

ON THE EVALUATION OF THREE-CENTRE HELLMANN-FEYNMAN FORCE INTEGRALS INVOLVING SLATER-TYPE ATOMIC ORBITALS

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A combination of analytical and semi-numerical methods is suggested for the accurate evaluation of three-centre Hellmann-Feynman force integrals occurring in some studies of molecular systems. The analytical method combines the expansion techniques of Coulson (1937) and Bader (1962) while the semi-numerical method employs a double-zeta expansion (Barnett 1963). Explicit analytical expressions are listed for all required non-vanishing integrals involving up to  $2s$  and  $2p$  Slater-type atomic orbitals.

INTRODUCTION

In the last ten years considerable interest has been aroused in the chemical applications of the Hellmann-Feynman theorem (Hellmann 1937; Feynman 1939). This theorem states that if  $\psi$  is an exact (unnormalized) eigenfunction of a Hamiltonian  $H$  and  $E$  is the corresponding energy eigenvalue, then if  $\lambda$  is any parameter on which  $H$  or  $\psi$  depends

$$\frac{\partial E}{\partial \lambda} \langle \psi | \psi \rangle = \left\langle \psi \left| \frac{\partial H}{\partial \lambda} \right| \psi \right\rangle \quad \dots \quad (1)$$

In other words, for a normalized wave function, the first derivative of the energy with respect to a parameter is equal to the expectation value of the corresponding first derivative of the Hamiltonian. A proof of this theorem has been given by Berlin (1951) as well as by Hirschfelder and Coulson (1962). Depending on the physical nature of the parameter  $\lambda$  (e.g. it may be an inter-nuclear distance, a semi-empirical parameter in an approximate theory, a nuclear charge, a nuclear coordinate and so on), one can obtain much valuable information about variations in the energy of a molecular system with the help of the above theorem.

If  $\lambda$  is taken as the  $x$ -coordinate, say, of the  $p$ th nucleus in a molecule, then, in the framework of the adiabatic approximation, the Hellmann-Feynman theorem leads (Feynman 1939) to the following expression for the  $x$ -component,  $F_{px}$ , of the force acting on the  $p$ th nucleus

$$F_{px} = Z_p \left[ \sum_{a \neq p} \frac{Z_a}{R_{pa}^2} \frac{\partial R_{pa}}{\partial x_p} - \int \rho(r) \frac{x_p}{r_p^3} d\tau \right] \quad \dots \quad (2)$$

where  $Z_p$  is the nuclear charge of  $p$ ,  $R_{pq}$  an internuclear distance and  $\rho(\mathbf{r})$  is the total one-electron density at the point  $\mathbf{r}$ . The integration in (2) is to be done in the one-electron space. Equation (2) has seen interesting applications in recent years, for example, in viewing molecular shapes and chemical binding in terms of the forces on nuclei (Berlin 1951; Bader and Jones 1961, 1963; Bader *et al.* 1967; Coulson and Deb 1971).

It may be remarked here that the Hellmann-Feynman theorem is valid for exact wave functions, genuine self-consistent-field wave functions and certain other wave functions (Goddard 1968) which are not of Hartree-Fock type. But in practical applications of the theorem, employing present-day wave functions, the validity of the theorem with such approximate functions is tacitly assumed even though equation (2) has been found to be rather sensitive to small inaccuracies in the wave function.

If one chooses a single-determinantal wave function, then

$$\rho(\mathbf{r}) = \sum_i \eta_i \phi_i^2(\mathbf{r}) \quad \dots \quad (3)$$

where  $\phi_i$  is the  $i$ th MO (real) with occupation number  $\eta_i$ . If  $\phi_i$  is a conventional poly-centre LCAO-MO, then the calculation of the electron-nuclear component of the net force in equation (2) involves the evaluation of one-centre integrals  $\langle x_\mu | f_\mu | x'_\mu \rangle$ , two-centre integrals  $\langle x_\mu | f_\mu | x'_\nu \rangle$  and  $\langle x_\nu | f_\mu | x'_\nu \rangle$ , and the three-centre integrals  $\langle x_\mu | f_\kappa | x'_\nu \rangle$ , where  $x_\mu$  is an AO centred on  $\mu$  and  $f_\kappa$  is a force operator; the prime in  $x'_\nu$  indicates that this AO may have a functional form different from the AO  $x_\mu$ . The analytical expressions for many one- and two-centre integrals have been listed by Coulson (1942) as well as by Bader and Jones (1961). In this paper we are concerned with only the three-centre integrals.

#### SUMMARY OF DIFFERENT METHODS

In general, there are four techniques for reducing the three-centre integrals into single-centre ones. None of these, however, is completely satisfactory. These are: (i) expanding the force operator  $f_\kappa$  and the AO  $x'_\nu$  about the centre  $\mu$ , the former in terms of Gegenbauer polynomials (Magnus and Oberhettinger 1954) and the latter in terms of Legendre polynomials and modified Bessel functions (Watson 1958); (ii) expanding both the AO's,  $x_\mu$  and  $x'_\nu$ , about the centre  $\kappa$  of the force operator (double-zeta technique, Barnett 1963); (iii) transforming two of the distances,  $r_\mu$ ,  $r_\kappa$  and  $r_\nu$ , in terms of the third distance, bond length and bond angle (*see* Bader and Jones 1963); and (iv) the use of spheroidal coordinates based on the centres  $\mu$  and  $\nu$ , with  $f_\kappa$  being a directional derivative of the Neumann expansion of  $1/r_\kappa$  (*see* Appendix for details). Of these, (i) and (iv) are purely analytical methods for evaluating these integrals and involve no other approximation besides truncating an infinite series. The second method is a semi-numerical one where the integrals of products of

modified Bessel functions  $I_{m+1/2}(z)$  and  $K_{m+1/2}(z)$  are obtained numerically, and also involves terminating an infinite series. The third method is one of straightforward numerical integration up to a prescribed accuracy and is less efficient, computationally, than the analytical methods. We would like to suggest that for evaluating a three-centre force integral one may choose either (i) or (ii) depending on whether the integral, expressed as an infinite sum, converges or not. In general, (i) leads to convergence, although slowly, if the centres  $\kappa$  and  $\nu$  are equidistant from the centre of expansion  $\mu$ . Where this is not satisfied (i) results in non-convergence and one then may resort to (ii) which has a faster rate of convergence; for (i) convergence of the infinite sum also depends on the orbital exponent of  $x_\mu$ , a higher exponent resulting in faster convergence. However, whenever convergent (i) provides a rapid method for computing such integrals, involving very little additional labour for several other values of the valence angle involved; this is not so with the numerical and the semi-numerical methods.

