

Design, performance, and calibration of the CMS hadron-outer calorimeter

CMS HCAL Collaborations

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Abstract The Outer Hadron Calorimeter (HCAL HO) of the CMS detector is designed to measure the energy that is not contained by the barrel (HCAL HB) and electromagnetic (ECAL EB) calorimeters. Due to space limitation the barrel calorimeters do not contain completely the hadronic shower and an outer calorimeter (HO) was designed, constructed and inserted in the muon system of CMS to measure the energy leakage. Testing and calibration of the HO was carried out in a 300 GeV/c test beam that improved the linearity and resolution. HO will provide a net improvement in missing E_T measurements at LHC energies. Information from HO will also be used for the muon trigger in CMS.

1 Introduction

The Compact Muon Solenoid (CMS) experiment [1] at the Large Hadron Collider (LHC) at CERN is a multi-purpose detector optimised to look for signatures of the Higgs boson, super-symmetric particles and other new physics processes in pp collisions at a centre of mass energy of 14 TeV. Signature for production of many of these new particles is large missing transverse energy, E_T . The momenta of all charged particles are measured using a high resolution tracker in a 4 T solenoidal magnetic field, while energies of all particles, charged as well as photons and neutral hadrons, are measured using electromagnetic and hadron calorimeters. Key elements in improving the missing transverse energy measurement are hermiticity of the calorimeter, excellent energy resolution and small energy leakage.

Figure 1 shows the longitudinal quarter view of the CMS detector. The hadron calorimeter [2], located behind the tracker and the electromagnetic calorimeter as seen from the interaction point, is a sampling calorimeter made of copper alloy¹ absorber and plastic scintillators. The barrel hadron calorimeter (HB) has a polygonal structure. It is made out of two halves each 4.33 m long. A half barrel is made of 18 wedges each subtending 20° in azimuth ϕ . The HB has 16 scintillator layers designed to be hermetic with

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¹ Cartridge brass # 260 with 70% Copper and 30% Zinc.

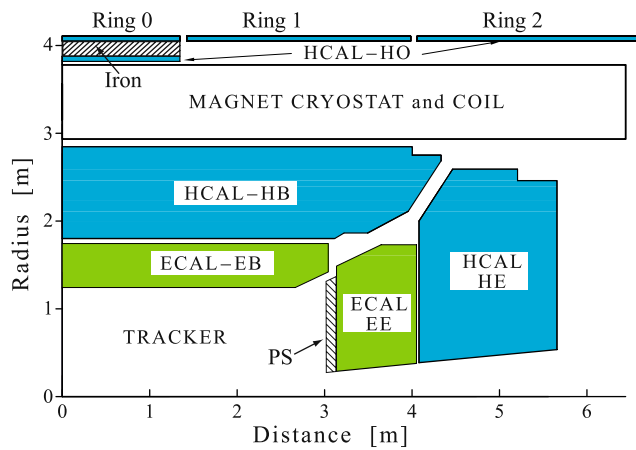


Fig. 1 Location of the hadron calorimeter in and around the CMS magnet. HCAL HB and HCAL HE are the hadron calorimeters. ECAL EB and ECAL EE are the electromagnetic calorimeters and PS is pre-shower detector

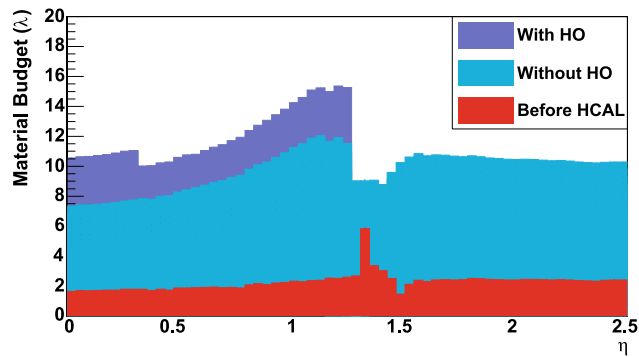


Fig. 2 Number of interaction lengths of the CMS calorimeter as a function of η . The two shaded regions correspond to the setups with or without the outer hadron calorimeter (HO)

minimal uninstrumented gaps. HB extends up to $|\eta| \approx 1.4$ ($\eta = -\ln \tan \frac{\theta}{2}$, where θ is the polar angle).

HB is radially restricted between the outer extent of the electromagnetic calorimeter ($R = 1.77$ m) and the magnet cryostat and coil ($R = 2.95$ m). This constrains the total amount of material for the calorimeters EB plus HB to about 7.8 interaction lengths λ_1 at $\eta = 0$ as can be seen in Fig. 2.

Extensive Monte Carlo simulations and test beam studies were carried out [3] to define the parameters of the hadron calorimeter. Figure 3 shows the fraction of events that are missing more than one third of the energy for a test beam of 300 GeV/c for HB alone. Adding the tracker and EB, at 300 GeV/c nearly 3% of pion events gives rise to missing energy corresponding to 100 GeV. Small number of events with large leakage of energy lead to large fluctuation on an event by event basis and cannot be corrected offline. This unmeasured energy is the primary justification for the outer hadron (HO) calorimeter.

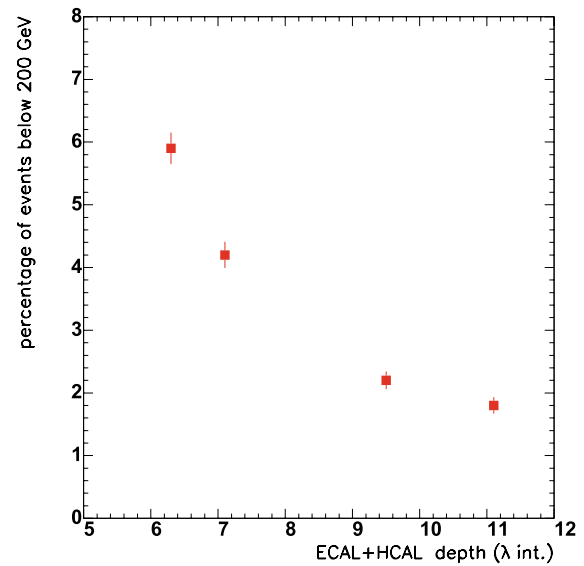


Fig. 3 Fraction of 300 GeV/c pion data with reconstructed energy less than 200 GeV (approximately 3σ below the mean, and corresponding to more than 100 GeV missing energy) versus total absorber depth obtained from 1996 test beam data at 3 T magnetic field. The total λ_{int} includes the contribution of the electromagnetic calorimeter

2 The outer hadron calorimeter (HO): constraints and expectation

The EB plus HB inside the 4 T cryostat and solenoid is relatively thin. To ensure adequate sampling depth for $\eta < 1.4$, the hadron calorimeter is extended outside the cryostat. The HO will utilise the cryostat and solenoid coil as an additional absorber ($1.4\lambda_{int}/\sin\theta$) and will also identify and measure the late starting showers.

