43
F-2	AJ 1-10	178	12.5	59.8	82.3	2.56	6.53	450	26.9	7.34	7.22	178	5.23	55.9	111	51.4	10.3	3.15	6.90	1.44	3.55	0.56
	KP-2		11.9	54.8	57.4	2.30			24.0	7.03	7.52	137	4.75	52.9	103	47.2	9.46	3.04	8.70	1.04	2.28	0.42
F-3	$\#32D^1$	2.2	9.10	45.7	8.57	1.07	6.70	296	23.9	2.72	3.76	84.7	2.81	23.3	40.7	33.5	4.90	1.99	3.47	0.78	1.71	0.35
	AJ2-14		11.2	51.1		1.95	6.39	309	27.7	3.71	4.43	111	2.88	30.7	61.3	32.9	6.83	2.28	5.13	1.04	2.72	0.38
	AJ3-10		10.9	50.8				319	28.1	3.71	4.57	117		31.1	62.6	29.2	6.86	2.32	5.28	1.08	2.73	0.43
	KP-3D		10.7	51.6		2.06			28.9	3.27	4.66	104	2.69	31.5	63.6	31.4	6.95	2.29	7.19	0.82	2.19	0.37
	3B		10.9	47.8		2.05			28.8	3.41	4.79	105	2.81	31.1	65.0	32.3	6.84	2.35	6.93	0.90	1.03	0.34
F-4	AJ 4-5	50	10.2	43.3	113	2.01	5.78	272	31.1	5.66	5.29	115	1.63	29.5	60.2	29.3	7.45	2.25	5.92	1.22	3.84	0.56
	KP-4D		10.1	43.0					31.9	5.42	5.57	95	1.52		61.6	34.9	5.36	2.23	7.69	0.64	3.56	0.53
F-5	AJ10-13		10.8	57.7	45.1	2.01	4.68	386	27.1	6.33	6.55	102	4.13	47.4	98.1	40.1	9.11	2.63	4.48	1.31	3.68	0.57
	4AD		9.15	54.1					30.3	8.29	6.12		2.35		73.4	49.8	12.1	2.47	6.70	1.31	3.26	0.59
	4B		10.8	51.6		1.99			27.3	6.43	6.65	110	3.85	46.6	93.5	43.4	8.41	2.68	8.76	1.10	2.83	0.48
F-6	AJ8-4	151	8.81	32.5	22.0	2.28	4.59	391	31.4	10.1	6.32	99.1	2.86	42.0	82.4	37.2	8.62	2.51	7.02	1.45	4.87	0.75
	AJ8-7		8.42	31.2	20.7	2.24			31.4	9.67	6.75	83.0	2.66	42.1	85.7	40.8	7.84	2.49	8.66	1.23	3.97	0.70
	KP-5D		8.74	31.5					30.6	9.78	6.70	78.7	2.80		86.0	42.9	6.09	2.48	8.81	1.13	3.97	0.71
F-7	AJ9-2		11.1	49.1	93.7	2.02	6.32	476	31.2	6.48	6.72	140	4.72	46.3	90.5	73.3	9.13	2.81	7.0	1.30	3.73	0.56
	AJ9-11		10.8	48.3	103	1.97			31.0	5.97	6.38	94.8	3.88	45.0	92.3	43.5	8.03	2.71	8.47	1.16	2.37	0.49
	5B		11.0	49.5	112	1.97	11.0		31.6	6.05	6.52	122	4.06	45.3	89.8	42.2	8.58	2.71	8.56	1.03	3.06	0.44
F-8	H.S.	108	7.75	31.3	27.2	2.29	5.61	692	30.0	7.68	6.99	92.0	2.57	42.8	67.2	52.0	8.94	2.37	7.60	1.31	4.90	0.68
	KP-6D	124	8.80	31.9	27.0				31.8	9.83	6.95	94.5	2.82		86.8	44.0	6.31	2.55	8.91	1.12	3.93	0.73
F-9	AJ6-3	33	9.27	52.0	128	1.47			41.6	0.80	2.10	58.0	0.41	7.45	16.8	10.6	3.16	1.21	3.80	0.67	1.65	0.39
	AJ6-8		9.87	53.1	120	1.49	8.93		41.6	1.32	2.29	65.0	0.47	7.69	14.8	13.8	3.53	1.27	4.20	0.74	2.36	0.54
	KP7	26	9.64	53.0	125				41.5	1.15	2.09	75.0	0.42		15.3	10.7	3.40	1.21	3.58	0.61	2.22	0.37
Error(%)		5	2.0	2.0	2.0	1.0	2.0	10	2.0	1.0	1.0	5	3.0	2.0	2.0	7.0	2.0	2.0	5.0	5.0	7.0	3.0

Table 2. Concentration of various elements in Anjar basalts based on INAA. Na, Ca and Fe in wt%, Ir in pg/g and remaining elements in  $\mu g/g$ .

<sup>1</sup> Bhandari *et al* (1996)