#### THE ANALYTICAL AND SEMI-NUMERICAL METHODS

Fig. 1 depicts the coordinate system employed in these calculations. The integrals may be evaluated by choosing two of the cartesian axes to lie in

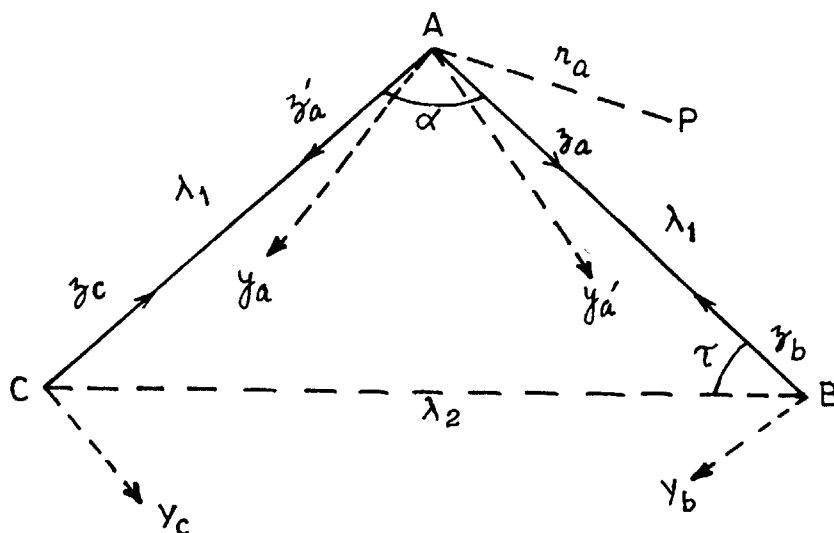


FIG. 1. Coordinate system for the evaluation of three-centre force integrals. The  $x$ -axes are perpendicular to the plane of the paper, and  $\lambda_1 < \lambda_2$ .

the plane containing the three centres involved. For this purpose the AO's and force vectors may be transformed, if necessary, by rotation to appropriate

basis sets. The analytical method of evaluating  $\langle x_\mu | f_\nu | x'_\nu \rangle$  consists of the following main steps:

- (1) expanding the radial function of  $x'_\nu$  about the centre  $\mu$ , as an infinite sum, in terms of Legendre functions and modified Bessel functions of first and second kinds,  $I_{m+1/2}(z)$  and  $K_{m+1/2}(z)$ , respectively (Coulson 1937; Barnett and Coulson 1951);
- (2) expanding  $f_\nu$  as another infinite sum, about the centre  $\mu$  in terms of products of Gegenbauer polynomials and certain radial functions (Bader 1962);
- (3) reduction of the double infinite sum to a single one by invoking the addition theorem for Legendre functions (see Magnus and Oberhettinger 1954).

Final simplifications are achieved with the help of various recurrence and orthogonality relations of the functions involved. A somewhat similar procedure is followed with the semi-numerical method except that the step (2) is to be replaced by a similar step like (1), since the centre of expansion for both the AO's is now that of the force operator  $f_\nu$ . We refrain here from providing an illustrative example for the above procedures since both Coulson (1937) and Bader (1962) have illustrated their individual techniques with simple representative examples. Therefore, we shall list in the next section only the explicit analytical expressions for the integrals obtained by combining these two techniques without indicating the tedious but simple algebra involved in deriving them.

The number of convergents to be taken for an infinite sum in an actual calculation depends on the prescribed accuracy. In order to obtain a value with four significant figures 20 convergents (sometimes less for method (ii)) usually suffice. For a test integral  $\langle h_c \left| \frac{y_b}{r_b^3} \right| 1s_a \rangle$  (see Fig. 1), where  $h_c$  and  $1s_a$  indicate Slater-type hydrogen and central atom  $1s$  orbitals respectively, the two procedures gave computed values agreeing up to the fourth significant figure. The agreement is expected to improve when one chooses a finer quadrature grid for the numerical integration in method (ii).

A word is needed about the use of a Mulliken-type approximation

$$x_\mu x_\nu \simeq \frac{1}{2} S_{\mu\nu} (x_\mu^2 + x_\nu^2) \quad \dots \quad \dots \quad \dots \quad \dots \quad (4)$$

which apparently might be applied to evaluate these three-centre integrals.

Bader and Jones (1963) have used (4) to calculate the integral  $\langle h_b \left| \frac{z_a''}{r_a^3} \right| h_c \rangle$  (see Fig. 1,  $z_a''$  bisects  $\angle BAC$ ). In Table I we compare the values, obtained by using (4), of this and another integral with the corresponding values obtained in the present calculations. We notice that for both the integrals the difference between the approximate and the accurate value—

which is quite appreciable—decreases as the three centres approach a linear configuration. This is presumably because of the fact that the approximation (4), in effect, distributes the overlap density on to the separate atoms. In classical language, this amounts to expressing the force vector directed to the centre of overlap density  $x_\mu x_\nu$  as the resultant of two vectors directed to the two centres  $\mu$  and  $\nu$ ; obviously this becomes more and more valid as the centres  $\kappa, \mu$  and  $\nu$  tend to linearity. For further discussion regarding approximation (4) we refer to Newton *et al.* (1966), Dahl and Ballhausen (1968) and Doggett and McKendrick (1968).

