Research Papers in Physics and Astronomy

Gregory Snow Publications

University of Nebraska - Lincoln

Year 2001

Search for New Physics Using QUAERO: A General Interface to D0 Event Data

V. M. Abazov^{*} Gregory $\operatorname{Snow}^{\dagger}$

D0 Collaboration[‡]

 * Joint Institute for Nuclear Research, Dubna, Russia

 † gsnow@unlhep.unl.edu ‡

This paper is posted at DigitalCommons@University of Nebraska - Lincoln. http://digitalcommons.unl.edu/physicssnow/55

Search for New Physics Using QUAERO: A General Interface to D0 Event Data

V.M. Abazov,²³ B. Abbott,⁵⁸ A. Abdesselam,¹¹ M. Abolins,⁵¹ V. Abramov,²⁶ B.S. Acharya,¹⁷ D.L. Adams,⁶⁰ M. Adams,³⁸ S. N. Ahmed,²¹ G. D. Alexeev,²³ G. A. Alves,² N. Amos,⁵⁰ E. W. Anderson,⁴³ Y. Arnoud,⁹ M. M. Baarmand,⁵⁵ V. V. Babintsev,²⁶ L. Babukhadia,⁵⁵ T. C. Bacon,²⁸ A. Baden,⁴⁷ B. Baldin,³⁷ P. W. Balm,²⁰ S. Banerjee,¹⁷ E. Barberis,³⁰ P. Baringer,⁴⁴ J. Barreto,² J. F. Bartlett,³⁷ U. Bassler,¹² D. Bauer,²⁸ A. Bean,⁴⁴ M. Begel,⁵⁴ A. Belyaev,³⁵ S. B. Beri,¹⁵ G. Bernardi,¹² I. Bertram,²⁷ A. Besson,⁹ R. Beuselinck,²⁸ V. A. Bezzubov,²⁶ P. C. Bhat,³⁷ V. Bhatnagar,¹¹ M. Bhattacharjee,⁵⁵ G. Blazey,³⁹ S. Blessing,³⁵ A. Boehnlein,³⁷ N. I. Bojko,²⁶ F. Borcherding,³⁷ K. Bos,²⁰ A. Brandt,⁶⁰ R. Breedon,³¹ G. Briskin,⁵⁹ R. Brock,⁵¹ G. Brooijmans,³⁷ A. Bross,³⁷ D. Buchholz,⁴⁰ M. Buehler,³⁸ V. Buescher,¹⁴ V. S. Burtovoi,²⁶ J. M. Butler,⁴⁸ F. Canelli,⁵⁴ W. Carvalho,³ D. Casey,⁵¹ Z. Casilum,⁵⁵ H. Castilla-Valdez,¹⁹ D. Chakraborty,³⁹ K. M. Chan,⁵⁴ S. V. Chekulaev,²⁶ D. K. Cho,⁵⁴ S. Choi,³⁴ S. Chopra,⁵⁶ J. H. Christenson,³⁷ M. Chung,³⁸ D. Claes,⁵² A. R. Clark,³⁰ J. Cochran,³⁴ L. Coney,⁴² B. Connolly,³⁵ W.E. Cooper,³⁷ D. Coppage,⁴⁴ S. Crépé-Renaudin,⁹ M. A. C. Cummings,³⁹ D. Cutts,⁵⁹ G. A. Davis,⁵⁴ K. Davis,²⁹ K. De,⁶⁰ S. J. de Jong,²¹ K. Del Signore,⁵⁰ M. Demarteau,³⁷ R. Demina,⁴⁵ P. Demine,⁹ D. Denisov,³⁷ S. P. Denisov,²⁶ S. Desai,⁵⁵ H. T. Diehl,³⁷ M. Diesburg,³⁷ G. Di Loreto,⁵¹ S. Doulas,⁴⁹ P. Draper,⁶⁰ Y. Ducros,¹³ L. V. Dudko,²⁵ S. Duensing,²¹ L. Duflot,¹¹ S. R. Dugad,¹⁷ A. Duperrin,¹⁰ A. Dyshkant,³⁹ D. Edmunds,⁵¹ J. Ellison,³⁴ V. D. Elvira,³⁷ R. Engelmann,⁵⁵ S. Eno,⁴⁷ G. Eppley,⁶² P. Ermolov,²⁵ O. V. Eroshin,²⁶ J. Estrada,⁵⁴ H. Evans,⁵³ V. N. Evdokimov,²⁶ T. Fahland,³³ S. Feher,³⁷ D. Fein,²⁹ T. Ferbel,⁵⁴ F. Filthaut,²¹ H. E. Fisk,³⁷ Y. Fisyak,⁵⁶ E. Flattum,³⁷ F. Fleuret,³⁰ M. Fortner,³⁹ H. Fox,⁴⁰ K. C. Frame,⁵¹ S. Fu,⁵³ S. Fuess,³⁷ E. Gallas,³⁷ A. N. Galyaev,²⁶ M. Gao,⁵³ V. Gavrilov,²⁴ R. J. Genik II,²⁷ K. Genser,³⁷ C. E. Gerber,³⁸ Y. Gershtein,⁵⁹ R. Gilmartin,³⁵ G. Ginther,⁵⁴ B. Gómez,⁵ G. Gómez,⁴⁷ P.I. Goncharov,²⁶ J. L. González Solís,¹⁹ H. Gordon,⁵⁶ L. T. Goss,⁶¹ K. Gounder,³⁷ A. Goussiou,²⁸ N. Graf,⁵⁶ G. Graham,⁴⁷ P. D. Grannis,⁵⁵ J. A. Green,⁴³ H. Greenlee,³⁷ S. Grinstein,¹ L. Groer,⁵³ S. Grünendahl,³⁷ A. Gupta,¹⁷ S. N. Gurzhiev,²⁶ G. Gutierrez,³⁷ P. Gutierrez,⁵⁸ N. J. Hadley,⁴⁷ H. Haggerty,³⁷ S. Hagopian,³⁵ V. Hagopian,³⁵ R. E. Hall,³² P. Hanlet,⁴⁹ S. Hansen,³⁷ J. M. Hauptman,⁴³ C. Hays,⁵³ C. Hebert,⁴⁴ D. Hedin,³⁹ J. M. Heinmiller,³⁸ A. P. Heinson,³⁴ U. Heintz,⁴⁸ T. Heuring,³⁵ M. D. Hildreth,⁴² R. Hirosky,⁶³ J. D. Hobbs,⁵⁵ B. Hoeneisen,⁸ Y. Huang,⁵⁰ R. Illingworth,²⁸ A. S. Ito,³⁷ M. Jaffré,¹¹ S. Jain,¹⁷ R. Jesik,²⁸ K. Johns,²⁹ M. Johnson,³⁷ A. Jonckheere,³⁷ M. Jones,³⁶ H. Jöstlein,³⁷ A. Juste,³⁷ W. Kahl,⁴⁵ S. Kahn,⁵⁶ E. Kajfasz,¹⁰ A. M. Kalinin,²³ D. Karmanov,²⁵ D. Karmgard,⁴² Z. Ke,⁴ R. Kehoe,⁵¹ A. Khanov,⁴⁵ A. Kharchilava,⁴² S. K. Kim,¹⁸ B. Klima,³⁷ B. Knuteson,³⁰ W. Ko,³¹ J. M. Kohli,¹⁵ A. V. Kostritskiy,²⁶ J. Kotcher,⁵⁶ B. Kothari,⁵³ A. V. Kotwal,⁵³ A. V. Kozelov,²⁶ E. A. Kozlovsky,²⁶ J. Krane,⁴³ M. R. Krishnaswamy,¹⁷ P. Krivkova,⁶ S. Krzywdzinski,³⁷ M. Kubantsev,⁴⁵ S. Kuleshov,²⁴ Y. Kulik,⁵⁵ S. Kunori,⁴⁷ A. Kupco,⁷ V.E. Kuznetsov,³⁴ G. Landsberg,⁵⁹ W. M. Lee,³⁵ A. Leflat,²⁵ C. Leggett,³⁰ F. Lehner,^{37,*} J. Li,⁶⁰ Q. Z. Li,³⁷ X. Li,⁴ J. G. R. Lima,³ D. Lincoln,³⁷ S. L. Linn,³⁵ J. Linnemann,⁵¹ R. Lipton,³⁷ A. Lucotte,⁹ L. Lueking,³⁷ C. Lundstedt,⁵² C. Luo,⁴¹ A. K. A. Maciel,³⁹ R. J. Madaras,³⁰ V. L. Malyshev,²³ V. Manankov,²⁵ H. S. Mao,⁴ T. Marshall,⁴¹ M. I. Martin,³⁹ R. D. Martin,³⁸ K. M. Mauritz,⁴³ B. May,⁴⁰ A. A. Mayorov,⁴¹ R. McCarthy,⁵⁵ T. McMahon,⁵⁷ H. L. Melanson,³⁷ M. Merkin,²⁵ K. W. Merritt,³⁷ C. Miao,⁵⁹ H. Miettinen,⁶² D. Mihalcea,³⁹ C. S. Mishra,³⁷ N. Mokhov,³⁷ N. K. Mondal,¹⁷ H. E. Montgomery,³⁷ R. W. Moore,⁵¹ M. Mostafa,¹ H. da Motta,² E. Nagy,¹⁰ F. Nang,²⁹ M. Narain,⁴⁸ V. S. Narasimham,¹⁷ H. A. Neal,⁵⁰ J. P. Negret,⁵ S. Negroni,¹⁰ T. Nunnemann,³⁷ D. O'Neil,⁵¹ V. Oguri,³ B. Olivier,¹² N. Oshima,³⁷ P. Padley,⁶² L. J. Pan,⁴⁰ K. Papageorgiou,³⁸ A. Para,³⁷ N. Parashar,⁴⁹ R. Partridge,⁵⁹ N. Parua,⁵⁵ M. Paterno,⁵⁴ A. Patwa,⁵⁵ B. Pawlik,²² J. Perkins,⁶⁰ M. Peters,³⁶ O. Peters,²⁰ P. Pétroff,¹¹ R. Piegaia,¹ B. G. Pope,⁵¹ E. Popkov,⁴⁸ H.B. Prosper,³⁵ S. Protopopescu,⁵⁶ J. Qian,⁵⁰ R. Raja,³⁷ S. Rajagopalan,⁵⁶ E. Ramberg,³⁷ P. A. Rapidis,³⁷ N. W. Reay,⁴⁵ S. Reucroft,⁴⁹ M. Ridel,¹¹ M. Rijssenbeek,⁵⁵ F. Rizatdinova,⁴⁵ T. Rockwell,⁵¹ M. Roco,³⁷ P. Rubinov,³⁷ R. Ruchti,⁴² J. Rutherfoord,²⁹ B.M. Sabirov,²³ G. Sajot,⁹ A. Santoro,² L. Sawyer,⁴⁶ R.D. Schamberger,⁵⁵ H. Schellman,⁴⁰ A. Schwartzman,¹ N. Sen,⁶² E. Shabalina,³⁸ R. K. Shivpuri,¹⁶ D. Shpakov,⁴⁹ M. Shupe,²⁹ R. A. Sidwell,⁴⁵ V. Simak,⁷ H. Singh,³⁴ J. B. Singh,¹⁵ V. Sirotenko,³⁷ P. Slattery,⁵⁴ E. Smith,⁵⁸ R. P. Smith,³⁷ R. Snihur,⁴⁰ G. R. Snow,⁵² J. Snow,⁵⁷ S. Snyder,⁵⁶ J. Solomon,³⁸ V. Sorín,¹ M. Sosebee,⁶⁰ N. Sotnikova,²⁵ K. Soustruznik,⁶ M. Souza,² N. R. Stanton,⁴⁵ G. Steinbrück,⁵³ R. W. Stephens,⁶⁰ F. Stichelbaut,⁵⁶ D. Stoker,³³ V. Stolin,²⁴ A. Stone,⁴⁶ D. A. Stoyanova,²⁶ M. Strauss,⁵⁸ M. Strovink,³⁰ L. Stutte,³⁷ A. Sznajder,³ M. Talby,¹⁰ W. Taylor,⁵⁵ S. Tentindo-Repond,³⁵ S. M. Tripathi,³¹ T. G. Trippe,³⁰ A. S. Turcot,⁵⁶ P. M. Tuts,⁵³ P. van Gemmeren,³⁷ V. Vaniev,²⁶ R. Van Kooten,⁴¹ N. Varelas,³⁸ L. S. Vertogradov,²³ F. Villeneuve-Seguier,¹⁰ A. A. Volkov,²⁶ A. P. Vorobiev,²⁶ H. D. Wahl,³⁵ H. Wang,⁴⁰ Z. -M. Wang,⁵⁵ J. Warchol,⁴² G. Watts,⁶⁴ M. Wayne,⁴² H. Weerts,⁵¹ A. White,⁶⁰ J. T. White,⁶¹ D. Whiteson,³⁰ J. A. Wightman,⁴³ D. A. Wijngaarden,²¹ S. Willis,³⁹ S. J. Wimpenny,³⁴ J. Womersley,³⁷ D. R. Wood,⁴⁹ R. Yamada,³⁷

