Neutrino masses, leptogenesis and dark matter in hybrid seesaw

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We suggest a hybrid seesaw model where relatively "light" right-handed neutrinos give no contribution to the neutrino mass matrix due to a special symmetry. This allows their Yukawa couplings to the standard model particles to be relatively strong, so that the standard model Higgs boson can decay dominantly to a left and a right-handed neutrino, leaving another stable right-handed neutrino as cold dark matter. In our model neutrino masses arise via the type-II seesaw mechanism, the Higgs triplet scalars being also responsible for the generation of the matter-antimatter asymmetry via the leptogenesis mechanism.

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I. INTRODUCTION

Laboratory experiments with reactors and accelerator neutrinos have confirmed the observations of solar and atmospheric neutrinos [1], establishing the phenomenon of neutrino oscillations and hence the need for small but nonzero neutrino masses [2]. Since neutrinos are massless in the $SU(3)_c \times SU(2)_L \times U(1)_Y$ Standard Model (SM) this implies the need for new physics, whose detailed nature constitutes one of our deepest current challenges in particle physics [3]. The simplest extension of the SM to explain the neutrino masses is to include singlet righthanded neutrinos and/or triplet scalars [4]. The inclusion of the former is justified by the fact that there are righthanded partners for all other fermions of the SM. The singlet right-handed neutrinos will in general have Majorana masses as well as a Dirac mass term, of the order of the charged fermion masses. The latter arise from the usual Yukawa interaction between left- and right-handed leptons, once the SM Higgs doublet acquires a vacuum expectation value (vev). The terms involving the righthanded neutrinos are,

$$\mathcal{L} \supset -\frac{1}{2} (M_N)_{ij} \overline{N_{R_i}^c} N_{R_j} - h_{ij} \overline{\ell_{L_i}} \phi N_{R_j} + \text{H.c.}$$
(1)

Here N_R , ℓ_L and ϕ denote the right-handed neutrinos, the left-handed leptons and Higgs doublet, respectively.

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In the basis (ν_L, N_R^c) , the neutrino mass matrix reads,

$$M_{\nu} = \begin{pmatrix} 0 & m_D \\ m_D^T & M_N \end{pmatrix}, \qquad (2)$$

where $(m_D)_{ij} = h_{ij} \langle \phi \rangle$ with $\langle \phi \rangle \simeq 174$ GeV. After blockdiagonalization, one gets

$$m_{\nu} \simeq -m_D \frac{1}{M_N} m_D^T \,. \tag{3}$$

If the Majorana masses of the right-handed neutrinos are much larger than the Dirac masses, the left-handed neutrinos will naturally acquire a tiny Majorana through the type-I seesaw mechanism [4] [5, 6]. In order to account for neutrino masses in the eV range one needs,

$$h_{ij}h_{kj}(M_N)_{jk}^{-1} \sim 10^{-23} \text{ GeV}^{-1}$$
. (4)

so that the Majorana masses M_N are required to be orders of magnitude larger than the electroweak symmetry breaking scale $\sim \langle \phi \rangle$, unless the coefficients h_{ij} are very small,

$$h_{ij} \lesssim 10^{-11}$$
. (5)

There is an attractive way to avoid this conclusion in the framework of the inverse seesaw model [7, 8, 9]. In this case the right-handed neutrinos pair-off with extra singlet leptons to form Dirac-type neutral heavy leptons in such a way that their mixing with the doublet neutrinos does not lead to light neutrino masses in the limit of conserved lepton number. As a result of this "symmetryprotection" right-neutrinos can lie at the TeV scale and produce signals at accelerators [10, 11], without conflict with the observed smallness of neutrino masses.

Here we focus on a seesaw mechanism without additional fermion degrees of freedom. If one assumes

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"generic" violation of lepton number through righthanded neutrino Majorana masses, acceptable neutrino masses require very tiny effective Yukawa couplings connecting the right-handed neutrinos with the left-handed neutrinos, Eqs. (4) and (5), hence the right-handed neutrinos can not be produced at the LHC, nor can the Higgs boson decay into neutrinos. We explore the possibility that the neutrino mass m_{ν} can be exactly zero even if the Majorana mass M_{N} and the Dirac mass m_{D} are both non-vanishing [12, 13], as a result of a cancellation. We present a model where the type-I seesaw contribution to the neutrino masses vanishes identically due to a suitable symmetry, avoiding the main constraints on M_N and m_D . Neutrino masses arise from the type-II triplet seesaw mechanism [4]. As a result, relatively light righthanded neutrinos could have sizeable Yukawa couplings. If the Majorana masses of the right-handed neutrinos are of the order of TeV or less, they may be produced at the LHC. We also note that, if it is heavier than the right-handed neutrino, the Higgs boson in this scenario will also have a distinct decay channel into a left plus a right-handed neutrino. Our explicit symmetry will also protect one of the right-handed neutrinos from decaying, as it has no SM Yukawa interactions. This naturally accounts for a stable cold dark matter candidate [14, 15]. Finally, our scenario for the origin of neutrino masses also provides successful leptogenesis [16, 17] induced by the out-of-equilibrum decays of the heavy scalar triplets [18].