TABLE I

*The accuracy of Mulliken-type approximation (4)*  
H-atom exponent = 1.0; A-atom exponent = 2.275;  
 $\lambda_1 = 1.8101$  a.u.,  $z_a''$  bisects  $\angle BAC$  (Fig. 1)

Integral	$\left\langle h_c \left  \frac{z_b}{r_b^3} \right  2s_a \right\rangle$	$\left\langle h_b \left  \frac{z_a''}{r_a^3} \right  h_c \right\rangle$		
Valence angle $\alpha$ in deg.	Approximate	Accurate	Approximate	Accurate
90	0.05950	0.1048	0.06714	0.09855
100	0.05880	0.09944	0.05441	0.08513
105	0.05847	0.09706	0.04878	0.07874
110	0.05816	0.09487	0.04360	0.07255
120	0.05761	0.09103	0.03954	0.06073
180	0.05615	0.08105	0	0

## MASTER FORMULAE

In this section we list explicit analytical expressions for all required non-vanishing integrals involving up to  $2s$  and  $2p$  STO's with exponents  $\zeta_i$ . In cases where a scale factor has been employed to satisfy the virial theorem (Kaldor and Shavitt 1966) the orbital exponents and bond lengths are to be multiplied and divided, respectively, by this factor. Most of the quantities written below, viz.  $p_m, q_m, A_m, B_m, C_m$ , etc., together with their recurrence relations have been defined by Coulson\* (1937) while  $P_m(\cos \alpha), P_m^1(\cos \alpha)$  are the Legendre and associated Legendre functions respectively. In order to avoid the accumulation of round-off errors the basic functions  $A_m$  and  $B_m$  are to be built up *in decreasing order* of  $m$  by using sum-type recurrence relations instead of difference-type. For this purpose,  $A_m$  for a highest required

\* Coulson's (1937) relation,  $\rho^2 A_n F_{n-1} + A_{n-1} F_n = 2e^{-2\rho}$ , should read

$$\rho^2 A_{n-1} F_n + A_n F_{n-1} = 2e^{-2\rho} \quad \dots \quad (5)$$

$m$  may be computed, from which the corresponding  $B_m$  may be evaluated by using an asymptotic formula (Coulson 1937). For  $m = 24$  this procedure results in a  $B_0$  accurate up to the tenth significant figure.\*

\* For computing  $A_m$  for the highest  $m$  one should use, for the modified Bessel function  $I_{m+1/2}(z)$ , the formula

$$I_{m+1/2}(z) = \left(\frac{2z}{\pi}\right)^{1/2} z^m \sum_{j=0}^{\infty} \frac{(m+j)!}{j!(2m+2j+1)!} z^{m+2j} \dots \dots \dots (6)$$

The alternative closed formula,

$$I_{m+1/2}(z) = \frac{1}{(2\pi z)^{1/2}} \left[ e^z \sum_{j=0}^m \frac{(-1)^j (m+j)!}{j!(m-j)! 2z^j} + (-1)^{m+1} e^{-z} \sum_{j=0}^m \frac{(m+j)!}{j!(m-j)! 2z^j} \right] \dots (7)$$

which is mathematically equivalent to (6), was found to be unsuitable for computational purposes.

From Fig. 1 we notice that

$$\left. \begin{aligned} z_c &= \lambda_1 - z'_a \\ &= \lambda_1 - z_a \cos \alpha - y_a \sin \alpha \\ y_c &= y'_a \\ &= z_a \sin \alpha - y_a \cos \alpha \end{aligned} \right\} \dots \dots \dots (8)$$

There are eight integrals (see later) which could not be simplified like the others according to method (i). We, therefore, give their expressions according to method (ii); limitations of space prevent us from giving similar expressions for other integrals. The various integrals  $P_1(m)$ ,  $Q_1(m)$ ,  $R_1(m)$ , etc., may be computed numerically since they can be expressed as sums of integrals involving products of functions  $I_{m+1/2}(z)$  and  $K_{m+1/2}(z)$ .

Let  $\Lambda = \zeta_c \lambda_1$ ,  $\mu = \zeta_a / \zeta_c$ ,  $\lambda_1 < \lambda_2$ . One then replaces  $R$  and  $a$  in Coulson's (1937) quantities by  $\Lambda$  and  $\mu$  respectively. To reduce the final integral expressions to relatively simple forms we first define the following quantities:

- (A)  $c_m = F_{m-1}C_{m+1} - F_m C_{m+2}$ ;      $d_m = F_{m-1}D_{m+1} - F_m D_{m+2}$ ;
- $e_m = F_{m-1}E_{m+1} - F_m E_{m+2}$ ;      $g_m = A_m G_{m-1} - A_{m+1} G_m$ ;
- $h_m = A_m H_{m-2} - A_{m+1} H_{m-1}$ ;      $i_m = F_{m-1}I_m - F_m I_{m+1}$ ;
- $k_m = F_{m-1}K_{m+1} - F_m K_{m+2}$ ;      $l_m = F_{m-1}L_{m+1} - F_m L_{m+2}$ ;
- $m_m = A_m M_{m-3} - A_{m+1} M_{m-2}$ ;      $n_m = A_m N_{m-4} - A_{m+1} N_{m-3}$ ;
- $o_m = A_m O_{m-5} - A_{m+1} O_{m-4}$ ;      $s_m = A_m S_m - A_{m+1} S_{m+1}$ ;
- $t_m = A_m T_m - A_{m+1} T_{m+1}$ .

$$\begin{aligned}
 \text{(B) } c'_m &= F_m C_{m+1} - \frac{2(m+1)}{2m+1} F_{m+1} C_{m+2}; \\
 d'_m &= F_m D_{m+1} - \frac{2(m+1)}{2m+1} F_{m+1} D_{m+2}; \\
 e'_m &= F_m E_{m+1} - \frac{2(m+1)}{2m+1} F_{m+1} E_{m+2}; \\
 e''_m &= F_m E_{m+1} - \frac{2m}{2m+1} F_{m-1} E_m; \\
 g'_m &= A_m G_m - \frac{2m}{2m+1} A_{m-1} G_{m-1}; \\
 g''_m &= A_m G_m - \frac{2(m+1)}{2m+1} A_{m+1} G_{m-1}; \\
 h'_m &= A_m H_{m-1} - \frac{2m}{2m+1} A_{m-1} H_{m-2}; \\
 h''_m &= A_m H_{m-1} - \frac{2(m+1)}{2m+1} A_{m+1} H_m; \\
 i'_m &= F_m I_m - \frac{2m}{2m+1} F_{m-1} I_{m-1}; & i''_m &= F_m I_m - \frac{2(m+1)}{2m+1} F_{m+1} I_{m+1}; \\
 k'_m &= F_m K_{m+1} - \frac{2m}{2m+1} F_{m-1} K_m; & l'_m &= F_m L_{m+1} - \frac{2m}{2m+1} F_{m-1} L_m; \\
 m'_m &= A_m M_{m-2} - \frac{2(m+1)}{2m+1} A_{m+1} M_{m-1}; & m''_m &= A_m M_{m-2} - \frac{2m}{2m+1} A_{m-1} M_{m-3}; \\
 n'_m &= A_m N_{m-3} - \frac{2(m+1)}{2m+1} A_{m+1} N_{m-2}; & o'_m &= A_m O_{m-4} - \frac{2(m+1)}{2m+1} A_{m+1} O_{m-3}; \\
 s'_m &= A_m S_{m+1} - \frac{2m}{2m+1} A_{m-1} S_m; & t'_m &= A_m T_{m+1} - \frac{2m}{2m+1} A_{m-1} T_m,
 \end{aligned}$$