P. Yamin,⁵⁶ T. Yasuda,³⁷ Y. A. Yatsunenko,²³ K. Yip,⁵⁶ S. Youssef,³⁵ J. Yu,³⁷ Z. Yu,⁴⁰ M. Zanabria,⁵ H. Zheng,⁴² Z. Zhou,⁴³ M. Zielinski,⁵⁴ D. Zieminska,⁴¹ A. Zieminski,⁴¹ V. Zutshi,⁵⁶ E. G. Zverev,²⁵ and A. Zylberstejn¹³

(D0 Collaboration)

¹Universidad de Buenos Aires, Buenos Aires, Argentina ²LAFEX, Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, Brazil ³Universidade do Estado do Rio de Janeiro, Rio de Janeiro, Brazil ⁴Institute of High Energy Physics, Beijing, People's Republic of China ⁵Universidad de los Andes, Bogotá, Colombia ⁶Charles University, Center for Particle Physics, Prague, Czech Republic ⁷Institute of Physics, Academy of Sciences, Center for Particle Physics, Prague, Czech Republic ⁸Universidad San Francisco de Quito, Quito, Ecuador ⁹Institut des Sciences Nucléaires, IN2P3-CNRS, Universite de Grenoble 1, Grenoble, France ¹⁰CPPM, IN2P3-CNRS, Université de la Méditerranée, Marseille, France ¹¹Laboratoire de l'Accélérateur Linéaire, IN2P3-CNRS, Orsay, France ¹²LPNHE, Universités Paris VI and VII, IN2P3-CNRS, Paris, France ¹³DAPNIA/Service de Physique des Particules, CEA, Saclay, France ¹⁴Universität Mainz, Institut für Physik, Mainz, Germany ¹⁵Panjab University, Chandigarh, India ¹⁶Delhi University, Delhi, India ¹⁷Tata Institute of Fundamental Research, Mumbai, India ¹⁸Seoul National University, Seoul, Korea ¹⁹CINVESTAV, Mexico City, Mexico ²⁰FOM-Institute NIKHEF and University of Amsterdam/NIKHEF, Amsterdam, The Netherlands ²¹University of Nijmegen/NIKHEF, Nijmegen, The Netherlands ²²Institute of Nuclear Physics, Kraków, Poland ²³Joint Institute for Nuclear Research, Dubna, Russia ²⁴Institute for Theoretical and Experimental Physics, Moscow, Russia ²⁵Moscow State University, Moscow, Russia ²⁶Institute for High Energy Physics, Protvino, Russia ²⁷Lancaster University, Lancaster, United Kingdom ²⁸Imperial College, London, United Kingdom ²⁹University of Arizona, Tucson, Arizona 85721 ³⁰Lawrence Berkeley National Laboratory and University of California, Berkeley, California 94720 ³¹University of California, Davis, California 95616 ³²California State University, Fresno, California 93740 ³³University of California, Irvine, California 92697 ³⁴University of California, Riverside, California 92521 ³⁵Florida State University, Tallahassee, Florida 32306 ³⁶University of Hawaii, Honolulu, Hawaii 96822 ³⁷Fermi National Accelerator Laboratory, Batavia, Illinois 60510 ³⁸University of Illinois at Chicago, Chicago, Illinois 60607 ³⁹Northern Illinois University, DeKalb, Illinois 60115 ⁴⁰Northwestern University, Evanston, Illinois 60208 ⁴¹Indiana University, Bloomington, Indiana 47405 ⁴²University of Notre Dame, Notre Dame, Indiana 46556 ⁴³Iowa State University, Ames, Iowa 50011 ⁴⁴University of Kansas, Lawrence, Kansas 66045 ⁴⁵Kansas State University, Manhattan, Kansas 66506 ⁴⁶Louisiana Tech University, Ruston, Louisiana 71272 ⁴⁷University of Maryland, College Park, Maryland 20742 ⁴⁸Boston University, Boston, Massachusetts 02215 ⁴⁹Northeastern University, Boston, Massachusetts 02115 ⁵⁰University of Michigan, Ann Arbor, Michigan 48109 ⁵¹Michigan State University, East Lansing, Michigan 48824 ⁵²University of Nebraska, Lincoln, Nebraska 68588 ⁵³Columbia University, New York, New York 10027 ⁵⁴University of Rochester, Rochester, New York 14627 ⁵⁵State University of New York, Stony Brook, New York 11794 ⁵⁶Brookhaven National Laboratory, Upton, New York 11973 ⁵⁷Langston University, Langston, Oklahoma 73050