	F-1	F-1	F-2	F-3	F-3	F-3	F-3	F-4	F-4	F-5	F-5	F-5	F-6	F-6	F-7	F-8	F-9	F-9	MB
	KP-1	AJ11-2	AJ1-15	#32	AJ2-3	AJ3-11	KP-3	AJ4-1	KP-4	AJ10-16	4A	4B	AJ812	KP-5	AJ9-12	KP-6	AJ6-9	KP-7	81#1
$\mathrm{SiO}_2(\%)$	45.34	45.16	45.11	45.73	46.00	45.62	45.15	49.78	49.35	46.14	45.69	45.48	52.37	50.32	45.15	51.84	45.78	45.99	48.53
$\mathrm{TiO}_2$	3.77	3.77	3.95	3.58	3.77	3.74	3.80	2.85	2.78	3.94	3.82	3.94	3.11	3.30	3.65	3.15	1.56	1.60	1.03
Al <sub>2</sub> O <sub>3</sub>	15.61	15.54	15.76	15.71	15.30	15.41	14.64	14.86	15.40	15.05	15.87	15.04	15.48	15.76	16.03	15.29	17.03	16.87	15.61
$Fe_2O_3$	15.72	16.03	15.32	15.61	15.26	15.82	16.11	14.68	14.49	16.35	15.79	16.71	12.46	13.54	15.74	12.18	14.40	14.09	12.16
MnO	0.19	0.20	0.22	0.21	0.24	0.22	0.22	0.21	0.19	0.23	0.20	0.23	0.20	0.19	0.20	0.18	0.21	0.20	0.20
MgO	5.48	4.96	5.13	5.28	5.50	5.15	5.72	4.27	4.35	4.74	4.95	4.62	3.56	3.63	5.22	3.67	5.99	5.95	10.78
CaO	8.10	8.32	8.57	9.40	9.03	9.24	10.12	8.49	8.49	9.05	8.13	9.07	7.71	7.67	8.78	7.76	12.05	11.93	8.40
Na <sub>2</sub> O	2.90	2.92	3.08	2.50	2.73	2.81	2.54	2.94	3.02	2.61	2.92	2.64	3.22	3.47	2.84	3.14	2.19	2.22	1.99
K <sub>2</sub> O	1.67	1.68	1.68	1.13	1.12	1.17	0.94	1.03	1.07	1.10	1.60	1.07	1.41	1.42	1.49	1.66	0.17	0.15	0.69
$P_2O_5$	0.50	0.54	0.51	0.38	0.42	0.42	0.35	0.35	0.35	0.39	0.55	0.39	0.44	0.39	0.52	0.47	0.16	0.15	0.12
$\overline{\mathrm{Ba}(\mu\mathrm{g}/\mathrm{g})}$	347.5	298.5	332.6	250.5	277.6	237.5	263.0	237.7	250.3	275.9	311.2	331.3	308.3	324.7	269.6	546.3	102.7	116.3	128.7
Rb	40.9	41.25	41.89	32.27	18.62	24.58	14.60	39.77	36.90	45.86	47.36	43.78	70.97	78.43	43.71	66.63	2.64	5.98	18.3
Sr	453.7	414.4	403.6	885.6	436.9	402.8	450.3	240.1	239.4	372.7	352.2	354.8	222.2	214.6	343.4	1838.7	165.6	165.1	165.0
Y	34.25	33.89	34.71	27.03	28.54	27.58	28.78	38.22	36.94	40.06	40.97	37.86	47.46	48.38	39.65	48.63	24.91	23.37	18.76
Zr	289.1	290.6	292.3	180.1	189.9	179.5	184.9	215.9	209.5	288.7	297.0	272.2	267.3	269.2	275.5	271.1	77.50	76.22	80.91
Nb	57.27	63.02	61.12	36.22	36.25	37.32	35.51	21.39	18.80	60.81	64.12	53.62	35.79	36.88	40.91	39.64	7.03	7.20	4.86
U	1.02	0.92	1.04	0.51	0.68	0.54	0.63	0.89	0.99	0.95	0.88	1.04	1.48	1.63	0.86	1.58	0.22	0.29	0.40
Th	3.33	2.92	3.20	1.58	1.87	1.88	1.89	2.64	2.89	3.10	3.24	3.38	4.86	5.08	2.97	5.00	0.65	0.64	1.32
Pb	4.31	7.09	9.59	3.68	5.28	16.92	2.87	9.66	5.95	8.16	2.19	3.00	6.88	7.24	5.42	3.21	5.64	6.89	5.18
Ga	25.34	27.54	26.04	22.79	24.76	22.50	22.57	26.52	24.50	25.71	27.19	25.10	23.93	25.04	28.19	24.93	20.93	20.28	18.77
Zn	146.6	170.0	163.8	136.6	140.4	116.7	122.9	143.3	131.9	171.9	169.5	145.3	144.7	131.3	163.6	138.3	112.7	215.5	90.25

Table 3. Composition of Anjar flows based on XRF and ICP-MS analyses.

Cu	274.1	308.2	305.3	178.4	170.9	157.7	164.8	226.5	216.2	284.2	312.1	252.7	274.5	268.7	309.6	281.8	187.6	203.6	123.3
Ni	57.08	83.65	87.67	37.77	32.06	25.15	28.34	56.89	51.52	68.54	62.02	63.61	47.95	48.76	56.32	52.36	73.09	76.70	143.0
V	402.3	406.7	408.5	480.6	442.9	361.7	456.4	508.7	489.6	440.1	429.6	437.6	536.3	547.5	554.0	486.4	426.8	358.2	387.6
$\overline{\mathrm{Cr}}$	33.38	45.07	61.16	25.57	20.91	11.67	20.21	68.69	65.36	38.71	37.36	45.17	31.26	25.99	50.89	28.83	52.92	52.99	143.4
Hf	7.16	5.75	6.06	3.53	4.40	4.57	4.39	4.82	5.51	5.92	5.69	6.52	5.68	6.32	5.37	6.03	1.88	2.01	1.63
Cs	0.11	0.09	0.25	0.05	0.19	0.24	0.29	0.52	0.52	0.46	0.25	0.46	1.03	0.97	0.16	0.77	0.07	0.08	0.16
$\mathbf{Sc}$	24.57	29.95	27.35	32.91	30.76	29.23	29.45	37.32	34.99	35.20	36.42	28.81	37.48	34.99	37.21	35.75	49.03	43.25	38.39
Co	54.96	62.80	54.80	53.41	51.85	48.58	50.28	48.68	46.04	60.86	55.60	52.32	39.74	37.18	50.89	36.17	56.18	55.44	47.36
La	50.54	45.26	47.71	26.78	30.80	30.58	29.94	28.93	29.13	43.43	46.31	47.23	44.36	42.64	41.43	41.94	7.42	7.34	11.11
Ce	103.7	91.89	96.41	56.58	64.83	65.08	63.79	59.03	62.36	87.73	94.34	95.80	80.13	86.37	85.18	84.82	16.59	16.39	21.88
Pr	10.89	9.60	10.18	6.04	6.90	7.13	6.85	6.43	6.85	9.22	9.73	10.15	8.48	9.10	8.55	8.81	2.06	2.05	2.34
Nd	43.60	38.89	40.30	25.35	29.08	28.65	27.71	27.54	28.18	37.98	39.39	39.73	34.24	37.14	34.64	36.50	9.79	10.06	9.76
Sm	9.39	8.09	9.10	6.01	6.43	6.99	6.81	6.76	6.60	7.94	8.25	8.17	7.61	8.48	8.34	7.92	3.25	3.04	2.37
Eu	2.97	2.56	2.54	1.97	2.11	2.24	2.33	2.09	2.23	2.55	2.56	2.69	2.40	2.72	2.30	2.52	0.97	1.10	0.87
Gd	10.13	9.09	9.40	6.80	7.12	7.11	7.61	7.94	8.22	8.82	9.95	9.24	9.15	10.11	9.12	9.81	3.96	4.07	3.02
Tb	1.28	1.15	1.21	0.85	1.03	1.00	0.97	1.11	1.20	1.20	1.22	1.30	1.32	1.37	1.21	1.40	0.66	0.71	0.49
Dy	6.24	5.60	5.61	4.08	4.75	4.99	4.72	5.53	6.05	5.78	6.03	6.61	7.14	7.36	6.02	7.18	3.57	3.50	2.39
Но	1.08	0.90	1.10	0.71	0.88	0.90	0.87	1.14	1.10	1.14	1.05	1.22	1.32	1.49	1.09	1.42	0.72	0.71	0.47
Er	3.85	3.28	3.27	2.67	3.03	2.95	3.02	3.54	4.06	3.94	3.82	4.07	4.61	4.95	3.60	4.85	2.46	2.48	1.75
Tm	0.58	0.54	0.49	0.39	0.48	0.49	0.51	0.63	0.62	0.52	0.64	0.67	0.86	0.92	0.57	0.83	0.43	0.40	0.30
Yb	2.93	2.49	2.43	1.86	2.28	2.33	2.55	3.00	3.42	2.80	3.06	3.41	4.01	4.17	2.74	4.34	1.94	2.13	1.44
Lu	0.42	0.33	0.35	0.27	0.32	0.32	0.30	0.41	0.50	0.40	0.40	0.46	0.56	0.59	0.42	0.58	0.28	0.29	0.21
$\overline{mg^*}$	0.51	0.50	0.51	0.50	0.52	0.50	0.51	0.45	0.46	0.46	0.48	0.45	0.45	0.45	0.49	0.47	0.51	0.52	0.58