where

$$\begin{aligned}
 I_m &= (2\pi)^{1/2} P_{m+9/2, m+1/2}, & I_m &= D_{m+3} + (2m+3)D_{m+2}; \\
 K_m &= (2\pi)^{1/2} P_{m+9/2, m-1/2}, & K_m &= E_{m+2} + (2m+1)E_{m+1}; \\
 L_m &= (2\pi)^{1/2} P_{m+11/2, m-1/2}, & L_m &= I_{m+1} + (2m+1)I_m; \\
 O_m &= \left(\frac{2}{\pi}\right)^{1/2} Q_{m+1/2, m+9/2}, & O_m &= M_m + (2m+7)M_{m+1}; \\
 S_m &= \left(\frac{2}{\pi}\right)^{1/2} Q_{m+3/2, m-1/2}, & S_{m+1} &= \frac{1}{2m+1} (G_{m+1} - S_m), \\
 S_1 &= \left(\frac{1}{2\Lambda^2} - \frac{\mu+1}{2\Lambda}\right) e^{-(\mu+1)\Lambda} + \frac{1}{2}(\mu+1)^2 G_0; \\
 T_m &= \left(\frac{2}{\pi}\right)^{1/2} Q_{m+1/2, m-1/2}, & T_{m+1} &= \frac{1}{2m+1} (H_m - T_m), \\
 T_1 &= \frac{e^{-(\mu+1)\Lambda}}{\Lambda} - (\mu+1)G_0 \text{ [Coulson (1937)]}
 \end{aligned}$$

The various non-vanishing three-centre force integrals may now be expressed as the following infinite sums:

Let

$$\frac{y_b}{r_b^3} \equiv f_b, \quad \frac{z_b}{r_b^3} \equiv g_b, \quad \frac{x_b}{r_b^3} \equiv \pi_b.$$

$$(1) \langle 1s_c | f_b | 1s_a \rangle \\ = 2\zeta_c^2 \mu^{3/2} \frac{e^{\mu\Lambda}}{\Lambda^2} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) [c_m + \Lambda g_m]$$

$$(2) \langle 1s_c | g_b | 1s_a \rangle \\ = 4 \left( \frac{\zeta_c^3 \zeta_a^3}{\lambda_1 \lambda_2} \right)^{1/2} \sum_{m=0}^{\infty} [(m+1) P_{m+1}(\cos \tau) \{P_1(m) + Q_1(m) + R_1(m)\} \\ + m P_{m-1}(\cos \tau) \{P_2(m) + Q_2(m) + R_2(m)\}]$$

where

$$P_1(m) = \int_0^{\lambda_1} \frac{1}{r} p_m^<(\zeta_a, r; \lambda_1) p_{m+1}^<(\zeta_c, r; \lambda_2) dr,$$

$$Q_1(m) = \int_{\lambda_1}^{\lambda_2} \frac{1}{r} p_m^>(\zeta_a, r; \lambda_1) p_{m+1}^<(\zeta_c, r; \lambda_2) dr,$$

$$R_1(m) = \int_{\lambda_2}^{\infty} \frac{1}{r} p_m^>(\zeta_a, r; \lambda_1) p_{m+1}^>(\zeta_c, r; \lambda_2) dr.$$

$P_2(m)$ ,  $Q_2(m)$ ,  $R_2(m)$  are obtained by replacing  $p_{m+1}$  in  $P_1(m)$ ,  $Q_1(m)$ ,  $R_1(m)$ , respectively, by  $p_{m-1}$ ; < and > indicate whether  $r$  is less or greater, respectively, than the bond length argument ( $\lambda_1$  or  $\lambda_2$ ) of  $p_k$ .

$$(3) \langle 1s_c | f_b | 2s_a \rangle \\ = \frac{2}{3} \zeta_c^2 (3\mu^5)^{1/2} \frac{e^{\mu\Lambda}}{\Lambda^2} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) [d_m + \Lambda h_m]$$

$$(4) \langle 1s_c | g_b | 2s_a \rangle \\ = 4 \left( \frac{\zeta_c^3 \zeta_a^5}{3\lambda_1 \lambda_2} \right)^{1/2} \sum_{m=0}^{\infty} [(m+1) P_{m+1}(\cos \tau) \{P_3(m) + Q_3(m) + R_3(m)\} \\ + m P_{m-1}(\cos \tau) \{P_4(m) + Q_4(m) + R_4(m)\}]$$

where  $P_3(m)$ ,  $R_4(m)$  are obtained by replacing  $p_m$  in  $P_1(m)$ ,  $R_2(m)$ , respectively, with  $q_m$ .