⁵⁸University of Oklahoma, Norman, Oklahoma 73019
 ⁵⁹Brown University, Providence, Rhode Island 02912
 ⁶⁰University of Texas, Arlington, Texas 76019
 ⁶¹Texas A&M University, College Station, Texas 77843
 ⁶²Rice University, Houston, Texas 77005
 ⁶³University of Virginia, Charlottesville, Virginia 22901
 ⁶⁴University of Washington, Seattle, Washington 98195
 (Received 7 June 2001; published 14 November 2001)

We describe QUAERO, a method that (i) enables the automatic optimization of searches for physics beyond the standard model, and (ii) provides a mechanism for making high energy collider data generally available. We apply QUAERO to searches for standard model WW, ZZ, and $t\bar{t}$ production, to searches for these objects produced through a new heavy resonance, and to the first direct search for $W' \rightarrow WZ$. Through this interface, we make three data sets collected by the D0 experiment at $\sqrt{s} = 1.8$ TeV publicly available.

DOI: 10.1103/PhysRevLett.87.231801

PACS numbers: 13.85.Rm, 12.60.-i, 29.85.+c

It is generally recognized that the standard model, a successful description of the fundamental particles and their interactions, must be incomplete. Models that extend the standard model often predict rich phenomenology at the scale of a few hundred GeV, an energy regime accessible to the Fermilab Tevatron. In part because of the complexity of the apparatus required to test models at such large energies, experimental responses to these ideas have not kept pace. Any technique that reduces the time required to test a particular candidate theory would allow more such theories to be tested, reducing the possibility that the data contain overlooked evidence for new physics.

Once data are collected and the backgrounds have been understood, the testing of any specific model in principle follows a well-defined procedure. In practice, this process has been far from automatic. Even when the basic selection criteria and background estimates are taken from a previous analysis, the reinterpretation of the data in the context of a new model often requires a substantial length of time.

Ideally, the data should be "published" in such a way that others in the community can easily use those data to test a variety of models. The publishing of experimental distributions in journals allows this to occur at some level, but an effective publishing of a multidimensional data set has, to our knowledge, not yet been accomplished by a large particle physics experiment. The problem appears to be that such data are context specific, requiring detailed knowledge of the complexities of the apparatus. This knowledge must somehow be incorporated either into the data or into whatever tool the nonexpert would use to analyze those data.

Many data samples and backgrounds have been defined in the context of SLEUTH [1], a quasi-model-independent search strategy for new high p_T physics that has been applied to a number of exclusive final states [2,3] in the data collected by the D0 detector [4] during 1992-1996 in Run I of the Fermilab Tevatron. In this Letter, we describe a tool (QUAERO) that automatically optimizes an analysis for a particular signature, using these samples and standard model backgrounds. SLEUTH and QUAERO are complementary approaches to searches for new phenomena, enabling analyses that are both general (SLEUTH) and focused (QUAERO). We demonstrate the use of QUAERO in eleven separate searches: standard model WW and ZZ production; standard model $t\bar{t}$ production with leptonic and semileptonic decays; resonant WW, ZZ, WZ, and $t\bar{t}$ production; associated Higgs boson production; and pair production of first generation scalar leptoquarks. The data described here are accessible through QUAERO on the World Wide Web [5], for general use by the particle physics community.

The signals predicted by most theories of physics beyond the standard model involve an increased number of predicted events in some region of an appropriate variable space. In this case the optimization of the analysis can be understood as the selection of the region in this variable space that minimizes $\overline{\sigma^{95\%}}$, the expected 95% confidence level (C.L.) upper limit on the cross section of the signal in question, assuming the data contain no signal. The optimization algorithm consists of a few simple steps:

(i) Kernel density estimation [6] is used to estimate the probability distributions $p(\vec{x} | s)$ and $p(\vec{x} | b)$ for the signal and background samples in a low-dimensional variable space \mathcal{V} , where $\vec{x} \in \mathcal{V}$. The signal sample is contained in a Monte Carlo file provided as input to QUAERO. The background sample is constructed from all known standard model and instrumental sources.