study, several hand specimens were collected during 1991–93 for chemical analysis and in June 1994 for magnetic polarity studies and <sup>40</sup>Ar-<sup>39</sup>Ar dating. Subsequently, 11 oriented drill cores (AJ series) for palaeomagnetic and geochronological studies were collected in April 1995 jointly with the French team of Gallet, Rocchia, Feraud and Hofmann. Several samples (KP series), appearing relatively less weathered were also collected from nearby locations for <sup>40</sup>Ar-<sup>39</sup>Ar dating by K. Pande. Two northsouth traverses were made and wherever possible the samples were collected from the west and east margins of each flow as shown in figure 1. A number of samples from various intertrappeans (IT 1, 2, 3, 5 and 6) were also collected for chemical and mineralogical analysis. The results of analysis of intertrappeans will be discussed elsewhere and here we confine ourselves to the basalt samples from various flows. The sample locations are shown in figure 1. The flow assignment for various samples and their <sup>40</sup>Ar-<sup>39</sup>Ar ages measured by Venkatesan *et al* (1996), Venkatesan and Pande (personal communication) and Hofmann (1997) are given in table 1. Chemical and field studies carried out here indicate that there are nine flows, one or two of which could possibly be dykes. The flow assignment (except for the lower three flows) in this work is therefore not the same as that of Hofmann (1997) and Bhandari et al (1995, 1996).

## 3. Experimental techniques

Thirty six samples belonging to the nine flows collected from different locations were analyzed for their chemical composition using neutron activation analysis (INAA), inductively coupled plasma mass spectrometer (ICPMS) or X-ray fluorescence (XRF) methods. Standard procedures were used for neutron activation analyses. Powder samples (typically 200mg) were irradiated in the CIRUS reactor of the Bhabha Atomic Research Centre, Bombay with a neutron fluence of  $\sim 10^{18}$  neutrons, together with BCR, Allende meteorite and other standards (AGV, SDO) kindly provided by the U.S. Geological Survey. The concentrations of various elements, listed in table 2, were estimated by periodical gamma ray spectrometry of the irradiated samples using a hyper pure coaxial germanium diode (relative efficiency 48%) and a low energy planar detector. Sixteen samples were analyzed in this way by INAA and Ir was also measured in all the flows following radiochemical separation (Bhandari et al 1994). Measurements of major and trace elements (table 3) were performed using XRF and ICPMS on twenty of these samples at the National Geophysical Research

Institute, Hyderabad. A sample of Mahabaleshwar Flow (MB81#1 of J.J. Mahoney collection) was also included in these analyses for the sake of comparison. XRF measurements were carried out using Philips PW 1400 model microprocessorcontrolled sequential X-ray fluorescence spectrometer following Govil (1985). Measurement of major and minor elements were found to be reproducible within  $\pm 1\%$ . The trace elements and REE were determined using ICP-MS (Plasma Quad, PQ1). Japanese basalt standard (JB) was used as a reference. Details of the various procedures adopted are discussed in Balaram et al (1995, 1996). The trace element and REE data were found to be reproducible with precision better than  $\pm 7\%$ in the standards. U, Th and K were measured by gamma ray spectrometry using a hyperpure germanium detector having a relative efficiency of 25%, located in a low background 10cm thick lead shield. The concentration of U and Th, given in table 4, are based on the characteristic gamma ray energies of their daughter radionuclides (<sup>214</sup>Bi and <sup>208</sup>Tl respectively), assuming a secular equilibrium. Basalt 107 (U = 5.69 ppm, Th = 14.5 ppm and K = 2.63%) in which the natural series is known to be in equilibrium was used as a standard.