Outside the cryostat the magnetic field is returned through an iron yoke designed in the form of five 2.536 m wide (along z -axis) rings. The HO is placed as the first sensitive layer in each of these five rings. The rings are identified by numbers $-2, -1, 0, +1$ and $+2$. The numbering increases with z and the nominal central z positions of the five rings are respectively -5.342 m, -2.686 m, 0 , $+2.686$ m and $+5.342$ m. Since the EB plus HB at the central ring (ring 0) has the smallest value of λ_1 , two layers of HO scintillators were placed on either side of a 19.5 cm thick piece of iron (the tail catcher iron) at radial distances of 3820 mm and 4070 mm, respectively. All other rings have a single HO layer at a radial distance of 4070 mm. The total depth of the calorimeter system is thus extended to a minimum of $10\lambda_{int}$ (see Fig. 2) except at the barrel-endcap boundary region.

The HO is constrained by the geometry and construction of the muon system. Figure 1 shows the position of the HO layers in the rings of the muon stations. The segmentation of these detectors closely follows that of the barrel muon system. Each ring has the 12 identical ϕ -sectors. The 12 sectors are separated by 75 mm thick stainless steel beams which

hold successive layers of iron of the return yoke. The space between successive muon rings in the η direction and also the space occupied by the stainless steel beams in the ϕ direction cause some reduction in the active area of the HO. In addition, the space occupied by the CMS ‘chimneys’ in the vertical sector of ring -1 and $+1$ are also not available for the HO as well. The chimneys are used for the cryogenic transfer lines and power cables of the magnet system. Finally, the mechanical structures needed to position the scintillator trays further constrains HO along ϕ .

In the radial direction, space for HO is limited to only 16 mm in the iron yoke, as there are aluminium honey comb support structures for the muon detector system. In addition, the HO modules are independently supported from the steel beams located on either side of each ϕ sector. The thickness and position of the iron ribs in the yoke structure further constrains the shape and segmentation of the HO.

The sizes and positions of the HO 1.0 cm thick scintillator tiles (BC408) roughly map onto the HB towers of granularity 0.087×0.087 in η and ϕ . The HO consists of one (rings ± 1 and ± 2) or two (ring 0) layers of scintillator tiles located in front of the first layer of the barrel muon detector. Scintillation light from the tiles is collected using multi-clad Y11 Kuraray wave-length shifting (WLS) fibres, of diameter 0.94 mm, and transported to the photo detectors located on the structure of the return yoke by splicing a multi-clad Kuraray clear fibre (also of 0.94 mm diameter) with the WLS fibre. In order to simplify installation of the HO, the scintillator tiles are packed into a single unit called a tray. Each tray corresponds to one ϕ slice ($\sim 5^\circ$ wide in ϕ). However, along the $z(\eta)$ direction, a tray covers the entire span of a muon ring. Figure 4 shows a schematic view of a HO tray where one tile is mapped to a tower of the HB and the optical cable from the tray is connected to the readout box.

The physics impact of HO has been studied [4] using the simulation tool CMSIM developed for the CMS detector. The simulation was carried out for single pions at specific η values. The deposited energy in EB plus HB show a deficit in energy measurement starting at 70 GeV/c at $\eta = 0$ (ring 0). The results with HO are more Gaussian in energy distribution indicating that the addition of the HO properly recovers the energy leakage. Even with HO in ring 0 there is some small amount of energy leakage. There is evidence of leakage beyond HO in ring 1 but with reduced intensity. The amount of leakage in ring 2 is found to be negligible.

A critical signature of several new physics signals is the evidence of unobserved particle(s), whether it is supersymmetry, or leptoquark or just plain production of dark matter. CMS detects unobservable particle(s) by measuring missing transverse energy E_T and will use it as part of the trigger and data analysis. Any energy that escapes observation will increase the background making data analysis harder. HO has been designed to reduce the unobserved energy background.

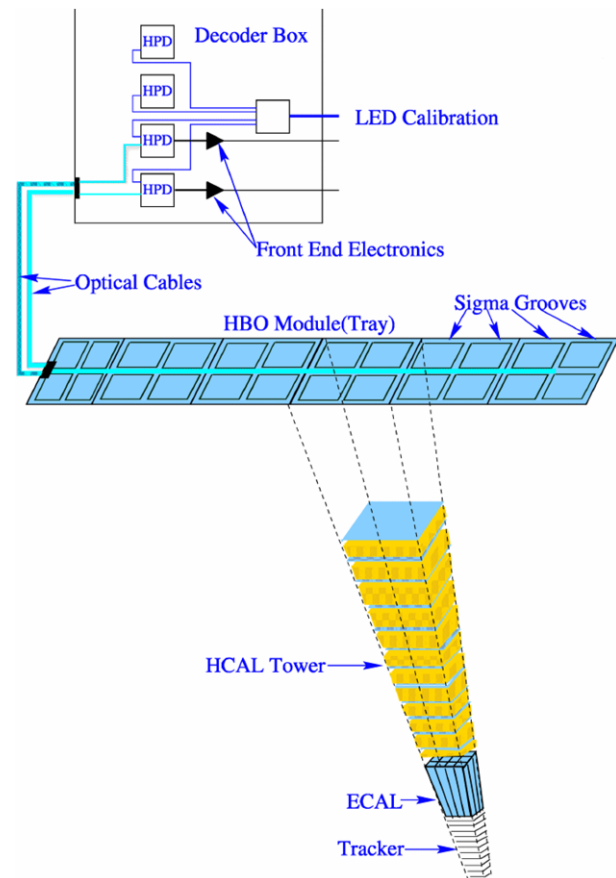


Fig. 4 Schematic view of a HO tray shown with individual tiles and the corresponding grooves for WLS fibres. Each tile is mapped to a tower of HB. Optical fibres from the tray extend to the decoder box which contains the photo-detector and readout electronics