(ii) A discriminant function $D(\vec{x})$ is defined by [7]

$$D(\vec{x}) = \frac{p(\vec{x} \mid s)}{p(\vec{x} \mid s) + p(\vec{x} \mid b)}.$$
 (1)

The semi-positive-definiteness of $p(\vec{x} | s)$ and $p(\vec{x} | b)$ restricts $D(\vec{x})$ to the interval [0, 1] for all \vec{x} .

(iii) The sensitivity S of a particular threshold $D_{\rm cut}$ on the discriminant function is defined as the reciprocal of $\overline{\sigma^{95\%}}$. $D_{\rm cut}$ is chosen to maximize S.

(iv) The region of variable space having $D(\vec{x}) > D_{\text{cut}}$ is used to determine the actual 95% C.L. cross section upper limit $\sigma^{95\%}$ [8].

TABLE I. A summary of the data available within QUAERO, including the selection cuts applied and the efficiency of identification requirements. The final states are inclusive, with many events containing one or more additional jets. Reconstructed jets satisfy $p_T^j > 15$ GeV and $|\eta_{det}^j| < 2.5$, and reconstructed electrons satisfy $p_T^e > 15$ GeV and $|\eta_{det}^e| < 1.1$ or $1.5 < |\eta_{det}^e| < 2.5$, where η_{det} is the pseudorapidity measured from the center of the detector.

Final state	Selection criteria	$\epsilon_{ ext{ID}}$	$\int \mathcal{L} dt$
eμ	$p_T^{e,\mu} > 15 \text{ GeV} \ \eta_{ m det}^{\mu} < 1.7$	0.30	$108 \pm 5 \text{ pb}^{-1}$
$e \not\!\!\!E_T 2j$	$p_T^{e,j_{1,2}} > 20 \text{ GeV}$	0.61	$115 \pm 6 \text{ pb}^{-1}$
	$\not\!$		
	$p_T^{e\not\!\!\!E_T}>40~{ m GeV}$		
ee2j	$p_T^{e_{1,2},j_{1,2}} > 20 \text{ GeV}$	0.70	$123 \pm 7 \text{ pb}^{-1}$

When provided with a signal model and a choice of variables \mathcal{V} , QUAERO uses this algorithm and D0 Run I data to compute an upper limit on the cross section of the signal. Instructions for use are available from the QUAERO web site.

Table I shows the data available within QUAERO, and Table II summarizes the backgrounds. These data and their backgrounds are described in more detail in Ref. [3]. The final states are inclusive, with many events containing one or more additional jets. Kolmogorov-Smirnov tests have been used to demonstrate agreement between data and the expected backgrounds in many distributions. The fraction of events with true final state objects satisfying the cuts shown that satisfy these cuts after reconstruction is given as an "identification" efficiency ($\epsilon_{\rm ID}$). Because electrons are more accurately measured and more efficiently identified than muons in the D0 detector, the corresponding muon channels $\mu \not \!\!\!\! E_T 2j$ and $\mu \mu 2j$ have been excluded from these data.

To check standard model results, we remove WW and ZZ production from the background estimate and search (i) for standard model WW production in the space defined by the transverse momentum of the electron (p_T^e) and missing transverse energy $(\not\!\!E_T)$ in the final state $e \mu \not\!\!E_T$, and (ii) for standard model ZZ production in the space defined

TABLE II. Standard model backgrounds (often produced with accompanying jets) to the final states considered. VV denotes WW, WZ, and ZZ; "data" indicates backgrounds from jets misidentified as electrons estimated using data. Monte Carlo programs (ISAJET [9], PYTHIA [10], HERWIG [11], and VECBOS [12]) are used to estimate several sources of background.

	Standard model backgrounds					
Final state	multijets	W	Ζ	VV	tĪ	
eμ	data	data	ISAJET	PYTHIA	HERWIG	
$e \not\!\!\!E_T 2j$	data	VECBOS		PYTHIA	HERWIG	
ee2j	data		PYTHIA	PYTHIA		

by the invariant mass of the two electrons (m_{ee}) and two jets (m_{jj}) in the final state ee2j. Removing $t\bar{t}$ production from the background estimate, we search for this process (iii) in the final state $e\not\!\!\!E_T 4j$ using the two variables laboratory aplanarity (A) and $\sum p_T^j$, and (iv) in the final state $e\mu\not\!\!\!E_T 2j$, using the two variables p_T^e and $\sum p_T^j$, assuming a top quark mass of 175 GeV.