Measurements of magnetic polarity were made on oriented hand specimens collected during June 1994 and drill core samples collected during 1995 using a Schonstedt spinner magnetometer (U.S.A.). Measurements were made on 5 to 15 samples of each oriented core (or specimen) and the average values are given in table 5. After the measurement of Natural Remanent Magnetisation (NRM), the samples were subjected to thermal and/or alternate field demagnetisation (AFD) and the magnetic field was measured at several stages of demagnetisation.

## 4. Results and discussions

The results of analyses based on INAA are given in table 2 and those determined with XRF and ICPMS are given in table 3. The precision of measurements for each element is given in these tables. In general, there is an agreement between the INAA, XRF and ICP-MS results although some values lie outside the expected range of errors. A similar situation exists for U, Th and K concentrations measured by gamma ray spectrometry and other techniques. The dispersion may be due to several reasons such as weathering, alteration, sample inhomogeneity, use of different standards in different analyses as well as due to genuine variation in samples taken from different locations of the flows. Sample #32 (Flow 3), published earlier (Bhandari et al 1996) appeared heavily weathered

Table 4. Concentration of U, Th and K determined by  $\gamma$  -ray spectrometry<sup>1</sup>.

Flow $\#$	Sample	U (ppm)	Th ( ppm )	K ( wt %)	$\mathrm{Th}/\mathrm{U}$	K/U
F-9	AJ 6-8	0.29	1.54	0.09	5.3	3103
F-7	AJ 9-2	1.61	7.29	1.29	4.5	8012
F-6	AJ 8-4	2.56	10.9	1.20	4.2	4687
F-5	AJ 10-13	1.66	7.65	1.02	4.8	7692
F-5	4A	1.69	8.23	1.30	4.6	6144
F-4	AJ 4-5	1.57	6.24	0.83	3.9	5286
F-3	AJ 3-10	0.87	3.84	0.86	4.4	9885
F-3	AJ 2-14	0.92	4.29	0.97	4.6	10543
F-2	AJ 1-10	1.84	7.74	1.42	4.2	7717
F-1	AJ 11-13	1.16	4.32	1.43	3.7	12327

<sup>1</sup> U and Th concentrations were estimated via their daughter radionuclides assuming secular equilibrium of the radioactive series.

when examined petrographically and its low concentration of Ba, Na, Cr, Ni etc may be due to this reason. Weathering can also lower concentration of elements like iron and sodium. Lead concentrations in different samples were also found to lie in a large range. In neutron activation analysis, the statistical errors due to counting are low as indicated in table 2 but values of certain elements like Tb, Gd, and Zn have interferences from high concentrations of elements giving strong gamma peaks in the neighbourhood of their gamma ray lines. K and Na in some samples could not be determined to a good accuracy because of significant decay before they could be counted. In spite of these uncertainties, various flows can be distinguished based on characteristic trace element concentrations and ratios as discussed later and the precision or reproducibility does not affect the arguments developed below. In this discussion we either use mean values of different analyses or those considered to be the most reliable, based on considerations of errors of measurements and interferences as discussed above.

### 5. Chemical stratigraphy

The Anjar flows have high amount of Fe<sub>2</sub>O<sub>3</sub> (12.18 – 16.7%) and TiO<sub>2</sub> (1.56 – 4.0%) which is similar to that generally found in some other Deccan basalts. Except for the upper-most flow (F9) in which TiO<sub>2</sub> is ~1.58 %, the concentration in all other flows is  $\geq 3$  %. These flows, except for F-3 and F-4, have high Al<sub>2</sub>O<sub>3</sub> (14.64–17.03 %). Therefore, they can be classified as high alumina basalts. The most striking feature is that the SiO<sub>2</sub> content, with the exceptions of F-4, F-6 and F-8, range between 45.11 and 46.14%, similar to that

in alkali basalts. In comparison, F-4, F-6 and F-8 have relatively high SiO<sub>2</sub> content (49.35–52.37 %). Similarly the  $K_2O$ , Rb and Ba are also enriched in Anjar flows as compared to the flows of Mahabaleshwar and Narmada region (Mahonev 1988). The mg<sup>\*</sup> number  $(Mg^{2+}/(Mg^{2+} + Fe^{2+}))$ , is low (0.45 to 0.52) as shown in table 3, in comparison to the other Deccan flows (e.g. MB81#1 = 0.58). The values indicate that the Anjar basalts are relatively more evolved compared to the primitive members of the Deccan, which have  $mg^* = 0.7$ . Flows 1, 2 and 3 cluster in one group  $(mg^* = 0.5)$ , and 4, 5, 6 fall in another group ( $mg^* = 0.46$ ). The remaining flows 7, 8 and 9, have distinct values and do not fall in a group, thus their mg<sup>\*</sup> suggests change in magma composition. The REE patterns for various flows are shown in figure 2. Flow 9 is distinctly different from other flows and is similar to the least contaminated Ambenali type of basalts (Cox and Hawkesworth 1985; Mahoney 1988; Sano and Fujii 1996). The REE pattern of F9 is flatter and more depleted in LREE compared to Ambenali, indicating that it may represent even less contaminated magma. All other basalts have a high enrichment of LREE.

The Deccan basalts of the Western Ghats have been classified into several formations based on variation of some elements and elemental ratios e.g. Sr, Ba, Rb, Zr/Nb, Ba/Y, and TiO<sub>2</sub> (e.g. Devey and Lightfoot 1986; Beane *et al* 1986; Lightfoot *et al* 1990). These criteria are reproduced in table 6(a) and we can use them to classify Anjar flows. It may be noted that concentrations of some of these elements like Rb and Ba, however, are known to change as a result of weathering and alteration, whereas others like Zr and Nb are not much affected and therefore are considered more diagnos-

Table 5. Measurements of D and I of NRM and after thermal or AF demagnetisation and their comparison with measurements made at IPGP (Hofmann 1997).