QCD events also have missing E_T due to the production of neutrinos. The cross section of QCD events, where at least one particle has E_T above 500 GeV, is estimated to be several pb. Clearly these events will be affected due to leakage of energy in the hadron calorimeter, and the HO would help to decrease the background and improve the energy measurement. Figure 5 shows the accepted integrated cross section due to QCD processes for missing E_T above a certain value measured with and without HO. It is clear that the inclusion of HO reduces the background cross section by a factor of 1.5 or more for moderate E_T values.

3 HO module: design and assembly

3.1 Mechanical design

HO [5] is physically divided into 5 rings in η conforming to the muon ring structure. The rings are numbered -2 , -1 , 0 , $+1$ and $+2$ with increasing η . Each ring of the HO is divided into 12 identical ϕ sectors (numbered 1 to 12 counting

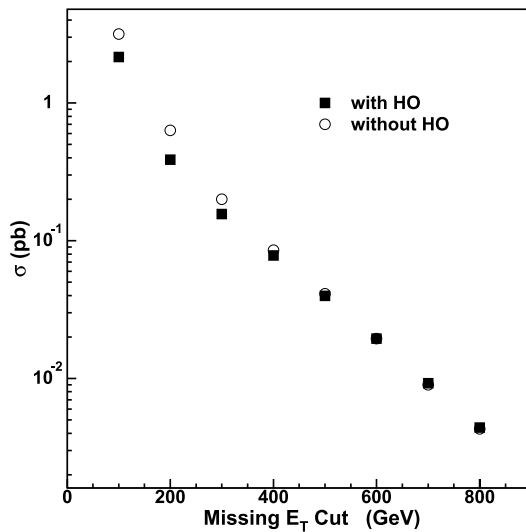


Fig. 5 Accepted integrated cross section above threshold as a function of missing E_T when the missing E_T is measured with and without the HO

Table 1 Dimension of tiles along ϕ for different trays. Each tray corresponds to one ϕ -slice in a ϕ sector

Tray #	Ring 0		Ring $\pm 1, \pm 2$ Layer 1
	Layer 0	Layer 1	
Tray 1	272.6 mm	298.6 mm	315.6 mm
Tray 2	341.6 mm	362.6 mm	364.6 mm
Tray 3	330.6 mm	350.6 mm	352.6 mm
Tray 4	325.6 mm	345.6 mm	347.6 mm
Tray 5	325.6 mm	345.6 mm	347.6 mm
Tray 6	266.6 mm	290.6 mm	404.6 mm

clockwise starting from 9 o'clock position) and each sector has 6 slices (numbered 1 to 6 counting clockwise) in ϕ . The ϕ slices of a layer are identical in all sectors. The widths of the slices along ϕ are given in Table 1. In each ϕ slice, there is a further division along η . The smallest scintillator unit is called a tile. The scintillator tiles in each ϕ sector belong to a plane. Perpendicular distance of this plane from the z -axis is 3.82 m for layer 0 and 4.07 m for layer 1. The tiles in each ϕ slice of a ring are mechanically held together in the form of a tray (details in Sect. 3.2).

Both layers of ring 0 have 8 η -divisions (i.e. 8 tiles in a tray): $-4, -3, -2, -1, +1, +2, +3, +4$. Ring 1 has 6 divisions: 5, ..., 10 and ring 2 has 5 divisions: 11, ..., 15. Ring -1 and ring -2 have the same number of divisions as rings 1 and 2 but with negative indices. The η -dimensions of any tile with negative tower number is the same as the one with positive number. Tile dimensions along η for the towers are shown in Table 2.

Length of a full tray is 2510 mm whereas the shorter trays, the sizes of which are constrained because of the

Table 2 HO tile dimensions along η for different rings and layers. The tile sizes, which are constrained by muon ring boundaries, are mentioned in parenthesis

Tower	η_{\max}	Length (mm)	Tower	η_{\max}	Length (mm)
Ring 0 Layer 0			Ring 0 Layer 1		
1	0.087	331.5	1	0.087	351.2
2	0.174	334.0	2	0.174	353.8
3	0.262	339.0	3	0.262	359.2
4	0.326	(248.8)	4	0.307	(189.1)
Ring 1 Layer 1			Ring 2 Layer 1		
5	0.436	391.5	11	0.960	420.1
6	0.524	394.2	12	1.047	545.1
7	0.611	411.0	13	1.135	583.3
8	0.698	430.9	14	1.222	626.0
9	0.785	454.0	15	1.262	(333.5)
10	0.861	(426.0)			

chimney (trays 4 and 5 in sector 4 of ring $+1$ and trays 3, 4, 5 and 6 in sector 3 of ring -1), are 2119 mm long. The shorter trays are constructed without the tile corresponding to tower number ± 5 . Because of the constraints imposed by the gap between the two rings, a part of tower # ± 4 , which falls in ring ± 1 (tower ± 4 is restricted only to ring 0) is merged with tower # ± 5 .