Including all standard model processes in the background estimate, we look for evidence of new heavy resonances. We search (v) for resonant WW production in constraining $m_{e\nu}$ and m_{jj} to M_W , and (vi) for resonant ZZ production in the final state ee2j, using the variable m_{eejj} after constraining m_{ii} to M_Z . In both cases we remove events that cannot be so constrained. To obtain a specific signal prediction, we assume that the resonance behaves like a standard model Higgs boson in its couplings to the W and Z bosons. Constraining $m_{e\nu}$ to M_W and m_{ii} to M_Z , we use the quality of the fit and $m_{e\nu jj}$ to search (vii) for a massive W' boson in the extended gauge model of Ref. [13]. Using $m_{e\nu 4i}$ after constraining $m_{e\nu}$ to M_W , we search (viii) for a massive narrow Z' resonance with Z-like couplings decaying to $t\bar{t} \rightarrow W^+ bW^- \bar{b} \rightarrow e\nu 4j$.

Nonresonant new phenomena are also considered. The variables m_{jj} and either m_{ev}^T or m_{ee} are used to search for a light Higgs boson produced (ix) in association with a W boson, and (x) in association with a Z boson. Finally,

TABLE III. Limits on cross section \times branching fraction for the processes discussed in the text. All final states are inclusive in the number of additional jets. The fraction of the signal sample satisfying QUAERO's selection criteria is denoted ϵ_{sig} ; \hat{b} is the number of expected background events satisfying these criteria; and N_{data} is the number of events in the data satisfying these criteria. The subscripts on h, W', Z', and LQ denote assumed masses, in units of GeV.

Process	$\boldsymbol{\epsilon}_{\mathrm{sig}}$	\hat{b}	N _{data}	$\sigma^{95\%} imes \mathcal{B}$
$WW \rightarrow e \mu \not\!\!\!E_T$	0.14	19.0 ± 4.0	23	1.1 pb
$ZZ \rightarrow ee2j$	0.12	19.7 ± 4.1	19	0.8 pb
$t\bar{t} \rightarrow e \not\!$	0.13	3.1 ± 0.9	8	0.8 pb
$t\bar{t} \to e \mu \not\!$	0.14	0.6 ± 0.2	2	0.4 pb
$h_{175} \rightarrow WW \rightarrow e \not\!\!\!E_T 2j$	0.02	29.6 ± 6.5	32	11.0 pb
$h_{200} \rightarrow WW \rightarrow e \not\!\!E_T 2j$	0.07	66.0 ± 13.8	69	4.4 pb
$h_{225} \rightarrow WW \rightarrow e \not\!\!\!E_T 2j$	0.06	43.1 ± 9.2	44	3.6 pb
$h_{200} \rightarrow ZZ \rightarrow ee2j$	0.15	17.9 ± 3.7	15	0.6 pb
$h_{225} \rightarrow ZZ \rightarrow ee2j$	0.15	18.8 ± 3.8	12	0.4 pb
$h_{250} \rightarrow ZZ \rightarrow ee2j$	0.17	18.1 ± 3.7	18	0.6 pb
$W'_{200} \rightarrow WZ \rightarrow e \not\!\!\!E_T 2j$	0.05	27.7 ± 6.3	29	3.4 pb
$W'_{350} \rightarrow WZ \rightarrow e \not\!\!\!E_T 2j$	0.23	22.7 ± 5.2	27	0.7 pb
$W'_{500} \rightarrow WZ \rightarrow e \not\!\!\!E_T 2j$	0.26	2.1 ± 0.8	2	0.2 pb
$Z'_{350} \rightarrow t\bar{t} \rightarrow e \not\!\!\!E_T 4j$	0.11	18.7 ± 4.0	20	1.1 pb
$Z'_{450} \rightarrow t\bar{t} \rightarrow e \not\!\!E_T 4j$	0.14	18.7 ± 4.0	20	0.9 pb
$Z'_{550} \rightarrow t\bar{t} \rightarrow e \not\!\!\!E_T 4j$	0.14	3.8 ± 1.0	2	0.3 pb
$Wh_{115} \rightarrow e \not\!\!E_T 2j$	0.08	37.3 ± 8.2	32	2.0 pb
$Zh_{115} \rightarrow ee2j$	0.20	19.5 ± 4.1	25	0.8 pb
$LQ_{225}\overline{LQ}_{225} \rightarrow ee2j$	0.33	0.3 ± 0.1	0	0.07 pb





we search (xi) for first generation scalar leptoquarks with mass 225 GeV in the final state ee2j using m_{ee} and S_T , the summed scalar transverse momentum of all electrons and jets in the event. The numerical results of these searches are listed in Table III. Figures 1 and 2 present plots of the signal density, background density, and selected region in the variables considered.