Flow	Sample codes	Pl	$RL^1$	IP	GP	Inferred	Adopted Age $(Ma)^2$
		D	Ι	D	Ι	Polarity	
F-9	$\mathrm{F7}$	150.50	62.0			Reverse	61.6
		$150.\ 46$	$46.38^{*}$				
	AJ6	127.74	66.12	149	41.8		
F-8	F-6H.S.	3.0	18.0			Reverse	64.8
		131.91	$58.63^{*}$				
F-7	AJ9	356.45	18.04			Secondary vis- cous component	
F-6	AJ 8	116.8	74.25	159.2	50.9	Reverse	65.1
F-6	F 5B	144.25	39.44			Reverse	
		146.6	$17.4^{*}$				
F-5	F-4A'	33.59	36.30			Normal	
		349.0	$-38.55^{*}$				
	F4	360.02	-37.71				
		348.97	$-38.55^{*}$				
F-5	AJ10	338.04	-20.4			Normal	
				343.6	-30.6		
F-4	AJ4	148.01	51.51			Reverse	65.4
		155.80	21.67	164.7	18.1		
F-3	F3#32	337.8	-31.89			Normal	65.1
		344.65	-41.52				
F-3	AJ2	332.72	-32.05	326.6	-55.8	Normal	
		326.0	-48.69	156.5	66.2		
F-3	AJ3	349.18	-4.51	353.7	-47.8	Normal	
		311.12	-47.24	142.3	32.9		
F-2	AJ 1	316.75	-12.05			Normal	66.6
		308.94	-17.76	309.7	-24.0		
	F2	317.94	32.94				
		331.52	$-65.46^{*}$				
	F1	299.87	- 42.53				
F-1		310.0	-33.35*				
F-1	AJ11	352.14 315.9	$19.65 \\ 7.57$	322.5	-2.9	Secondary vis- cous component	67.8

 $^1$  The results after Alternate Field Demagnetisation (AFD) with 500 Oe or thermal demagnetisation at 550°C (\*). **D**: Declination, **I**: inclination

 $^{2}$  Adopted based on data given in table 1 and figure 4.



The concentration of U, Th and K, particularly the ratio Th/U and K/U can also be used to distinguish different basalts. It is clear from table 4 that the sample F9 and F1 have distinctly different Th/U and K/U whereas all the others fall within a narrow field.

Iridium is usually found to be low in the Deccan tholeiites, varying between 3 and 30pg/g (Orth *et al* 1990). However, we find that the iridium values in Anjar flows range between 2 and 178 pg/g (table 2). Surprisingly the flows 3 and 4, between which the iridium rich intratrappeans are found have low values of 2 and 50pg/g respectively. The unusually high values of iridium (30-178pg/g) in Anjar basalts, because of their small volume, however, do not substantially increase the contribution of the Deccan volcanism to the K/T iridium enhancement and the high iridium inventory at KTB requires a bolide input (Alvarez *et al* 1980).

# 6. Chronostratigraphy



Figure 2. Typical Rare Earth Element (REE) patterns of each of the Anjar flows (a) normalized to chondrite and (b) Ocean Island Basalts (OIB).

tic for classification. The average concentrations of these elements in various Anjar flows are given in table 6(b). Comparison of these data with table 6(a) shows that F-9 is similar to the Ambenali type basalts whereas the other flows have no similarity to any of the Western Ghat flows. These observations suggest that the Anjar flows, except F9, are different from the typical Deccan flows. Measurement of isotopic ratio ( $^{87}$ Sr/ $^{86}$ Sr) in Anjar flows, which can provide additional confirmation of this conclusion, is in progress.

The various Anjar flows have been dated using the <sup>40</sup>Ar-<sup>39</sup>Ar method at PRL as well as at IPGP, France. Earlier, Venkatesan et al (1996) reported the plateau age of F-3 (65.2  $\pm$  0.6, 64.9  $\pm$  0.8 Ma) and plateau like age of F-4 (65.7  $\pm$  0.7Ma) based on whole rock analyses. These ages put some constraint on the time span of deposition of the intertrappean-III lying in between F-3 and F-4, in which Ir anomaly has been reported (Bhandari et al 1995 and 1996). To understand the complete picture of Anjar sequence, we have compiled the available ages of the flows in table 1, based on the work of Venkatesan et al (1996), Hofmann (1997) and Venkatesan and Pande (personal communication). Considering multiple analyses by these groups, reliability and errors of measurements, each flow can be assigned a certain mean age. These assigned ages are also given in table 1. In case of F-1, F-2, F-3 and F-4, the weighted mean of plateau ages have been taken. In case of F-6, F-8 and F-9, the available integrated or estimated ages have been subjected to the weighted mean analyses. In this way, we have obtained the time span of the Anjar flow sequence, which ranges from  $67.8 \pm 0.5$  to 61 $\pm$  1.6 Ma. The uppermost flow F-9 did not yield an age plateau and is assigned an age of  $61 \pm 1.6$  Ma, being the most precise value available from the two laboratories. The flows F-8, F-6, F-4 and F-3 are similar in age and indistinguishable from the age of the K/T transition ( $65.2 \pm 0.1$  Ma; Izett *et al* 1991). However, the plateau ages of mineral separates and whole rock of flows F-2 and F-1 are older, 66.6  $\pm$ 0.5 Ma and  $67.8 \pm 0.5$  Ma respectively. This range of ages fall close to the ages of the basal flows of

Table 6. (a) Discrimination criteria of different formations of Deccan traps adopted by Lightfoot et al (1990). The Ba/Y and  ${}^{87}Sr/{}^{86}Sr$  ratio criteria are taken from Devey and Lightfoot (1986).

Formation	${ m Sr}$ (ppm)	Ba (ppm)	Rb (ppm)	$\rm Zr/Nb$	Ba/Y	$ \begin{array}{c} {\rm TiO}_2 \\ ({\rm wt}\%) \end{array} $	$^{87}\mathrm{Sr}/^{86}\mathrm{Sr}$
Mahabaleshwar	> 225	> 90		< 13.0	4	> 2.00	> 0.705
Ambenali	< 240	< 100	< 8	< 11.0	< 3.5	< 2.70	< 0.705
Poladpur		> 100	> 8	> 11.0	> 3.5		0.705 - 0.713
Bushe	< 185	> 100	> 15	> 17.0		< 1.50	> 0.713
Panhala	< 200	< 100	< 10	> 12.5		< 2.10	
Desar	> 230	> 150	> 15	< 12.5		< 2.25	
Kolhapur	> 200			< 14.0		> 2.25	

Table 6. (b) Average concentrations of some characteristic elements and elemental ratios in Anjar flows.