There is a maximum of 40 mm of clearance for placing the HO trays together with the mechanical structures. Each tray with all its packing is expected to have a thickness of 15.44 mm. This necessitated a space of $17.0^{+0.2}_{-0.0}$ mm for inserting the trays. The total weight of a ϕ sector (six trays) is about 30.0 kg. To support this weight, a pair of aluminium honeycomb panels (10 mm thick on the bottom and 6 mm thick at the top) are clamped between three aluminium C-channels on the two sides and one at the end while the other end is kept open for inserting the trays. A set of twelve 20 mm \times 20 mm aluminium C-channels are glued on the inner side at a fixed separation for inserting the trays. The separation is properly adjusted for the appropriate tray width. All the panels are 2530^{+5}_{-0} mm long. The widths of the three types of panels (layer 0 and 1 of ring 0 and layer 1 of rings $\pm 1, \pm 2$) are respectively 1920^{+2}_{-0} mm, 2050^{+2}_{-0} mm and 2190^{+2}_{-0} mm. Panels for rings ± 1 and ± 2 also carry cooling tubes on one side. These serve as thermal screens to prevent heat percolation to the RPC/muon chamber assembly. The cooling tubes are made of copper and have inner and outer diameter of 6 mm and 8 mm. These are glued to the 6 mm thick honeycomb surface using araldite and fixed with copper collars. For rings ± 1 , a special panel is built to accommodate the short trays due to the chimneys.

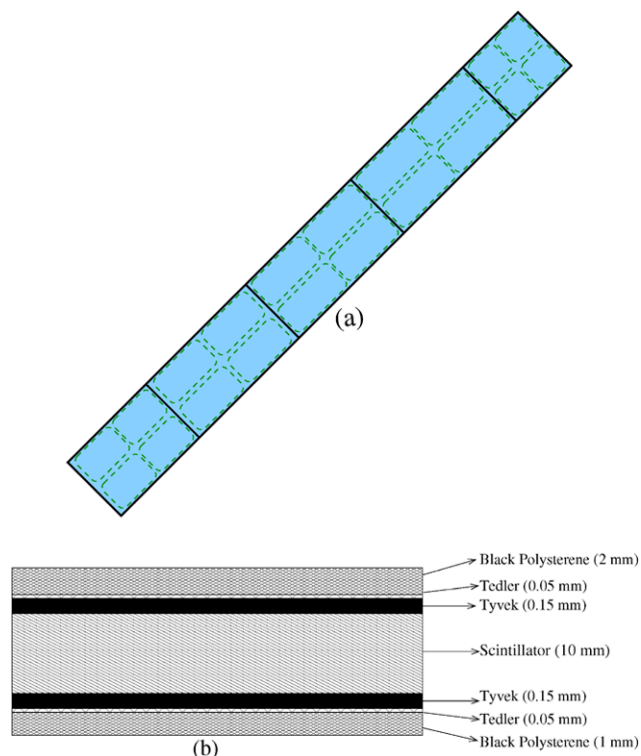


Fig. 6 (a) Layout of scintillator tiles in a typical tray of ring 2. (b) Cross section of a HO tray showing the different components

3.2 Trays

All tiles in each ϕ slice of a sector are grouped together in the form of a tray. Figure 6(a) shows a schematic view of a HO tray. Each tray contains 5 tiles in rings ± 2 ; 6 tiles in rings ± 1 and 8 tiles in ring 0. The edges of the tiles are painted with Bicorn reflecting white paint for better light collection as well as isolating the individual tiles of a tray. Further isolation of tiles is achieved by inserting a piece of black tedlar strip in between the adjacent tiles. The tiles in a tray are covered with a single big piece of white, reflective tyvek paper. Then they are covered with black tedlar paper to prevent light leak. This package is placed between two black plastic plates for mechanical stability and ease of handling. The top plastic cover is 2 mm thick and the bottom one is 1 mm thick. Figure 6(b) shows a cross section of a tray to illustrate the different components.

The 2 mm plastic sheet on the top has 1.6 mm deep channels grooved on it (on the outer side) to route the fibres from individual tiles to an optical connector placed in a groove at the edge of the tray. A 1.5 mm wide straight groove runs along the edge of the top cover to accommodate a stainless steel tube. This is used for the passage of a radioactive source for calibrating the modules. Each connector has two holes and these are fixed to the scintillator-plastic assembly through matching holes in them. Each ϕ sector in each ring

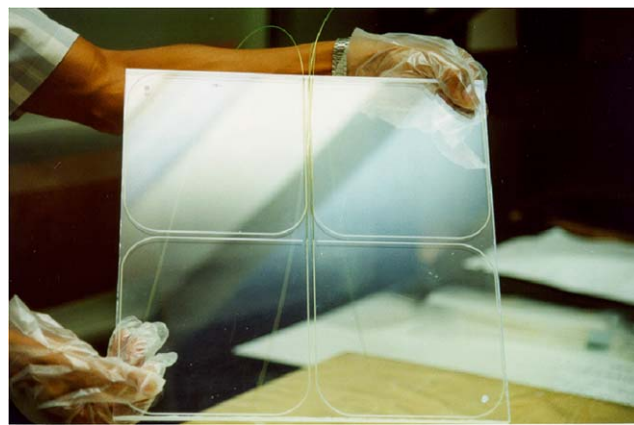


Fig. 7 View of a typical HO tile with WLS fibres inserted in the 4 σ -grooves

has 6 trays. There are 360 trays for layer 1 and 72 trays for layer 0.

3.3 Tiles

Scintillator tiles are made from Bicorn BC408 scintillator plates of thickness 10_{-1}^{+0} mm. Figure 7 shows a typical HO scintillator tile. The WLS fibres are held inside the tile in grooves with a key hole cross section. Each groove has a circular part (of diameter 1.35 mm) inside the scintillator and a neck of 0.86 mm width. The grooves are 2.05 mm deep. Each tile has 4 identical σ -shaped grooves, one groove in each quadrant of the tile. The grooves closely follow the quadrant boundary. The corners of the grooves are rounded to prevent damage to the fibre at the bend and to ease fibre insertion. The groove design is slightly different for the tile where the optical connector is placed at the end of the tray. Since the tiles are large, 4 grooves ensure good light collection and reduced attenuation of light.

3.4 Pigtails

The light collected by the WLS fibres inserted in the tiles is transported to photo-detectors. Captive ends of the WLS fibres, which reside inside the groove, are polished, aluminised and protected using a thin polymer coating. The other end of the WLS fibre comes out of the tile through a slot made on the 2 mm thick black plastic cover sheet. To minimise the loss of light in transportation, the WLS fibre (attenuation length of ~ 1.8 m) is spliced to a clear fibre (attenuation length of ~ 7.0 m). The clear fibres from each tile follow the guiding grooves on the top plastic to the optical connector at the end. Each tray has two optical connectors mounted on one end of the tray.