II. NEUTRINO MASS MATRICES

We shall first discuss the structure of the Dirac and Majorana masses of the neutrinos and then present the model in detail. To demonstrate the basic idea, consider a two-generation scenario in which the type-I seesaw matrix takes a particular form,

$$M_{\nu} = \begin{pmatrix} 0 & 0 & a & a \\ 0 & 0 & b & b \\ a & b & M & 0 \\ a & b & 0 & -M \end{pmatrix},$$
(6)

with $M \gg a, b$. Indeed, in this example, the two lefthanded neutrinos remain massless despite the coexistence of Dirac and Majorana mass terms, as a result of the cancelation between the contributions from the two righthanded neutrinos. The latter may be due to some underlying symmetry. In this case the Yukawa couplings thus can be large even if the right-handed neutrinos have "low" Majorana masses. Consider the mass matrix

$$M_{\nu} = \begin{pmatrix} 0 & 0 & 0 & a \\ 0 & 0 & 0 & b \\ 0 & 0 & 0 & M \\ a & b & M & 0 \end{pmatrix}$$
(7)

It is easy to see that (i) this matrix reduces to the previous after diagonalizing out the right handed states and (ii) it emerges from a Z_3 symmetry

$$N_{R_1} \to \omega^2 N_{R_1} , \quad N_{R_2} \to \omega N_{R_2} , \quad \nu_{L_i} \to \omega \nu_{L_i} ,$$
 (8)

where ω is the cube-root of 1, $\omega^3 = 1$ and $1 + \omega + \omega^2 = 0$. All other charged fermions also transform under this Z_3 symmetry as $f \to \omega f$ and the Higgs doublet ϕ is invariant, so that all the usual couplings allowed in the SM remain the same. The Lagrangian contributing to the neutrino masses is then given by

$$\mathcal{L} \supset -M\overline{N_{R_1}^c}N_{R_2} - h_{i2}\overline{\ell_{L_i}}\phi N_{R_2} + \text{H.c.}$$
(9)

which clearly leads to the form in Eq. (7) once the *vev* $\langle \phi \rangle$ is generated. This symmetry makes the light neutrinos remain exactly massless. In order to generate the required neutrino masses one may either break this Z_3 softly or, alternatively, introduce triplet scalars for implementing the type-II seesaw [4], without affecting the symmetry in the right-handed neutrino sector.

This mass matrix can be generalized to three generations. Consider the 3×3 mass matrix

$$\mathcal{M}_{\nu} = \begin{pmatrix} 0 & 0 & 0 & a & a & 0 \\ 0 & 0 & 0 & b & b & 0 \\ 0 & 0 & 0 & c & c & 0 \\ a & b & c & M & 0 & 0 \\ a & b & c & 0 & -M & 0 \\ 0 & 0 & 0 & 0 & 0 & \tilde{M} \end{pmatrix},$$
(10)

with $M, \tilde{M} \gg a, b, c$. Clearly, the three left-handed neutrinos remain exactly massless, while the masses of the right-handed neutrinos are $+M, -M, \tilde{M}$ can be "low" enough to be accessible to the LHC. Note, however, that the third-generation right-handed neutrino decouples. In another basis of the right-handed neutrinos one has,

$$\mathcal{M}_{\nu} = \begin{pmatrix} 0 & 0 & 0 & a & 0 & 0 \\ 0 & 0 & 0 & b & 0 & 0 \\ 0 & 0 & 0 & c & 0 & 0 \\ a & b & c & 0 & M & 0 \\ 0 & 0 & 0 & M & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & \tilde{M} \end{pmatrix} .$$
(11)