Flow	Sample	$\operatorname{Sr}$	Ba	Rb	$\mathrm{Zr/Nb}$	$\mathrm{Ba/Y}$	${\rm TiO}_2$
		(ppm)	(ppm)	(ppm)			(wt%)
F-9	KP-7	165.4	109.5	4.3	10.78	3.05	1.58
F-8	KP-6	1838.9	564.25	66.63	6.84	11.60	3.15
F-7	AJ 9-12	343.35	269.56	43.71	6.74	6.71	3.65
F-6	KP-5	218.4	316.5	74.7	7.38	6.65	3.21
F-5	4 A	359.9	306.11	45.7	4.82	7.7	3.9
F-4	AJ 4-1	239.73	244.0	38.33	10.62	6.55	2.82
F-3	AJ 2-3	7430.0	259.4	19.2	5.08	9.0	3.77
F-2	AJ 1-15	403.62	332.61	41.89	4.78	9.58	3.95
F-1	AJ 11-2	434.04	323.0	41.08	4.83	948	3.77

the Western Ghats (Venkatesan and Pande 1996). This observation clearly indicates that the Deccan volcanism started before the K/T transition. We discuss this point further in the next section while considering the palaeomagnetic measurements.

## 7. Magnetic stratigraphy

The results of measurements of natural remanent magnetisation (NRM) are given in table 5. Many of the samples show a viscous component, which can be resolved by AFD (alternate field demagnetisation) or thermal treatment. The deduced declination and inclination for various flows are given in table 5. If we make a simplistic assumption that any viscous component present should acquire the present day field at the site, then such samples should not be considered for deducing the primary magnetisation acquired at the time of solidification of the flows. The inferred polarity given in table 5 excludes such samples. In some cases, it is difficult to erase the secondary component totally by thermal or AF demagnetisation, but there are samples where secondary imprints could be com-

pletely erased and direction of primary magnetisation measured. We show some typical examples of AFD behavior of Flows 1, 2, 3 and 4 in vectorial projections (figure 3). Replicate measurements of Flow 1, 2 and 3 gave consistent results showing normal polarity whereas Flow 4 shows reversed polarity. Hofmann (1997) has discussed the complexity of magnetic records in various flows and problems in deducing the direction of primary magnetization. In the flows above the iridium layer (except for F-5), the primary direction of magnetisation could be deduced with confidence. These results are also given in table 5 for comparison. First we compare our results with that of Hofmann (1997) since many of the measurements were made on the same core samples. There is a general agreement between the two data sets obtained at PRL and IPGP as far as the apparent direction of magnetization is concerned. In case of Flow 1 (AJ11) and Flow 7 (AJ9), Hofmann (1997) finds that magnetic records are too complex to determine the direction of primary polarity whereas in the case of Flows 2 and 3, she has reported a weak reversed primary field. These conclusions do not fully agree with our work and in view of the strong secondary imprint, these results require further confirmation.



Figure 3. Typical examples of Alternate Field Demagnetisation (AFD) behaviour for Flows 1, 2, 3 and 4. The vectorial diagrams (Zijerveldt plots) show projections on vertical and horizontal planes as a function of AFD field strength. The axes are in  $(x \ 10^{-3})$  gauss/g.



Figure 4. Magnetic and geochronological stratigraphy based on assigned ages, (see text) of various Anjar Flows. The magnetic chrons are taken from Cande and Kent (1995). F-5 and F-7 are not included because their ages are not known. The location of K/T boundary, in the third intertrappean bed where triple iridium anomaly has been found is marked.

In summary, considering only the reliable measurements, we find that the lower flows (1, 2, and 3) are normal and the upper flows (4, 6, 8 and 9) are reversed. It is possible that the flows 5 and 7, which were collected from the western fringe of the region, close to a cluster of dykes, have their magnetic polarity influenced by thermal effects associated with dyke formation. This uncertainty must be kept in mind in the following discussion.

Combining the magnetic data with the geochronological framework, we find that Flows 4, 6, and 8 can be placed in Reversed chron 29R. The lower flows (F2 and 3) will probably lie in chron 30N and F-1 probably in 31N (figure 4). The age of Flow-9 is not reliable but it may belong to chron 28 or 27R. For F-7 and F-5, the magnetic data and ages have not been obtained and we can not ascertain their position in the chronostratigraphic sequence. In spite of these uncertainties, the results

on magneto- and chrono-stratigraphy confirm that the Anjar volcano-sedimentary sequence was not emplaced in a short period of time (<1Ma) but had an extended period of formation as proposed by Venkatesan *et al* (1993) and Venkatesan and Pande (1996).

Traditionally, the K/T boundary is believed to fall in Reverse chron 29R although Lerbekmo and Sweet (1994) have reported a short (~ 45ka) normal polarity chron, just at the KTB within 29R. The age bracket (68–61 Ma) of Anjar flows is the same as that of Mahabaleshwar sequence in Western Ghats, which were also dated in our laboratory yielding identical time span (67 to 61 Ma; Venkatesan *et al* 1993). The magnetic polarity of all the Mahabaleshwar flows is reversed (Khadri *et al* 1988) whereas the Anjar flows show at least one Normal-Reverse sequence (table 5, figure 4). The upper Anjar flows above the iridium layer, (except F-5 and F-7, where the results are ambiguous), therefore can possibly be correlated with the Mahabaleshwar sequence from the point of view of their magnetic polarity. However, the lower flows have the same age as the basal part of the Mahabaleshwar sequence (Venkatesan *et al* 1993) but show opposite polarity. This disagreement requires further study to be resolved.

The change from normal to reverse polarity seems to have occurred within IT-III, where the KTB has been identified. However, the data for the upper flows are not sufficiently reliable for us to conclude that the impact, which presumably gave rise to the iridium rich horizons in IT-III, was also responsible for the polarity change during the K/T transition. Careful and high resolution study of intertrappean sediments, which have very mild magnetic field, is required to address this point.