The pigtailed are tested with the help of a specially built scanner. A UV lamp moves over the pigtail in steps under a computer controlled mechanical stage. At specific intervals,

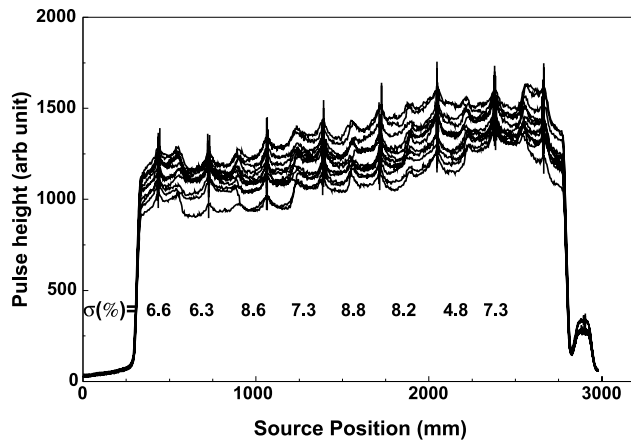


Fig. 8 Scan of a set of 12 fully assembled scintillator trays (of the 12 sectors) belonging to ring 0 layer 0 tray position 3 using the XY scanner. The lines on the figures refer to the peak positions corresponding to the centre of the tiles. The tiny peak on the right side is due to a calibration tile used in each set of measurements

the UV lamp shutter opens, the fibres absorb UV photons and the re-emitted light in each fibre is transmitted to the end to a set of photo diodes. Signals from the photo diodes are then digitised and analysed to study the combined transmission loss due to the spliced junctions, connector interface and attenuation. The mean transmission efficiency and its spread σ over the entire production was found to be about 95% and 3% respectively.

3.5 Quality assurance

Each tray is subjected to quality assurance tests using a moving radioactive source and cosmic ray muons. The radioactive source moves in X and Y direction (XY scanner) under computer control and the output is read by a data acquiring system. There is an additional set of measurements using radioactive source on a tip of a moving wire through the stainless steel tube fixed to each assembled tray.

XY scanner.

Figure 8 shows the ADC output at several x and y points over the surfaces of twelve trays belonging to layer 0 of ring 0. The collimated source excites all tiles in a similar way. One clearly sees that signal size for tiles close to the connector is largest and it slowly decreases for tiles with larger fibre length. The fluctuations among the trays are at a level much below 10% which is used as the acceptance criterion for the trays.

Wire source.

The tiles are also inter-calibrated using radioactive source technique [6] using a Co^{60} source with a strength of 5 mCi. Signals are sampled at the rate of 40 MHz and data are taken in histogramming mode. The mean of such a histogram represents the response of tile at the given reel position. While analysing the wire source data, the reel position is used to

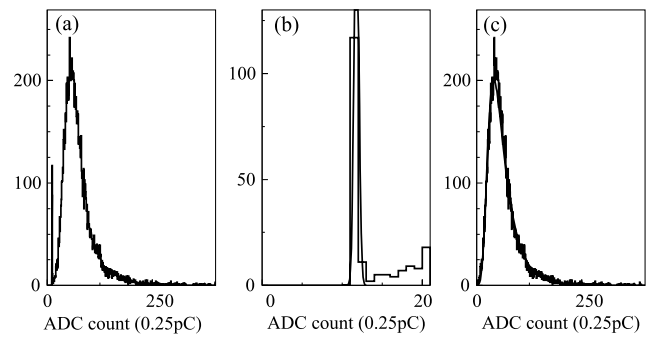


Fig. 9 (a) Raw ADC spectrum for tile # 8 for one of the trays of layer 0, ring 0. (b) A fit to the pedestal part to a Gaussian distribution. (c) Pedestal subtracted distribution fit to a Gaussian convoluted by Landau distribution

find out the correspondence between tile number and read-out channel. This method is also used during the beam test operation.

Cosmic ray muons.

The HO trays are expected to supply at least 10 photo electrons for a minimum ionising particle (MIP) if a photo-multiplier tubes (PMT) were used in the read out. The completed trays have been tested for their responses to MIPs using cosmic rays in a special stand designed for this purpose. Signal from the tiles are grouped into two bunches (alternate tiles) and the two bunches are read by two calibrated green extended PMT's. The electronic signal is sent to a charge-sensitive ADC. These PMT's are calibrated using a precisely controlled LED source to get the calibration for ADC count per photo electron.

Data are monitored on-line, and there is a calibration tile equipped with fibre readout which is placed separately and is always read out to monitor PMT gain. The pulse height spectrum is analysed offline to determine the number of primary photoelectrons produced. Figure 9(a) shows a typical ADC spectrum. One sees a very clear peak due to the passage of cosmic muons well separated from the pedestal peak. The pedestal subtracted pulse height (see Fig. 9(c)) is fit to a Landau distribution and the position of the mean value of the spectrum together with the calibration factor provides the number of photoelectrons yielded in a given tile.

Figure 10 shows a plot of average number of photoelectrons obtained using cosmic muons as a function of the tile number in a set of 12 trays belonging to ring 0 layer 0 tray. The number of photo electrons is always above 10. The number of photoelectrons are as high as 20 for some of the tiles close to the connector. The RMS spread in the number of photo-electrons is less than 10% which is the acceptance criterion for the trays.

Correlation between all the tests (with cosmic muons, source tube and XY scan) are shown in Fig. 11. The same set of readout system is used for cosmic and wire source test, whereas a different PMT, optical connectors are used

in the XY-scanner. Statistical error on these measured variables are less than 1%. HO towers will be calibrated in situ using muons from Cosmic Rays or from collision events. The strong correlation among the three set of measurements validates the use of muons in calibrating the system.