$$(5) \langle 1s_c | f_b | 2pz_a \rangle \\ = 2\zeta_c^2 \mu^{5/2} \frac{e^{\mu\Lambda}}{\Lambda^3} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) \left[ \frac{m+2}{2m+3} (e_m + \Lambda^3 g_m) + \frac{m-1}{2m-1} (\Lambda m_m + \Lambda^2 c_m) \right]$$

$$(6) \langle 1s_c | g_b | 2pz_a \rangle \\ = 4 \left( \frac{\zeta_c^3 \zeta_a^5}{\lambda_1 \lambda_2} \right)^{1/2} (\lambda_1 I_1 - I_2)$$

where  $I_1$  is given by the infinite sum of integral (2) above and

$$I_2 = \sum_{m=0}^{\infty} \left[ \frac{(m+1)(m+2)}{2m+3} P_{m+2}(\cos \tau) \{P_5(m) + Q_5(m) + R_5(m)\} \right. \\ \left. + \frac{4m^3 + 6m^2 - 1}{(2m-1)(2m+3)} P_m(\cos \tau) \{P_6(m) + Q_6(m) + R_6(m)\} \right. \\ \left. + \frac{m(m-1)}{2m-1} P_{m-2}(\cos \tau) \{P_7(m) + Q_7(m) + R_7(m)\} \right]$$



where

$$P_5(m) = \int_0^{\lambda_1} p_m^<(\zeta_a, r; \lambda_1) p_{m+2}^<(\zeta_c, r; \lambda_2) dr,$$

$$Q_5(m) = \int_{\lambda_1}^{\lambda_2} p_m^>(\zeta_a, r; \lambda_1) p_{m+2}^<(\zeta_c, r; \lambda_2) dr,$$

$$R_5(m) = \int_{\lambda_2}^{\infty} p_m^>(\zeta_a, r; \lambda_1) p_{m+2}^>(\zeta_c, r; \lambda_2) dr,$$

and  $P_6(m)$ ,  $P_7(m)$ ,  $Q_7(m)$ , etc., are obtained by replacing  $p_{m+2}$  in the corresponding above quantities by  $p_m$  and  $p_{m-2}$ .

$$(7) \quad \langle 1s_c | f_b | 2p_y a \rangle = \langle 1s_c | \pi_b | 2p_x a \rangle \\ = \zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^3} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{(m+1)(m+2)}{2m+3} (e_m + \Lambda^3 g_m) - \frac{m(m-1)}{2m-1} (\Lambda m_m + \Lambda^2 c_m) \right]$$

$$(8) \quad \langle 1s_c | g_b | 2p_y a \rangle \\ = 2 \zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^3} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) \left[ \frac{m}{2m-1} (\Lambda m_m + \Lambda^2 c_m) - \frac{m+2}{2m+3} (e_m + \Lambda^3 g_m) \right]$$

$$(9) \quad \langle 2s_c | f_b | 1s_a \rangle \\ = \frac{2}{3} \zeta_c^2 (3\mu^3)^{1/2} \frac{e^{\mu \Lambda}}{\Lambda^2} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) \left[ e'_m + \Lambda m'_m + \Lambda^2 c'_m + \Lambda^3 g_m \right]$$

$$(10) \quad \langle 2s_c | g_b | 1s_a \rangle \\ = 4 \left( \frac{\zeta_c^5 \zeta_a^3}{3 \lambda_1 \lambda_2} \right)^{1/2} \sum_{m=0}^{\infty} [(m+1) P_{m+1}(\cos \tau) \{P_8(m) + Q_8(m) + R_8(m)\} \\ + m P_{m-1}(\cos \tau) \{P_9(m) + Q_9(m) + R_9(m)\}]$$

where  $P_8(m)$ ,  $Q_9(m)$ , etc., are obtained by replacing  $p_{m+1}$  in  $P_1(m)$ ,  $p_{m-1}$  in  $Q_2(m)$ , etc., by  $q_{m+1}$  and  $q_{m-1}$  respectively.

$$(11) \quad \langle 2s_c | f_b | 2s_a \rangle \\ = \frac{2}{3} \zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^2} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) [i'_m + \Lambda n'_m + \Lambda^2 d'_m + \Lambda^3 h'_m]$$

$$(12) \quad \langle 2s_c | g_b | 2s_a \rangle \\ = \frac{4}{3} \left( \frac{\zeta_c^5 \zeta_a^5}{\lambda_1 \lambda_2} \right)^{1/2} \sum_{m=0}^{\infty} [(m+1) P_{m+1}(\cos \tau) \{P_{10}(m) + Q_{10}(m) + R_{10}(m)\} \\ + m P_{m-1}(\cos \tau) \{P_{11}(m) + Q_{11}(m) + R_{11}(m)\}]$$

where  $P_{10}(m)$ ,  $Q_{11}(m)$ , etc., are obtained by replacing  $p_{m+1}$  in  $P_3(m)$ ,  $p_{m-1}$  in  $Q_4(m)$ , etc., by  $q_{m+1}$  and  $q_{m-1}$  respectively.

$$(13) \quad \langle 2s_c | f_b | 2p_z a \rangle \\ = \frac{2}{3} \zeta_c^2 (3\mu^5)^{1/2} \frac{e^{\mu \Lambda}}{\Lambda^3} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) \left[ \frac{m+2}{2m+3} (k'_m + \Lambda^2 e'_m + \Lambda^3 m'_m + \Lambda^5 g'_m) \right. \\ \left. + \frac{m-1}{2m-1} (\Lambda o'_m + \Lambda^2 e''_m + \Lambda^3 m''_m + \Lambda^4 c'_m) \right]$$

$$(14) \langle 2s_c | g_b | 2pz_a \rangle = 4 \left( \frac{\zeta_c^5 \zeta_a^5}{3\lambda_1 \lambda_2} \right)^{1/2} (\lambda_1 I_3 - I_4)$$

where  $I_3$  is given by the infinite sum in integral (10) and

$$I_4 = \sum_{m=0}^{\infty} \left[ \frac{(m+1)(m+2)}{2m+3} P_{m+2}(\cos \tau) \{P_{12}(m) + Q_{12}(m) + R_{12}(m)\} \right. \\ \left. + \frac{4m^3 + 6m^2 - 1}{(2m-1)(2m+3)} P_m(\cos \tau) \{P_{13}(m) + Q_{13}(m) + R_{13}(m)\} \right. \\ \left. + \frac{m(m-1)}{2m-1} P_{m-2}(\cos \tau) \{P_{14}(m) + Q_{14}(m) + R_{14}(m)\} \right]$$

where

$$P_{12}(m) = \int_0^{\lambda_1} p_m^<(\zeta_a, r; \lambda_1) q_{m+2}^<(\zeta_c, r; \lambda_2) dr,$$

$$Q_{12}(m) = \int_{\lambda_1}^{\lambda_2} p_m^>(\zeta_a, r; \lambda_1) q_{m+2}^<(\zeta_c, r; \lambda_2) dr,$$

$$R_{12}(m) = \int_{\lambda_2}^{\infty} p_m^>(\zeta_a, r; \lambda_1) q_{m+2}^>(\zeta_c, r; \lambda_2) dr$$

and  $P_{13}(m)$ ,  $Q_{14}(m)$ , etc., are obtained by replacing  $q_{m+2}$  in the above corresponding quantities by  $q_m$  and  $q_{m-2}$  respectively.