FIG. 2. QUAERO's analysis of signatures involving undiscovered particles. From top to bottom the hypothetical signals are $h_{200} \rightarrow ZZ \rightarrow ee2j$, $Z'_{550} \rightarrow t\bar{t} \rightarrow eE_T4j$, $Wh_{115} \rightarrow eE_T2j$, and $LQ_{225}\overline{LQ}_{225} \rightarrow ee2j$. Plots (c) of the first two rows show the discriminant D (curve), the threshold D_{cut} (horizontal line), and the data (histogram); the region with $D > D_{cut}$ is selected.

check of the method, QUAERO almost exactly duplicates a previous search for $LQ\overline{LQ} \rightarrow ee2j$ [15].

QUAERO is a method both for automatically optimizing searches for new physics and for allowing D0 to make a subset of its data available for general use. In this Letter, we have outlined the algorithm used in QUAERO, and we have described the final states currently available for analysis using this method. QUAERO's performance on several examples, including both standard model and resonant WW, ZZ, and $t\bar{t}$ production, has been demonstrated. The limits obtained are comparable to those from previous searches at hadron colliders. The searches for $ZZ \rightarrow ee2j$, $Z' \rightarrow t\bar{t} \rightarrow e \not E_T 4j$, $Wh \rightarrow e \not E_T 2j$, and $Zh \rightarrow ee2j$ are the first from D0, and the searches for $W' \rightarrow WZ$ and resonant WW and ZZ production are the first of their kind. This tool should increase the facility with which new models may be tested in the future.

We thank the staffs at Fermilab and collaborating institutions, and acknowledge support from the Department of Energy and National Science Foundation (U.S.A.), Commissariat á L'Energie Atomique and CNRS/Institut National de Physique Nucléaire et de Physique des Particules (France), Ministry for Science and Technology and Ministry for Atomic Energy (Russia), CAPES and CNPq (Brazil), Departments of Atomic Energy and Science and Education (India), Colciencias (Colombia), CONACyT (Mexico), Ministry of Education and KOSEF (Korea), CONICET and UBACyT (Argentina), The Foundation for Fundamental Research on Matter (The Netherlands), PPARC (United Kingdom), Ministry of Education (Czech Republic), and the A. P. Sloan Foundation.

*Visitor from University of Zurich, Zurich, Switzerland.

- [1] D0 Collaboration, B. Abbott *et al.*, Phys. Rev. D **62**, 092004 (2000).
- [2] D0 Collaboration, B. Abbott *et al.*, Phys. Rev. Lett. 86, 3712 (2001).
- [3] D0 Collaboration, B. Abbott *et al.*, Phys. Rev. D **64**, 012004 (2001).

- [4] D0 Collaboration, S. Abachi *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **338**, 185 (1994).
- [5] D0 Collaboration, B. Abbott et al., http://quaero.fnal.gov/
- [6] David Scott, *Multivariate Density Estimation* (John Wiley & Sons, New York, 1992).
- [7] L. Holmström, S. Sain, and H. Miettinen, Comput. Phys. Commun. 88, 195 (1995).
- [8] $\sigma^{95\%}$ is a Bayesian limit, computed assuming a flat prior.
- [9] H. Baer *et al.*, BNL-HET-99-43; FSU-HEP-991218; UH-511-952-00; hep-ph/0001086; we used v7.22.
- [10] T. Sjöstrand *et al.*, Comput. Phys. Commun. **135**, 238 (2001).
- [11] G. Corcella *et al.*, J. High Energy Phys. **1** 10 (2001); hep-ph/0107071; we used v5.7.
- [12] F. A. Berends *et al.*, Nucl. Phys. **B357**, 32 (1991); we used v3.0.
- [13] G. Altarelli, B. Mele, and M. Ruiz-Altaba, Z. Phys. C 45, 109 (1989).
- [14] D0 Collaboration, S. Abachi *et al.*, Phys. Rev. Lett. **79**, 1203 (1997).
- [15] D0 Collaboration, B. Abbott *et al.*, Phys. Rev. Lett. **79**, 4321 (1997).