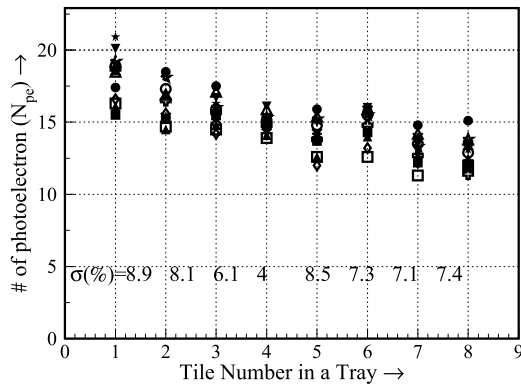


Fig. 10 Number of photo electrons for the tiles in fully assembled scintillator trays of layer 0, ring 0. Tile number is counted from the tray end with optical connectors. There are 12 measurements for a given tile number coming from 12 trays and the measurements have a spread between (4.0–8.9)% depending on the tile position

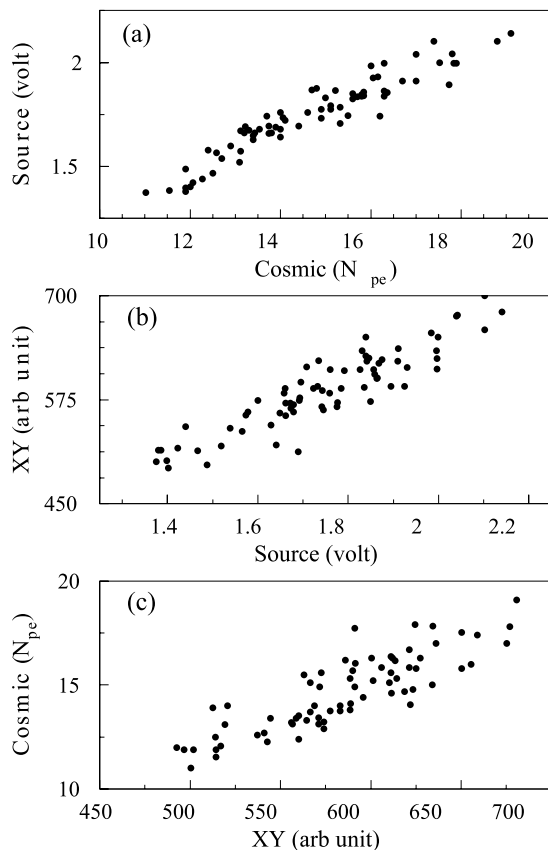


Fig. 11 Correlations between number of photo-electron in cosmic test and signal peaks in the wire source and XY source test

4 Test beam studies

Prototypes of HO trays were put together with 2 wedges of HB and a prototype of one sector of endcap hadron calorimeter module (HE) in the test beam facility at CERN. These prototype trays were exposed to a variety of charged beams of hadrons, electrons and muons over a wide energy range. These tests have been carried out over a number of years.

The signals from these detectors are read out using Hybrid Photo Diodes (HPD) which converts the light into electrical charges. These are then measured and encoded into a non-linear digital scale by the charge integrator IC (QIE). The QIE uses LHC clock to divide time into regular bins of 25 ns and measures the accumulated charge in each time bin. Internally, the QIE uses capacitors to accumulate the charge. The QIE's use non linear scale so that the QIE step size is a approximate constant fraction of the QIE value.

Arrival time of signals from a scintillator tile to the read-out element has a significant dependence on its location in η . Energy is measured by summing over a number of time steps and this η dependent time of arrival makes the energy measurement rather complex. This effect has been corrected by introducing an appropriate η dependent delay for each of the towers (see Fig. 12).

The HO trays were calibrated using the radioactive wire-source test as well as by exposing them to a beam of muons. Figure 13 shows a plot of ADC counts from ring 0, layer 0 tiles as the radioactive source passes along them. Pedestal is determined by fitting a straight line with zero slope to the part of the signal when radioactive source is far away from the tile under consideration. It is evident from the figure that there is a leakage of the signal to the neighbouring tiles when the source is at the edge of the tile. So the pedestal subtracted signals from the neighbouring tiles are added with proper weight to get the total signal. The signal from a tile and that

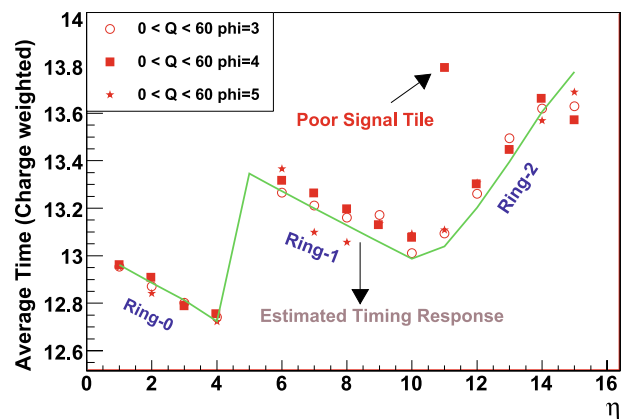


Fig. 12 Charge weighted average time slice as a function of η index of the tiles in HO before the time delay correction. The solid line is the estimated time delay as a function of rapidity. The vertical axis is in units of 25 ns

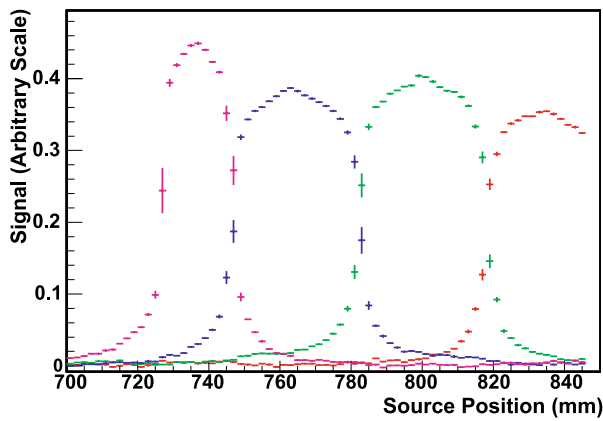


Fig. 13 ADC counts from ring 0, layer 0 tiles as the radioactive source passes along them