$$(15) \langle 2s_c | f_b | 2py_a \rangle = \langle 2s_c | \pi_b | 2px_a \rangle$$

$$= \frac{1}{3} \zeta_c^2 (3\mu^5)^{1/2} \frac{e^{\mu\Lambda}}{\Lambda^3} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{(m+1)(m+2)}{2m+3} (\kappa'_m + \Lambda^2 e'_m + \Lambda^3 m'_m + \Lambda^5 g'_m) \right. \\ \left. - \frac{m(m-1)}{2m-1} (\Lambda o'_m + \Lambda^2 e''_m + \Lambda^3 m''_m + \Lambda^4 c'_m) \right]$$

$$(16) \langle 2s_c | g_b | 2py_a \rangle$$

$$= \frac{2}{3} \zeta_c^2 (3\mu^5)^{1/2} \frac{e^{\mu\Lambda}}{\Lambda^3} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{m}{2m-1} (\Lambda o'_m + \Lambda^2 e''_m + \Lambda^3 m''_m + \Lambda^4 c'_m) \right. \\ \left. - \frac{m+2}{2m+3} (\kappa'_m + \Lambda^2 e'_m + \Lambda^3 m'_m + \Lambda^5 g'_m) \right]$$

$$(17) \langle 2pz_c | f_b | 1s_a \rangle$$

$$= \Lambda \langle 1s_c^{(2)} | f_b | 1s_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | 2pz_a^{(1)} \rangle - \sin \alpha \langle 1s_c^{(2)} | f_b | 2py_a^{(1)} \rangle$$

The forms of all these three integrals have been given above; we note that for the last two integrals  $\mu^{5/2}$  is now replaced by  $\mu^{3/2}$ . The superscript (1) indicates that the exponent  $\zeta_a$  is that of  $1s_a$  while (2) indicates  $\zeta_c$  to be that of a  $2p$  AO.

$$(18) \langle 2pz_c | g_b | 1s_a \rangle$$

$$= \Lambda \langle 1s_c^{(2)} | g_b | 1s_a \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | 2pz_a^{(1)} \rangle - \sin \alpha \langle 1s_c^{(2)} | g_b | 2py_a^{(1)} \rangle$$

$$(19) \langle 2py_c | f_b | 1s_a \rangle = \sin \alpha \langle 1s_c^{(2)} | f_b | 2pz_a^{(1)} \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | 2py_a^{(1)} \rangle$$

$$(20) \langle 2py_c | g_b | 1s_a \rangle = \sin \alpha \langle 1s_c^{(2)} | g_b | 2pz_a^{(1)} \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | 2py_a^{(1)} \rangle$$

$$(21) \langle 2pz_c | f_b | 2s_a \rangle$$

$$= \Lambda \langle 1s_c^{(2)} | f_b | 2s_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | r_a 2pz_a^{(2)} \rangle - \sin \alpha \langle 1s_c^{(2)} | f_b | r_a 2py_a^{(2)} \rangle$$

where  $2pz_a^{(2)}$ ,  $2py_a^{(2)}$  indicate the normalizing constant for the  $2s_a$  AO. The form of the first integral has been given previously and

$$\begin{aligned} & \langle 1s_c^{(2)} | f_b | r_a 2pz_a^{(2)} \rangle \\ &= \frac{2}{\sqrt{3}} \zeta_c^2 \mu^{5/2} \frac{e^{\mu \wedge}}{\Lambda^3} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) \left[ \frac{m+2}{2m+3} (i_m + \wedge^3 h_m) + \frac{m-1}{2m-1} (\wedge n_m + \wedge^2 d_m) \right], \\ & \langle 1s_c^{(2)} | f_b | r_a 2py_a^{(2)} \rangle \\ &= \frac{1}{\sqrt{3}} \zeta_c^2 \mu^{5/2} \frac{e^{\mu \wedge}}{\Lambda^3} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{(m+1)(m+2)}{2m+3} (i_m + \wedge^3 h_m) - \frac{m(m-1)}{2m-1} (\wedge n_m + \wedge^2 d_m) \right] \end{aligned}$$

$$(22) \quad \langle 2pz_c | g_b | 2s_a \rangle = \wedge \langle 1s_c^{(2)} | g_b | 2s_a \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | r_a 2pz_a^{(2)} \rangle - \sin \alpha \langle 1s_c^{(2)} | g_b | r_a 2py_a^{(2)} \rangle$$

where the first integral has been expressed before and

$$\langle 1s_c^{(2)} | g_b | r_a 2pz_a^{(2)} \rangle = 4 \left( \frac{\zeta_c^5 \zeta_a^5}{3 \lambda_1 \lambda_2} \right)^{1/2} (\lambda_1 I_5 - I_6)$$

where  $I_5$  is given by the infinite sum in integral (4) and

$$\begin{aligned} I_6 = \sum_{m=0}^{\infty} & \left[ \frac{(m+1)(m+2)}{2m+3} P_{m+2}(\cos \tau) \{ P_{15}(m) + Q_{15}(m) + R_{15}(m) \} \right. \\ & + \frac{4m^3 + 6m^2 - 1}{(2m-1)(2m+3)} P_m(\cos \tau) \{ P_{16}(m) + Q_{16}(m) + R_{16}(m) \} \\ & \left. + \frac{m(m-1)}{2m-1} P_{m-2}(\cos \tau) \{ P_{17}(m) + Q_{17}(m) + R_{17}(m) \} \right] \end{aligned}$$

where

$$P_{15}(m) = \int_0^{\lambda_1} q_m^<(\zeta_a, r; \lambda_1) p_{m+2}^<(\zeta_c, r; \lambda_2) dr,$$

$$Q_{15}(m) = \int_{\lambda_1}^{\lambda_2} q_m^>(\zeta_a, r; \lambda_1) p_{m+2}^<(\zeta_c, r; \lambda_2) dr,$$

$$R_{15}(m) = \int_{\lambda_2}^{\infty} q_m^>(\zeta_a, r; \lambda_1) p_{m+2}^>(\zeta_c, r; \lambda_2) dr,$$

and  $P_{16}(m)$ ,  $R_{17}(m)$ , etc., are obtained by replacing  $p_{m+2}$  in the corresponding quantities by  $p_m$ ,  $p_{m-2}$ , etc., respectively.