## 8. Geochemical nature of the magma

Certain chemical criteria have been developed to characterize the parent magmatic fluids from which basaltic lavas originate and to distinguish between different basalt types. Some of the well established criteria are based on TiO<sub>2</sub> content,  $(Na_2O + K_2O -$ SiO<sub>2</sub>),  $MnO_2 - TiO_2 - P_2O_5$ ,  $TiO_2 - Zr/P_2O_5$ ,  $Nb/Y-Zr/P_2O_5$  and Zr/Y-Zr discrimination diagrams. In order to use these criteria, we have plotted the major and trace element data in various discriminating diagrams (figures 5, 6, 7, 8) discussed below.

The geochemical data on Anjar basalts reveal that all the flows (except F-9) have an alkaline nature. On the total alkali-SiO<sub>2</sub> diagram (figure 5a), the Flows 1, 2, 3, 5 and 7 fall above the discriminating line between subalkalic and alkali fields whereas Flows 4, 6 and 8 lie close to the demarcating line. Flow F-9 lies in the subalkalic region. This is further supported by Zr/TiO<sub>2</sub>-Nb/Y diagram (figure 5b) where all flows except F-4 and F-9 lie in the alkali basalt field. F-4 lies close to the transition zone between tholeiites and alkali basalts. That the high concentration of alkali elements is not due to crustal contamination or subsequent alteration of basalt can be confirmed from the  $TiO_2 - Zr/P_2O_5$  (x10<sup>4</sup>) and Nb/Y-Zr/P<sub>2</sub>O<sub>5</sub>(x10<sup>4</sup>) diagrams of Floyd and Winchester (1975). These plots, shown in figure 6(a) and 6(b), confirm that the magma which gave rise to the Anjar basalts, particularly F-1, 2, 3, 5 and 7 was alkaline, F-4, 6 and 8 was transitional while F-9 was subalkaline. This is also supported by high Nb concentration (and high Nb/Y) as given in table 3, a criterion proposed by Winchester and Floyd (1977). Again, since Th, Yb and Ta are highly incompatible elements, which can not be readily incorporated in



Figure 5. (a) The total alkali and silica (TAS) diagram i.e.  $Na_2O + K_2O$  vs.  $SiO_2$ . Flows 1, 2, 3, 5 and 7 fall in the field of alkaline basalts whereas Flows 4, 6 and 8 fall near the demarcating line. The upper most flow F 9 falls within the subalkaline field. Data points of alkali basalts (half filled circles) from Bhujia hill, Kutch (Pande 1988) and tholeiites (half filled circles) from Saurashtra region (Shekhawat and Sharma 1996) are also shown for comparison (b). Data points of Anjar flows in Nb/Y vs.  $Zr/TiO_2 \times 10^{-4}$  diagram.

the magma during its passage through the crust, Th/Yb vs. Ta/Yb diagram (figure 7) gives another criterion for crustal contamination (Pearce 1980). Based on all these criteria, we classify Anjar basalts (except F-9) as alkaline and F-9 as tholeiitic.

It is clear from figure 7 that all the Anjar flows plot in the enriched mantle source zone near the region of intraplate basalts, parallel to vector W. The within-plate origin of Anjar flows is further confirmed by using the discrimination diagram of Pearce and Cann (1973) based on Ti-Zr-Y, which they used to distinguish between ocean or continental flood basalts from other basalt types (figure 8a). We find that the Anjar flows, except F-9, plot



Figure 6. (a). Data points of Anjar flows in  $TiO_2$  vs.  $Zr/P_2O_5 \times 10^4$  diagram; (b). Data points of Anjar flows in Nb/Y vs.  $Zr/P_2O_5 \times 10^4$  diagram.



Figure 7. Th/Yb vs. Ta/Yb diagram of Pearce (1980). The Anjar flows plot within the enriched mantle zone. The vectors shown indicate the influence of subduction component (S), within plate enrichment (W), crustal contamination (C), and fractional crystallization (F).



Figure 8(a). The discrimination diagram Ti-Zr-Y of Pearce and Cann (1973) showing that Anjar basalts lie within the within-Plate Basalt field [D].

within field D of within-plate basalts. The Zr/Y vs. Zr diagram of Pearce and Norry (1979) also confirms their within-plate origin (figure 8b).

The Ti - Cr plot (after Pearce *et al* 1975) shows that all the Anjar flows plot on the Ocean floor Basalt (OFB) field (figure 8c). This is further confirmed by the Hf-Th-Nb/16 (or Ta) diagram of Wood (1980), that differentiates various types of basalts of oceanic environment from each other (figure 8d). It can be noted that the alkaline flows (F-1, 2, 3, 5 and 7) plot in the field of alkaline within-plate and the subalkaline flows F-4, F-6 and F-8 plot at the transition zone of E-MORB (enriched mid oceanic ridge basalt) and alkali within-plate field. On the other hand, F-9 plots in the E-MORB field. Meschde (1986) has suggested that the incompatible element Nb can be used to separate different types of ocean floor basalts and also distinguish different types (Enriched, Plume or Normal ) of MORBs. We reproduce this discrimination diagram in figure 8(e) from which we conclude that all the Anjar flows lie in within-plate alkali basalt field. The ocean island alkali basaltic nature of most of the Anjar flows is confirmed by the major element  $(MnO-P_2O_5 - TiO_2)$  discrimination diagram (figure 8f, after Mullen 1983).

Thus, based on all the major, minor and trace element data discussed above, we conclude that the Anjar flows have originated within plate region in oceanic environment i.e. they have not been generated at plate margins. Rather these have origin similar to Ocean Island Basalt (OIB). Furthermore, the OIB normalized pattern of REE is flat, supporting the OIB nature (figure 2b). In table 7, we compare crucial trace elements like Ba, K, Rb, Sr, Nb and K/Ba in different types of Anjar basalts



Figure 8(b). The discrimination diagram for basalts based upon Zr vs. Zr/Y variation (after Pearce and Norry 1979). Anjar flows lie within the within-Plate Basalt field [A].



Figure 8(c). The Anjar basalt data points on the Cr vs. Ti discrimination diagram of Pearce *et al* (1975).

with MORB, OIT (Ocean Island tholeiites) and OIAB (Ocean island alkali basalts). These data as well as multi-element variation diagram i.e. spidergram (figure 9) also supports the conclusion that Anjar basalts (except F9) belong to OIAB group.