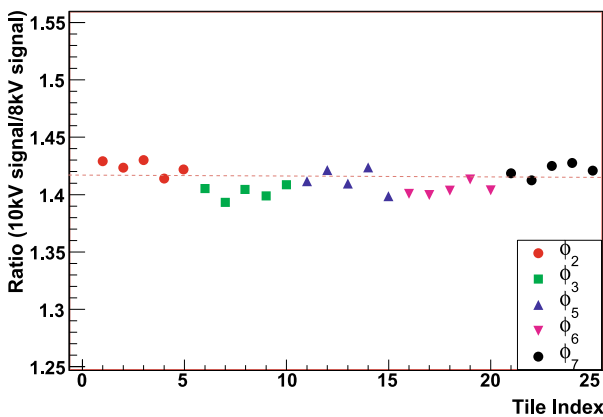


Fig. 14 Ratio of calibration constant for ring 1 HO tiles in the test beam set up with the HPD's being operated at 10 kV and at 8 kV

from its two adjacent ones are then fitted using an iterative procedure to extract the calibration constants. The HPD for HO readout was operated at voltages 8 kV and 10 kV and separate datasets were collected for the two settings. Calibration constants obtained with these two sets of data show the expected consistency (see Fig. 14) being different by a factor of 1.415.

For the muon runs, the HPD was operating at 10 kV. Muons of energy 150 GeV were pointed at the centre of each HO tile and the ADC signal from the tile was used to determine the muon calibration constant. To minimise noise, the sum of 4 time slices with the maximum amount of energy deposit is taken as the signal. The sum of four time slices away from the signal region is taken as the pedestal. The pedestal and signal peaks are not well separated in all the tiles (particularly in ring 1). The pedestal value is taken as the fitted mean obtained by fitting the pedestal distribution with a Gaussian curve. The peak region of the pedestal subtracted muon signal is then fitted with a Gaussian convoluted Landau distribution and the most probable value of this distribution from the fit is taken as the muon calibra-

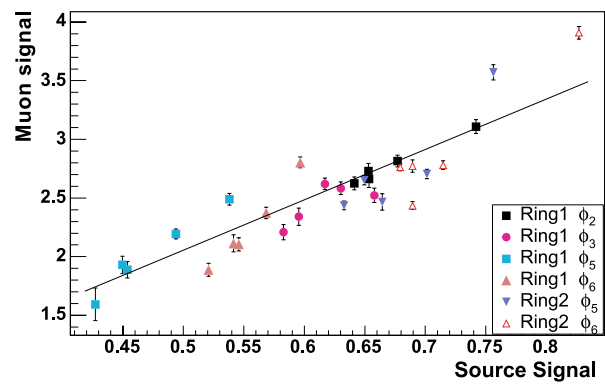


Fig. 15 Calibration constant from muon signal plotted against the corresponding value from the wire source analysis

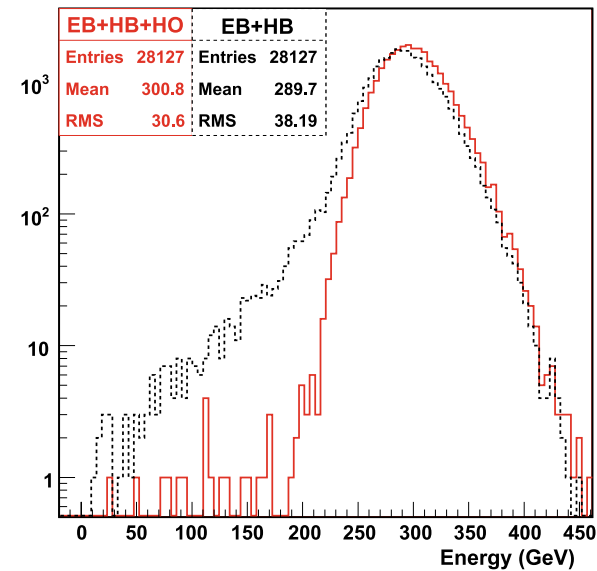


Fig. 16 Energy distribution for a 300 GeV pion beam measured with EB + HB and with EB + HB + HO

tion constant for the tile. Figure 15 shows correlation plot of calibration constants obtained through the two sets of measurements.

Figure 16 shows energy distributions for a 300 GeV pion beam with only EB + HB in the beam and with EB + HB + HO in the beam. The measured distributions for hadron beams of energy above 100 GeV are more symmetric and have smaller widths after adding HO to the energy calculation. This confirms the expectations from simulation studies shown in Sect. 2.

Weight factor for HO has been found to be 372 MeV/fC for beam being shot at $\eta = 0.22$ (Ring 0) and 407 MeV/fC at $\eta = 0.56$ (Ring 1). Energy resolution for π^- beam has been measured as a function of beam energy at those two configurations and are shown in Fig. 17. The two sets of measurements in each plot refer to measurements without and with HO. As can be seen in the figure, the energy reso-