$$\begin{aligned} & \langle 1s_c^{(2)} | g_b | r_a 2py_a^{(2)} \rangle \\ &= \frac{2}{3} \zeta_c^2 (3\mu^5)^{1/2} \frac{e^{\mu \wedge}}{\Lambda^3} \sum_{m=0}^{\infty} P_m^1(\cos \alpha) \left[ \frac{m}{2m-1} (\wedge n_m + \wedge^2 d_m) - \frac{m+2}{2m+3} (i_m + \wedge^3 h_m) \right] \end{aligned}$$

$$(23) \quad \langle 2py_c | f_b | 2s_a \rangle = \sin \alpha \langle 1s_c^{(2)} | f_b | r_a 2pz_a^{(2)} \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | r_a 2py_a^{(2)} \rangle$$

$$(24) \quad \langle 2py_c | g_b | 2s_a \rangle = \sin \alpha \langle 1s_c^{(2)} | g_b | r_a 2pz_a^{(2)} \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | r_a 2py_a^{(2)} \rangle$$

$$(25) \quad \langle 2pz_c | f_b | 2pz_a \rangle = \wedge \langle 1s_c^{(2)} | f_b | 2pz_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | z_a 2pz_a \rangle - \sin \alpha \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle$$

The first integral on the R.H.S. has been expressed previously.

We have

$$\begin{aligned}
 & \langle 1s_c^{(2)} | f_b | z_a 2pz_a \rangle \\
 &= 2\zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^5} \sum_{m=0}^{\infty} P_{m+3}(\cos \alpha) \left[ \frac{(m-1)^2}{(2m-1)(2m+3)} (\Lambda^3 n_m + \Lambda^4 d_m) \right. \\
 & \quad + \frac{m-1}{2m-1} (\Lambda^3 e_m + \Lambda^4 m_m) + \frac{(m+1)}{(2m-1)(2m+3)} (\Lambda^2 i_m + \Lambda^5 l_m) \\
 & \quad \left. + \frac{m+2}{2m+3} (\Lambda k_m + \Lambda^6 g_m) - \frac{(m+2)(m+3)}{(2m+3)(2m+5)} (l_m + \Lambda^7 k_m) \right] \\
 & \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle \\
 &= 2\zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^4} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{(m+1)(m+2)(m+3)}{(2m+3)(2m+5)} (k_m + \Lambda^5 g_m) \right. \\
 & \quad \left. + \frac{m(m-1)}{(2m-1)(2m+3)} (\Lambda^2 e_m + \Lambda^3 m_m) - \frac{m(m-1)(m-2)}{(2m-1)(2m-3)} (\Lambda^0 o_m + \Lambda^4 c_m) \right] \\
 (26) \quad & \langle 2pz_c | g_b | 2pz_a \rangle \\
 &= \langle \Lambda^0 1s_c^{(2)} | g_b | 2pz_a \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | z_a 2pz_a \rangle - \sin \alpha \langle 1s_c^{(2)} | g_b | y_a 2pz_a \rangle.
 \end{aligned}$$

The first of these three integrals has already been expressed and

$$\langle 1s_c^{(2)} | g_b | z_a 2pz_a \rangle = 4 \left( \frac{\zeta_c^5 \zeta_a^5}{\lambda_1 \lambda_2} \right)^{1/2} [\lambda_1^2 I_1 - 2\lambda_1 I_2 + I_7]$$

where  $I_1$  and  $I_2$  have been given in (6) and

$$\begin{aligned}
 I_7 = \sum_{m=0}^{\infty} & \left[ \frac{(m+1)(m+2)(m+3)}{(2m+3)(2m+5)} P_{m+3}(\cos \tau) \{P_{18}(m) + Q_{18}(m) + R_{18}(m)\} \right. \\
 & + \frac{12m^5 + 56m^4 + 81m^3 + 40m^2 - 9}{(2m-1)(2m+1)(2m+3)(2m+5)} P_{m+1}(\cos \tau) \{P_{19}(m) + Q_{19}(m) + R_{19}(m)\} \\
 & + \frac{3m(m^2-2)}{(2m-3)(2m+3)} P_{m-1}(\cos \tau) \{P_{20}(m) + Q_{20}(m) + R_{20}(m)\} \\
 & \left. + \frac{m(m-1)(m-2)}{(2m-1)(2m-3)} P_{m-3}(\cos \tau) \{P_{21}(m) + Q_{21}(m) + R_{21}(m)\} \right]
 \end{aligned}$$

where

$$P_{18}(m) = \int_0^{\lambda_1} r p_m^<(\zeta_a, r; \lambda_1) p_{m+3}^<(\zeta_c, r; \lambda_2) dr,$$

$$Q_{18}(m) = \int_{\lambda_1}^{\lambda_1} r p_m^>(\zeta_a, r; \lambda_1) p_{m+3}^<(\zeta_c, r; \lambda_2) dr,$$

$$R_{18}(m) = \int_{\lambda_2}^{\infty} r p_m^>(\zeta_a, r; \lambda_1) p_{m+3}^>(\zeta_c, r; \lambda_2) dr,$$

and  $P_{19}(m)$ ,  $Q_{20}(m)$ ,  $R_{21}(m)$ , etc., are obtained by replacing  $p_{m+3}$  in the corresponding quantities by  $p_{m+1}$ ,  $p_{m-1}$ ,  $p_{m-3}$ , etc., respectively.