This mass matrix could emerge from a Z_3 symmetry

$$\begin{split} N_{R_1} &\to \omega N_{R_1} \,, \quad N_{R_2} \to \omega^2 N_{R_2} \,, \quad N_{R_3} \to N_{R_3} \,, \\ & \nu_{L_i} \to \omega \nu_{L_i} \,. \end{split} \tag{12}$$

or, alternatively, from an U(1) global symmetry. The Lagrangian containing the right-handed neutrinos,

$$\mathcal{L} \supset -M\overline{N_{R_1}^c}N_{R_2} - \frac{1}{2}\tilde{M}\overline{N_{R_3}^c}N_{R_3} - h_{i1}\overline{\ell_{L_i}}\phi N_{R_1} + \text{H.c.}(13)$$

implies, in the limit $M\gg h_{i1}\left<\phi\right>,$ two degenerate right-handed neutrinos

$$\frac{1}{\sqrt{2}}(N_{R_1} \pm N_{R_2}) \to N_{R_{1,2}} \tag{14}$$

with masses $\pm M$ (opposite CP signs). For $M \lesssim 1 \text{ TeV}$ these two states $N_{R_{1,2}}$ would be accessible to LHC searches. On the other hand, a more interesting possibility may open up when the Higgs boson is heavier than N_{R_1} and N_{R_2} . Given the mild constraint on the Dirac mass term, the Yukawa couplings could be large enough that the Higgs boson would dominantly decay into a left-handed neutrino and a right-handed one, posing a new challenge to the Higgs search program at the LHC. The third right-handed neutrino N_{R_3} has mass \tilde{M} and no decay modes. Hence it could serve as the dark matter if its relic density is consistent with the cosmological observations.

III. LEPTOGENESIS, NEUTRINO MASS AND DARK MATTER

We now propose a realistic model, which contains the previous phenomenology of the right-handed neutrinos and explains the matter-antimatter asymmetry, the neutrino masses and the dark matter. We extend the SM by including three right-handed neutrinos, two triplet Higgs scalars $\xi_{1,2} \equiv (\mathbf{1}, \mathbf{3}, 2)$ and two singlet scalar fields $\sigma, \chi \equiv (\mathbf{1}, \mathbf{1}, 0)$. In addition to the SM gauge symmetry, we also impose a global $U(1)_{lep}$ symmetry of lepton number, under which the different fields transform as:

N_{R_1}	N_{R_2}	N_{R_3}	ℓ_{L_i}	ξ_1	ξ_2	σ	χ	ϕ
1	3	2	1	-2	$^{-1}$	-4	1	0

TABLE I: Lepton number assignments. For simplicity, we do not show the right-handed charged leptons, which carry the same lepton number as their left-handed partners.

The relevant part of the Lagrangian is given as,

$$\mathcal{L} \supset -\alpha_1 \sigma \overline{N_{R_1}^c} N_{R_2} - \frac{1}{2} \alpha_2 \sigma \overline{N_{R_3}^c} N_{R_3} - h_{i1} \overline{\ell_{L_i}} \phi N_{R_1} - \frac{1}{2} f_{ij} \overline{\ell_{L_i}^c} i \tau_2 \xi_1 \ell_{L_j} - \alpha_3 \chi \phi^T i \tau_2 \xi_2 \phi - \mu \chi \xi_2^{\dagger} \xi_1 + \text{H.c.}$$
(15)

After the singlet scalar σ develops its *vev* the first line will induce the Lagrangian (13), so that the right-handed neutrinos obtain their Majorana masses. The second line will generate the type-II seesaw in the presence of $\langle \chi \rangle$.

In our model, the global $U(1)_{lep}$ is assumed to break at a very large scale by $\langle \chi \rangle \sim 10^{13} \,\text{GeV}$. The triplet scalars ξ_1 and ξ_2 , whose masses $\sim M_{\xi}$ are of the order of $\langle \chi \rangle$, mix with each other and pick up tiny *vevs* after the electroweak symmetry breaking,

$$\langle \xi_2 \rangle \sim -\alpha_3 \frac{\langle \chi \rangle \langle \phi \rangle^2}{M_{\xi}^2}, \quad \langle \xi_1 \rangle \sim -\frac{\mu \langle \chi \rangle \langle \xi_2 \rangle}{M_{\xi}^2}.$$
 (16)

These triplet *vevs* will give rise to the left-handed neutrinos Majorana mass matrix,

$$m_{\nu_{ij}} = f_{ij} \left< \xi_1 \right> \,. \tag{17}$$

The CP-violating and out-of-equilibrium decays of the triplet scalars ξ_1 and ξ_2 into the SM lepton and Higgs doublets can generate a lepton asymmetry [18]. The sphaleron [17] processes active in the range 100 GeV $\lesssim T \lesssim 10^{12}$ GeV, will partially convert this lepton asymmetry to a baryon asymmetry for explaining the matterantimatter asymmetry of the universe. In order to successfully induce leptogenesis and suppress washout processes, we require that the singlet scalar σ develops its *vev* after the sphaleron epoch is over, for example, we take $\langle \sigma \rangle \sim 100$ GeV. Through their Yukawa couplings to σ the right-handed neutrinos acquire their Majorana masses $M = \alpha_1 \langle \sigma \rangle$ and $\tilde{M} = \alpha_2 \langle \sigma \rangle$, expected to lie below 100 GeV or so.