We use these data (tables 2, 3) to ascertain if there is any similarity between Anjar basalts and Réunion basalts, since the Réunion plume has been suggested as a source of Deccan volcanism (Morgan 1972, 1981). Y/Nb vs. Zr/Nb data plotted in figure 10 show that some of the Anjar basalts and Réunion basalts (Fisk *et al* 1988) overlap. The REE patterns of alkali basalts from Réunion Island A D Shukla et al





Figure 8(d). The Hf-Th-Nb/16 discrimination diagram of Wood (1980). The various fields are: [A] N type MORB; [B] E type MORB and within-plate tholeiites; [C] alkaline within-plate basalts; [D] volcanic-arc basalts. The Anjar basalts 1, 2, 3, 5 and 7 lie in alkaline within-plate region while F-4 and 9 in E-type MORB and within-plate tholeiites field. The flows 6 and 8 fall at the transition line of field C and B.



Figure 8(e). The Zr-Y-Nb discrimination diagram of Meschede (1986), the various fields: **[AI]** within-plate alkali basalts; **[AII]** within-plate alkali and within-plate tholeiites; **[B]** E-type MORB; **[C]** within- plate tholeiites and volcanicarc basalts; **[D]** N type MORB and volcanic-arc basalts. The flows 1, 2, 3, 5, 6 and 8 of Anjar fall in within-plate alkali basalt region. Flow-9 lies in E -MORB field. Flow 4 lies near the boundary of fields AII and C.

Figure 8(f). The MnO-TiO<sub>2</sub>-P<sub>2</sub>O<sub>5</sub> discrimination diagram of basalts and basaltic andesites (after Mullen 1983). The fields are: MORB; **OIT**: ocean island tholeiite or seamount tholeiite; **OIA**: ocean -island alkali basalt or seamount alkali basalt; **IAT**: island-arc tholeiite; **CAB**: calc-alkaline basalt. Typical flows of Anjar area are plotted. All except F-9 lies in OIA or at close to the boundary of OIA and OIT whereas F-9 falls within MORB field.

(Fisk et al 1988) are similar to that observed in Anjar basalts. Sm/Nd and Ba/Nb ratios have been used by Peng *et al* (1994) to compare Deccan and Réunion basalts. Following the same arguments, we find that most of the Anjar basalt data scatter close to the Réunion data (figure 11). This argument is not a strong evidence in favour of the Réunion origin of flows but only shows chemical similarities between the Réunion basalts with the Anjar traps. This aspect needs to be studied further by using various isotopic criteria. The uppermost flow F 9, however, plots away from Ambenali, although Sm/Nd is similar. Isotopic systematics of Pb-Sr-Nd can confirm the true affinity of these basalts as has been done for the northeastern and southwestern flows of Deccan.

### 9. Summary and conclusions

The chemical, chronological and magnetic characteristics of the basalt sequence at Anjar are discussed and compared with the southwestern and northeastern flows. The results show that some of the Anjar basalts are alkaline and resemble ocean island alkali basalts in their major and trace element characteristics. The Anjar basalt sequence represents a rare section comprising alkali basalts capped by tholeiites. Similarities of some diagnostic trace element ratios with Réunion Island

Table 7. Comparison of selected trace element abundances of Anjar flows (except F-9) with MORB, ocean island tholeiites (OIT) and ocean island alkali basalts (OIAB).

	MORB	OIT	OIAB	Anjar Basalts
$\mathrm{K}_2\mathrm{O}~\%$	< 0.1 - 0.3	0.2 - 1.0	1-7	0.94 - 1.68
Ba (ppm)	5 - 50	70 - 200	200 - 1400	237.5 - 564.3
Sr (ppm)	90 - 200	150 - 400	400 - 4000	222.2 - 1839
Rb (ppm)	< 5	5 - 12	15 - 400	14.6 - 71
$\operatorname{Zr}(\operatorname{ppm})$	15 - 150	100 - 300	200 - 1000	179.5 - 297
Nb (ppm)	1 - 15	5 - 25	20 - 160	18.8 - 64
K/Ba	20 - 160	25 - 40	$\sim 28$	24.5 - 46.7



Figure 9. Typical multi element Spidergram for Anjar flows normalized to (a) Primitive mantle and (b) average Ocean Island Basalts.

basalts are consistent with their plume origin in the Réunion hot spot but these chemical arguments are not unique or strong enough to prove the Réunion plume origin of Deccan. The ages of Anjar basalts lie in the same range as the ages of southwest-



Figure 10. The variation of Y/Nb and Zr/Nb in the Anjar flows showing similarity to Réunion basalts (half filled circles, Fisk *et al* 1988).

ern and northeastern flows of Deccan Traps (68–61 Ma) and they constitute a N-R and possibly a N-R-N-R magnetic polarity sequence, similar to the northeastern flows (Sreenivasa Rao *et al* 1985). The upper most flow of Anjar (F 9) is similar in chemical composition to Ambenali basalts. Its REE pattern suggests that it is even less contaminated compared to the Ambenali group. This flow may have originated in a similar 'uncontaminated' magma pool; the possibility of lava flowing great distances (a few hundred km) has been discussed by Peng *et al* (1998).

The iridium content of Anjar basalts is high; F-2 shows the highest value (178pg/g) observed in Deccan Traps so far. However, since alkali basalts form only a small fraction of Deccan Traps, their contribution to the iridium inventory at K/T boundary may be insignificant and its main source may be related to a bolide impact. In view of the high iridium content, dinosaur bearing intertrappeans, N-R magnetic polarity sequence and wide age span, the Anjar volcano-sedimentary sequence appears



Figure 11. The Sm/Nd vs. Ba/Nb ratio plot reproduced from Peng *et al* (1994). The data of Anjar flows (open circles) except F-9 lie close to Réunion basalt data.

to be a unique section for studying the relation of Deccan volcanism to the K/T boundary. Further study of isotopic (Pb-Sr-Nd) systematics may be useful in understanding its origin and relation to the Réunion plume.

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