$$\begin{aligned}
 & \langle 1s_c^{(2)} | g_b | y_a 2pz_a \rangle \\
 &= -\langle 1s_c^{(2)} | f_b | z_a 2pz_a \rangle + 2\zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^5} \sum_{m=0}^{\infty} P_{m+1}(\cos \alpha) \left[ \frac{m-1}{2m-1} (\Lambda^2 o_m + \Lambda^6 c_m) \right. \\
 & \quad \left. + \frac{m+2}{2m+3} (\Lambda^3 e_m + \Lambda^4 m_m) \right]
 \end{aligned}$$

$$(27) \langle 2py_c | f_b | 2pz_a \rangle = \sin \alpha \langle 1s_c^{(2)} | f_b | z_a 2pz_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle$$

$$(28) \langle 2py_c | g_b | 2pz_a \rangle = \sin \alpha \langle 1s_c^{(2)} | g_b | z_a 2pz_a \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | y_a 2pz_a \rangle$$

$$(29) \langle 2pz_c | f_b | 2py_a \rangle = \langle 1s_c^{(2)} | f_b | 2py_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | y_a 2pz \rangle - \sin \alpha \langle 1s_c^{(2)} | f_b | y_a 2py_a \rangle.$$

The first two integrals on the R.H.S. have been given previously. The third is given by

$$\begin{aligned} & \langle 1s_c^{(2)} | f_b | y_a 2py_a \rangle \\ &= \frac{1}{2} \zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^4} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{2m(m+1)}{(2m-1)(2m+3)} (\Lambda^2 e_m + \Lambda^3 m_m) \right. \\ & \quad \left. - \frac{(m+2)(m+3)}{(2m+3)(2m+5)} (k_m + \Lambda^5 g_m) - \frac{(m-1)(m-2)}{(2m-1)(2m-3)} (\Lambda^0 o_m + \Lambda^4 c_m) \right] \end{aligned}$$

$$(30) \langle 2pz_c | g_b | 2py_a \rangle = \langle 1s_c^{(2)} | g_b | 2py_a \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | y_a 2pz_a \rangle - \sin \alpha \langle 1s_c^{(2)} | g_b | y_a 2py_a \rangle.$$

Only the third integral on the R.H.S. remains to be expressed, -

$$\begin{aligned} & \langle 1s_c^{(2)} | g_b | y_a 2py_a \rangle \\ &= \zeta_c^2 \mu^{5/2} \frac{e^{\mu \Lambda}}{\Lambda^4} \sum_{m=0}^{\infty} P_m(\cos \alpha) \left[ \frac{2(m+1)(m^2+m-1)}{(2m-1)(2m+3)} (\Lambda^2 e_m + \Lambda^3 m_m) \right. \\ & \quad \left. - \frac{m(m-1)^2}{(2m-1)(2m+3)} (\Lambda^0 o_m + \Lambda^4 c_m) - \frac{(m+1)(m+2)(m+3)}{(2m+3)(2m+5)} (k_m + \Lambda^5 g_m) \right] \end{aligned}$$

$$(31) \langle 2py_c | f_b | 2py_a \rangle = \sin \alpha \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | y_a 2py_a \rangle$$

$$(32) \langle 2py_c | g_b | 2py_a \rangle = \sin \alpha \langle 1s_c^{(2)} | g_b | y_a 2pz_a \rangle - \cos \alpha \langle 1s_c^{(2)} | g_b | y_a 2py_a \rangle$$

$$(33) \langle 2px_c | f_b | 2px_a \rangle = \langle 2px_c | \pi_b | 2py_a \rangle = \frac{1}{2} \langle 1s_c^{(2)} | f_b | y_a 2py_a \rangle$$

$$(34) \langle 2px_c | \pi_b | 1s_a \rangle = \langle 1s_c^{(2)} | f_b | 2py_a^{(1)} \rangle, \text{ with } \mu^{5/2} \text{ replaced by } \mu^{3/2}$$

$$(35) \langle 2px_c | \pi_b | 2s_a \rangle = \langle 1s_c^{(2)} | f_b | r_a 2py_a^{(2)} \rangle$$

$$(36) \langle 2px_c | \pi_b | 2pz_a \rangle = \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle$$

$$(37) \langle 2pz_c | \pi_b | 2px_a \rangle = \langle 1s_c^{(2)} | f_b | 2py_a \rangle - \cos \alpha \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle - \frac{1}{2} \sin \alpha \langle 1s_c^{(2)} | f_b | y_a 2py_a \rangle$$

$$(38) \langle 2py_c | \pi_b | 2px_a \rangle = \sin \alpha \langle 1s_c^{(2)} | f_b | y_a 2pz_a \rangle - \frac{1}{2} \cos \alpha \langle 1s_c^{(2)} | f_b | y_a 2py_a \rangle.$$

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where

$$c_{l_0} = 2l+1; \quad c_{lm} = (-1)^m 2(2l+1) \left[ \frac{(l-m)!}{(l+m)!} \right]^2, \quad m \geq 1.$$

$\xi_>$  and  $\xi_<$  are the greater and less, respectively, of  $\xi$  and  $\xi_a$ , and  $P_l^m$ ,  $Q_l^m$  are unnormalized Legendre functions of first and second kinds respectively. Thus, one obtains

$$f_a = -\frac{2}{\lambda_2} \sum_{l=0}^{\infty} \sum_{m=0}^l c_{lm} [P_l^m(\xi) P_l^m(\eta) \text{grad}_a \{Q_l^m(\xi_a) P_l^m(\eta_a) \cos m(\phi - \phi_a)\} \\ + Q_l^m(\xi) P_l^m(\eta) \text{grad}_a \{P_l^m(\xi_a) P_l^m(\eta_a) \cos m(\phi - \phi_a)\}] \quad (\text{A3})$$

where the first term within the square brackets refers to  $\xi < \xi_a$  and the second to  $\xi > \xi_a$ . Using (A3) one can reduce the original three-centre integral,  $\langle x_b | f_a | x_c \rangle$ , into an infinite sum of two-centre integrals in spheroidal coordinates. Since the integrals involving  $P_l^m(x)$  and  $Q_l^m(x)$  die away with increasing  $l$  the above infinite sum is expected to converge fairly rapidly.