The spontaneous breakdown of the global lepton number global symmetry through $\langle \chi \rangle$, leads to a Goldstone boson, whose profile can be determined by the symmetry [6], leading to,

$$\mathcal{G} = \frac{1}{N} \left[\langle \chi \rangle \operatorname{Im}(\chi) + \langle \xi_1 \rangle \operatorname{Im}(\xi_1^0) + \langle \xi_2 \rangle \operatorname{Im}(\xi_2^0) \right] .$$
(18)

where N is a suitable normalization, which is of the order of the lepton number breaking scale $\sim \langle \chi \rangle$. Clearly, the triplet component of the Goldstone boson is highly suppressed by the ratio of the triplet *vevs* over the singlet *vev*, suppressing its coupling to the Z-bosons [6].

Note that since lepton number conservation forbids the couplings of χ to the right-handed neutrinos, $\langle \chi \rangle$ will not have any effect on them. The states N_{R_1} and N_{R_2} mix maximally and in the basis where their Majorana mass matrix is diagonal, the degenerate states N_{R_1} and N_{R_2} in Eq. (14) with mass $\pm M$ both couple to the left-handed neutrinos through equal Yukawa couplings. In addition, the state N_{R_3} has no couplings to the left-handed neutrinos. Therefore, the resulting neutrino mass matrix is a

null matrix. This implies that the strongest constraints on the couplings, for example those set by the smallness of neutrino masses, are absent, very much like the case of the inverse seesaw model [7, 9].

Another most severe constraint on the couplings of the right-handed neutrinos with the charged leptons comes from their contribution to $\mu \to e\gamma$, which is trivially removed now. Since here only the N_{R_1} and N_{R_2} would mediate the process and they are degenerate, the $\mu \to e\gamma$ amplitude would depend on the effective light neutrino masses, which now vanishes, hence avoiding the constraint on the couplings.

Note also that the physical Higgs boson can significantly decay into $h \rightarrow \nu_L + N_{R_1}$ $h \rightarrow \nu_L + N_{R_2}$, leading to a mono-jet-like signal. In contrast, the decay $h \rightarrow \nu_L + N_{R_3}$ does not take place.

The third right-handed neutrino N_{R_3} can not decay at all because its only Yukawa coupling is with the singlet scalar σ . This means that N_{R_3} will contribute a sizeable relic density to the Universe. One can indeed check that N_{R_3} with the mass of a few GeV can have a desired cross section to serve as the dark matter for $\langle \sigma \rangle \sim 100$ GeV.

As a last comment we note that there may be a sizeable quartic interaction between the singlet scalar σ and the SM Higgs doublet ϕ , i.e. $\lambda(\sigma^{\dagger}\sigma)(\phi^{\dagger}\phi)$. This term can not be forbidden by imposing extra symmetries, and there is no *a priori* reason for λ to be small. In the presence of such coupling this dark matter N_{R_3} can be searched in the decays of the Higgs boson produced at the LHC, $h \rightarrow N_{R_3} + N_{R_3}$, resulting in a missing momentum signal.

- See talks given at the Neutrino2008 conference Web-page http://www2.phys.canterbury.ac.nz~jaa53/
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IV. SUMMARY

Due to some special symmetry it may happen that the type-I seesaw mechanism does not generate the observed neutrino masses, despite the co-existence of sizeable Dirac mass terms and relatively low right-handed neutrino Majorana masses. Such null seesaw mechanism can be understood as a cancelation between the contributions from the right-handed neutrinos. In this case there is only a very mild constraint on the right-handed neutrinos, which can have sizeable Yukawa couplings to the SM particles even if they are light. The leading Higgs boson decay mode into a left-handed neutrino and a righthanded neutrino could be probed at the LHC. In the model we have presented, one of the right-handed neutrinos has no Yukawa couplings to the SM states, a fact that follows from our assumed symmetry. Hence it can provide the relic density required to solve the puzzle of the dark matter. Finally, in our model the source of neutrino masses is the type-II seesaw contribution arising from the induced *vevs* of scalar Higgs triplets. The CP-violating and out-of-equilibrium decays of these scalar triplets may also account for the matter-antimatter asymmetry of the Universe through the leptogenesis mechanism.

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