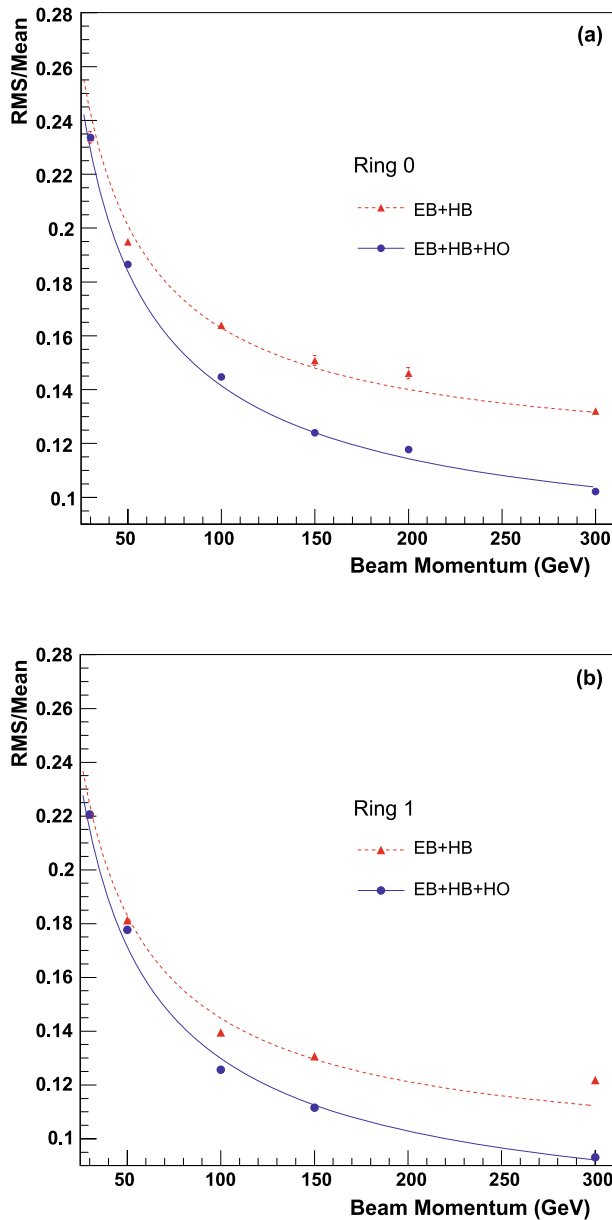


Fig. 17 Energy resolution for pions as a function of beam energy measured with EB + HB and with EB + B + HO for the beam being shot at (a) $\eta = 0.22$ and (b) $\eta = 0.56$

lution improves significantly with the inclusion of HO. Fits to the resolution distribution indicate that the constant term improves from 11.2% to 7.8% for Ring 0 and from 9.2% to 6.6% for Ring 1 with the inclusion of HO.

The signal size due to penetrating beam is compared with the noise level in the HO tiles from the 2002 test beam studies (see Fig. 18). This study has indicated that HO will be able to provide signals for minimum ionising particles with efficiency better than 90% keeping the noise level below 20%. CMS uses signal in the Resistive Plate Chambers (RPC) to trigger for muons in the barrel as well as in the end-cap region. In the barrel (RPC trigger towers 1–6) and in the

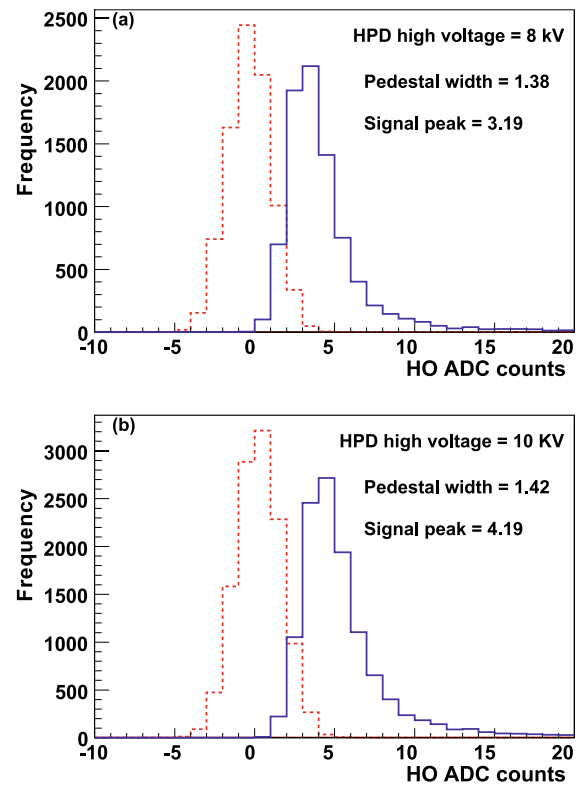


Fig. 18 Pedestal peak and muon signal for a ring 2 tile operated with a voltage of (a) 8 kV, (b) 10 kV on the HO HPD

overlap region (RPC trigger towers 7–9) for the RPC's, there is a considerable coverage by HO. Even with 95% chamber efficiency of the RPC's, muon trigger efficiency is rather poor for RPC trigger towers 6–9 (as low as 72%) if only RPC's are used in coincidence. This study suggested that HO could be a useful component in muon trigger together with the RPC at these solid angle [7].

As it has been observed for data collected with the radioactive source, size of the signal collected from HO tiles is considerably bigger when the HPD is operated at 10 kV as compared to 8 kV. This is also demonstrated in the HO signals obtained from the 2002 test beam studies with 225 GeV/c muon beam, as shown in Fig. 18. Use of HO will make the trigger efficiency better than 90% over most of the solid angle. Figure 19 shows a plot of efficiency of muon detection versus purity of the signal obtained from the 2004 test beam studies.

5 Summary

Constrained to lie outside of the inner CMS detector layers (pixels, tracker and electromagnetic calorimeter) and inside the 4 Tesla magnet, the hadron barrel calorimeter was found to be too thin to effectively absorb high energy hadrons specially near the $\eta = 0$ region. This would lead to degraded

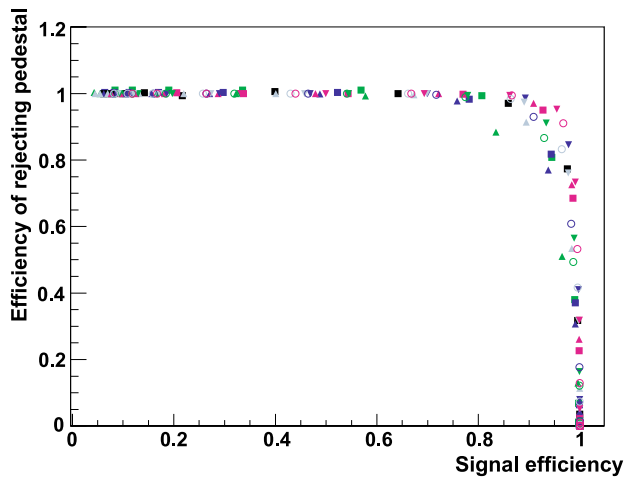


Fig. 19 Efficiency of rejecting pedestal vs. efficiency of muon detection in the HO detector. Different markers refer to different read out towers in ring 0

sensitivity for establishing new physics signals whose main signature is missing E_T . The addition of sensitive detector layers just outside the magnet (inside the first barrel muon station) was simulated and shown to ensure much smaller energy leakage and hence improvement in the energy resolution.

Sensitive detector layers of 1 cm thick plastic scintillator with embedded WLS fibre readout (similar technology

as the CMS barrel and endcap hadron calorimeters) were selected as the appropriate detectors.

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